

Physics potential of INO-ICAL

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(For the INO-ICAL Collaboration)

Workshop on neutrino programs with facilities in Japan, Aug 4-6th, 2015

Some good news started this year...

TATA INSTITUTE OF FUNDAMENTAL RESEARCH

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Press Release

The Union Cabinet of the Govt. of India chaired by the Prime Minister, Shri Narendra Modi, has given its approval for the establishment of India-based Neutrino Observatory (INO) at an estimated cost of Rs. 1500 crores.

The INO project is jointly supported by the Department of Atomic Energy and the Department of Science and Technology. Infrastructural support is provided by the Government of Tamil Nadu where the project is located. Tata Institute of Fundamental Research (TIFR), Mumbai is the host institute for INO.

But there is a long way to go.....

The INO-ICAL White Paper

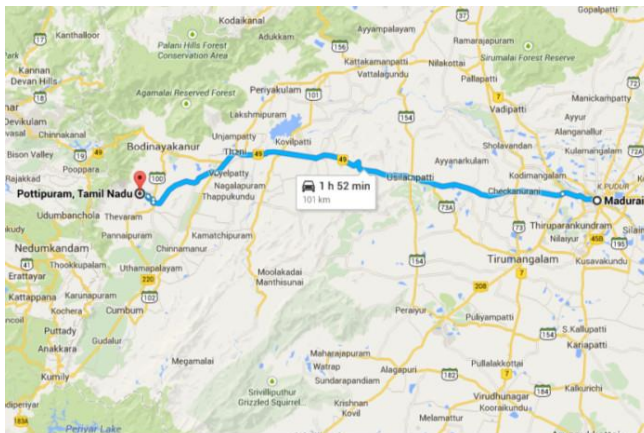
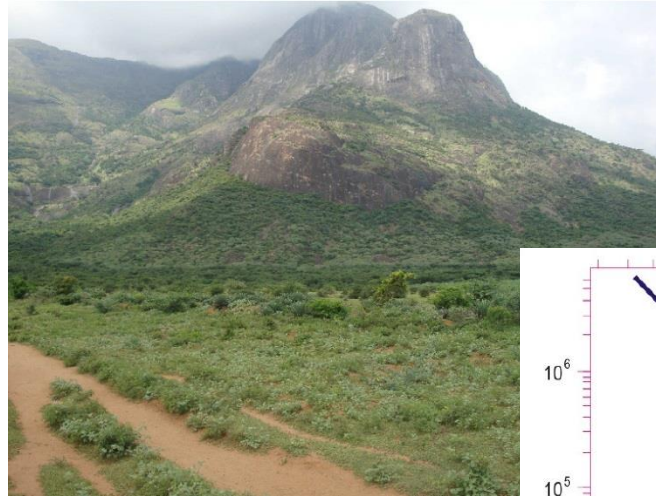
INO/ICAL/PHY/NOTE/2015-01

Physics Potential of the ICAL detector at the India-based Neutrino Observatory (INO)

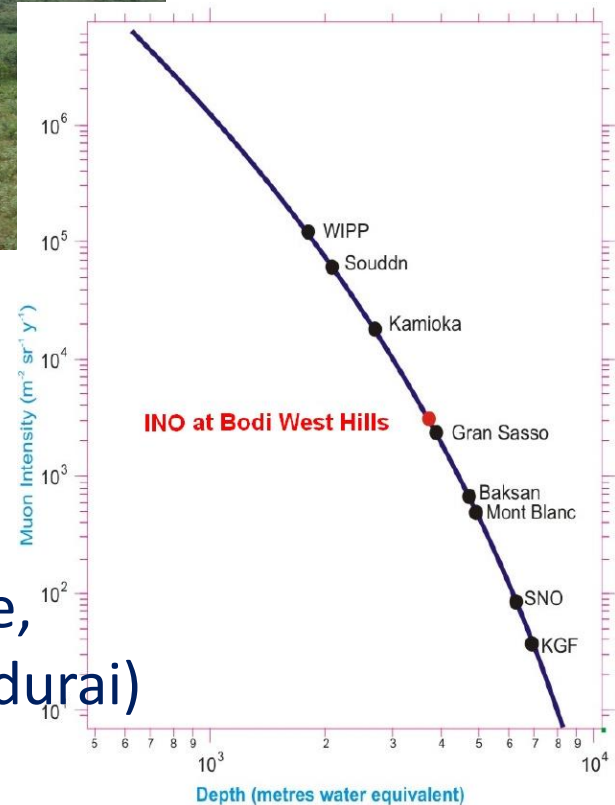
The ICAL Collaboration

arXiv:1505.07380v1 [physics.ins-det] 27 May 2015

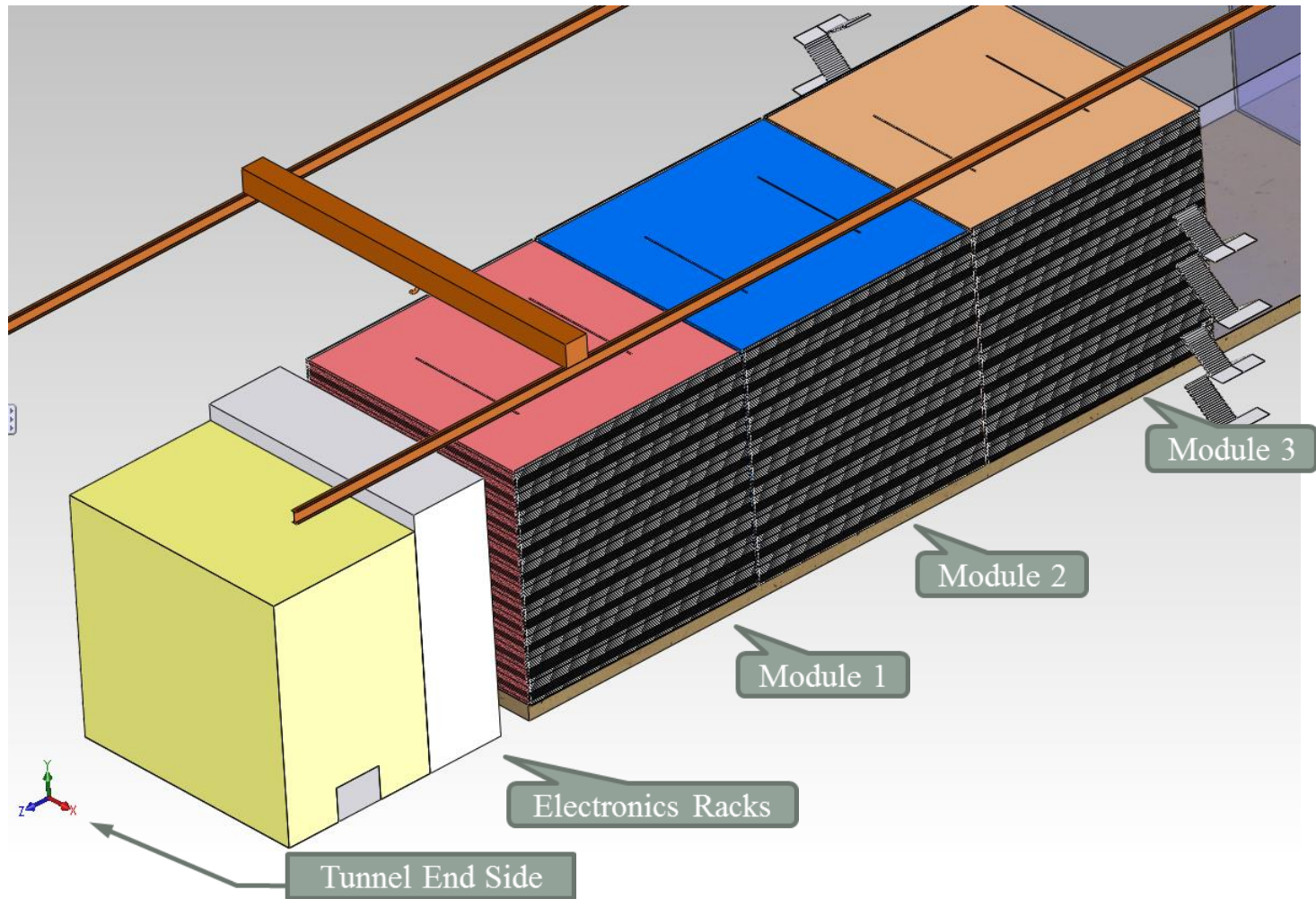
The location of INO



Bodi West Hills,
Pottipuram Village,
(100 km from Madurai)
Tamil Nadu State



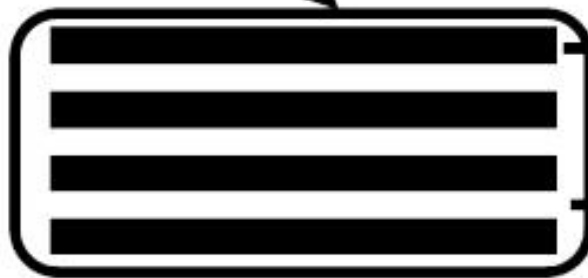
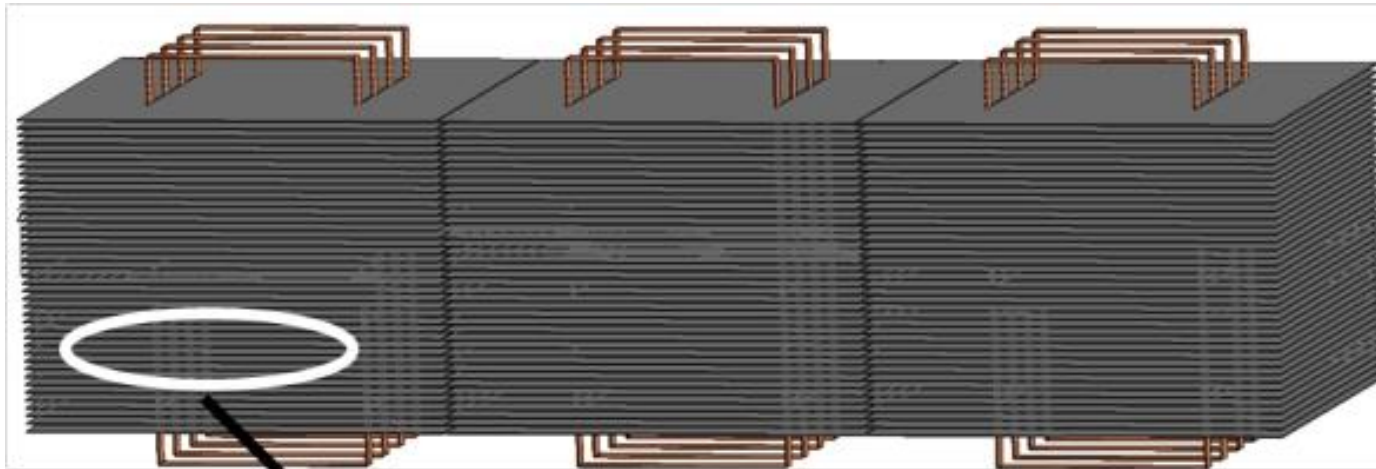
ICAL Modules in the main cavern



The ICAL detector: desiderata

- *Large target mass (50 – 100 kt)*
- *Good tracking and Energy resolution*
(Tracking calorimeter)
- *Good directionality for up/down discrimination*
(Nano-second time resolution)
- *Charge identification capability*
(uniform, homogeneous magnetic field)
- *Ease of construction & Modularity*
- *Complementarity to the other existing / proposed detectors*

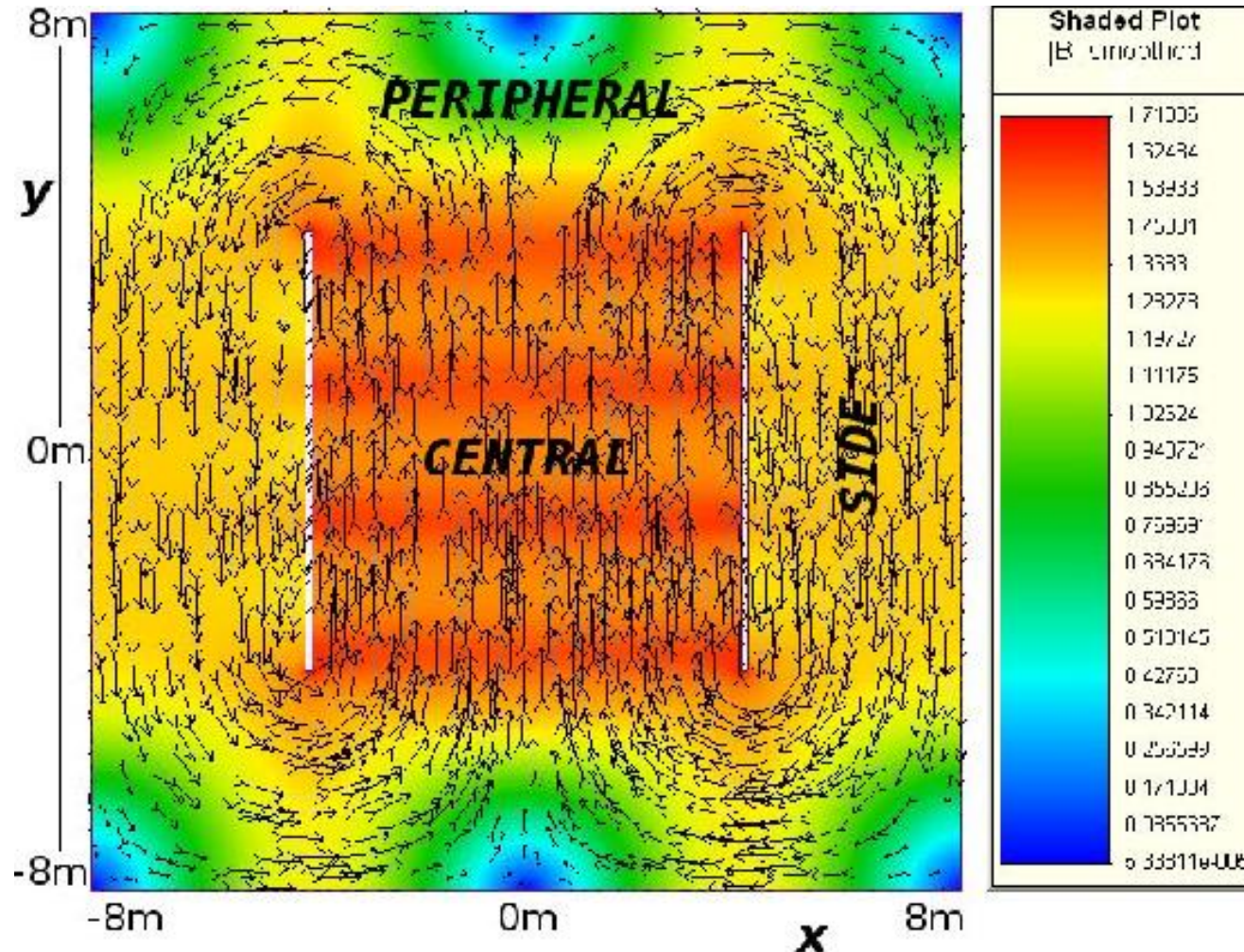
The ICAL detector



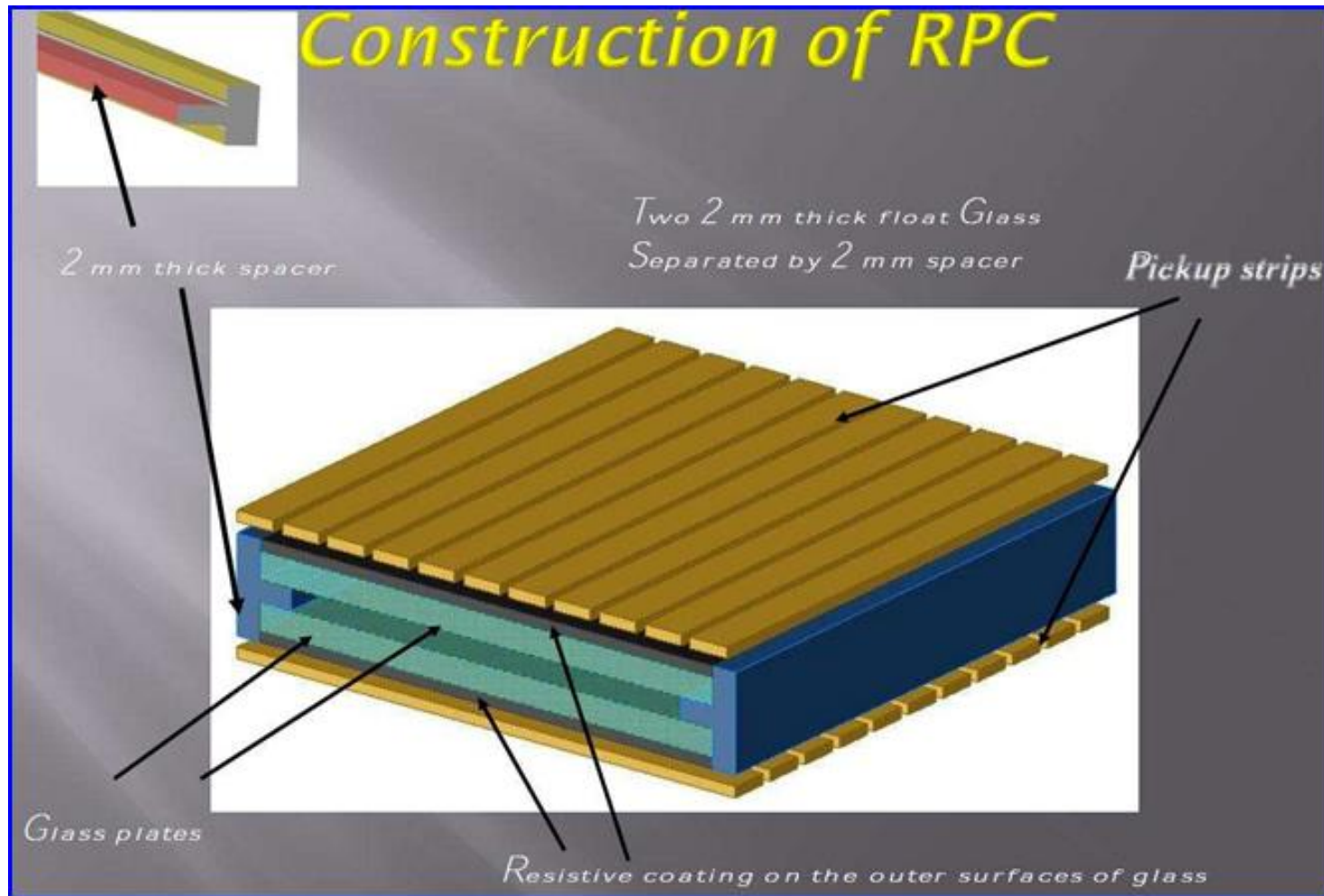
5.6 cm thick iron plate

4 cm air gap for RPC detector

The magnetic field



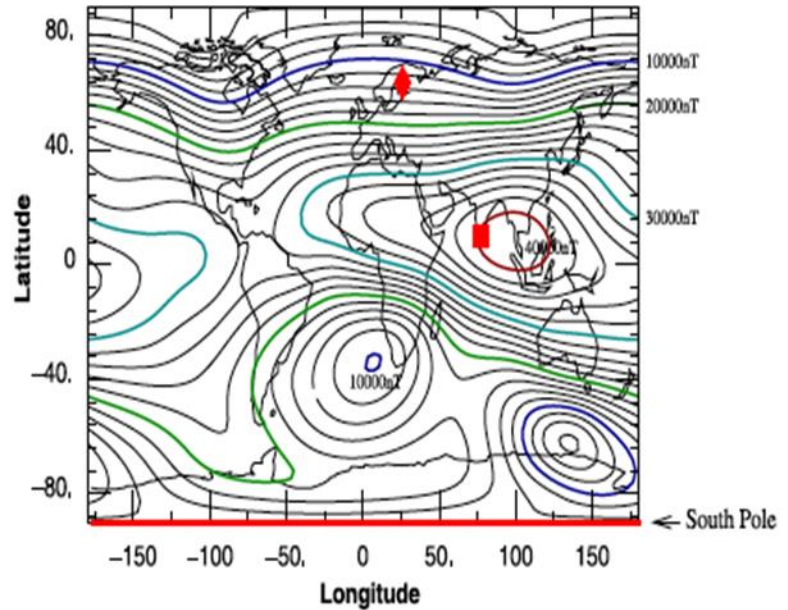
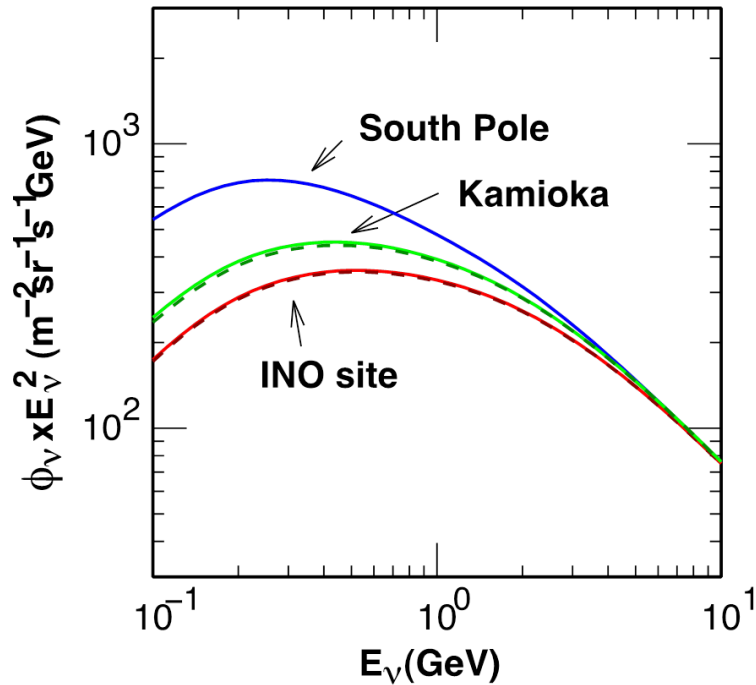
The active detector: RPC



The detector specifications

<i>No of modules</i>	<i>3</i>
<i>Module dimension</i>	<i>16 m X 16 m X 14.4m</i>
<i>Detector dimension</i>	<i>48.4 m X 16 m X 14.4m</i>
<i>No of layers</i>	<i>150</i>
<i>Iron plate thickness</i>	<i>5.6cm</i>
<i>Gap for RPC trays</i>	<i>4 cm</i>
<i>Magnetic field</i>	<i>1.4 Tesla</i>
<i>RPC unit dimension</i>	<i>195 cm x 184 cm x 2.4 cm</i>
<i>Readout strip width</i>	<i>3 cm</i>
<i>No. of RPCs/Road/Layer</i>	<i>8</i>
<i>No. of Roads/Layer/Module</i>	<i>8</i>
<i>No. of RPC units/Layer</i>	<i>192</i>
<i>Total no of RPC units</i>	<i>28800</i>
<i>No of Electronic channels</i>	<i>3.7 X 10⁶</i>

Atmospheric neutrino flux

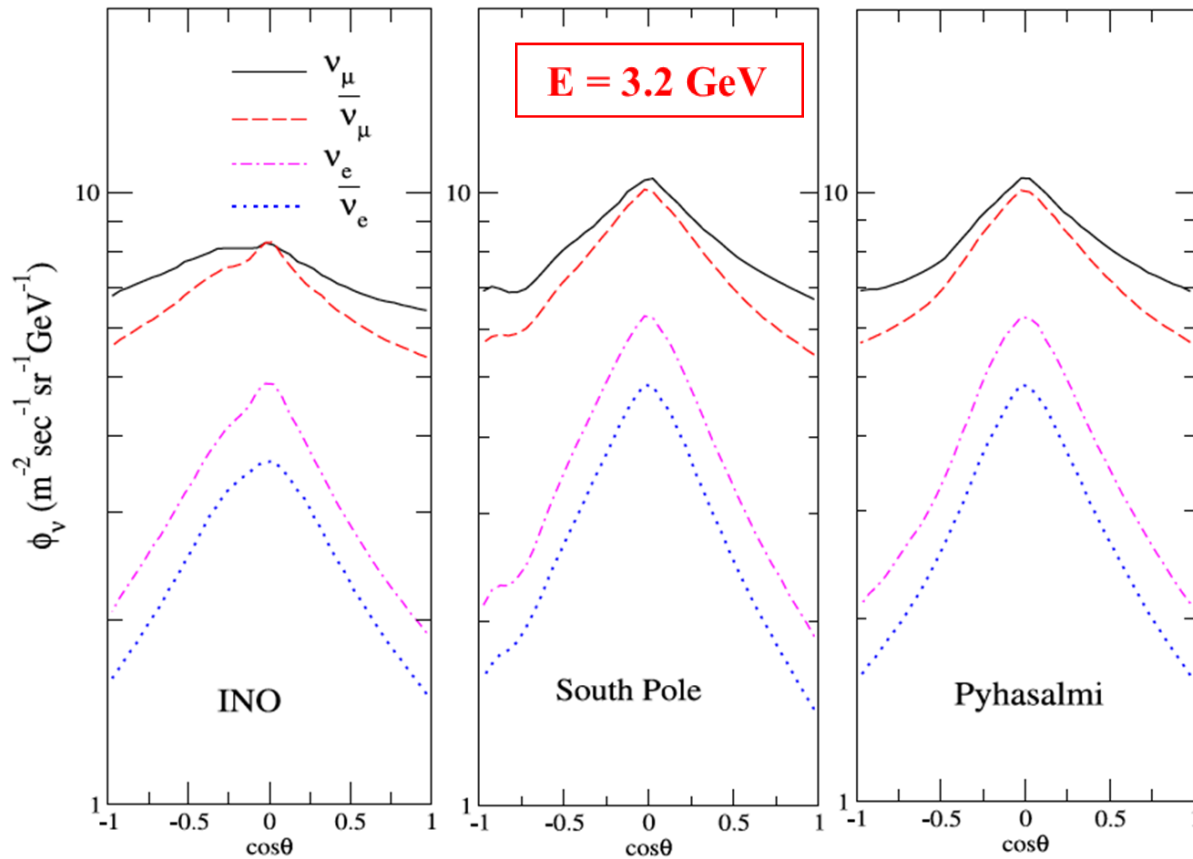


**Averaged over all directions
Summed over all flavors of neutrino and anti-neutrino**

**Horizontal component of the
geomagnetic field**

**Magnitude at the Earth's
surface ranges from 25 to 65
microtesla**

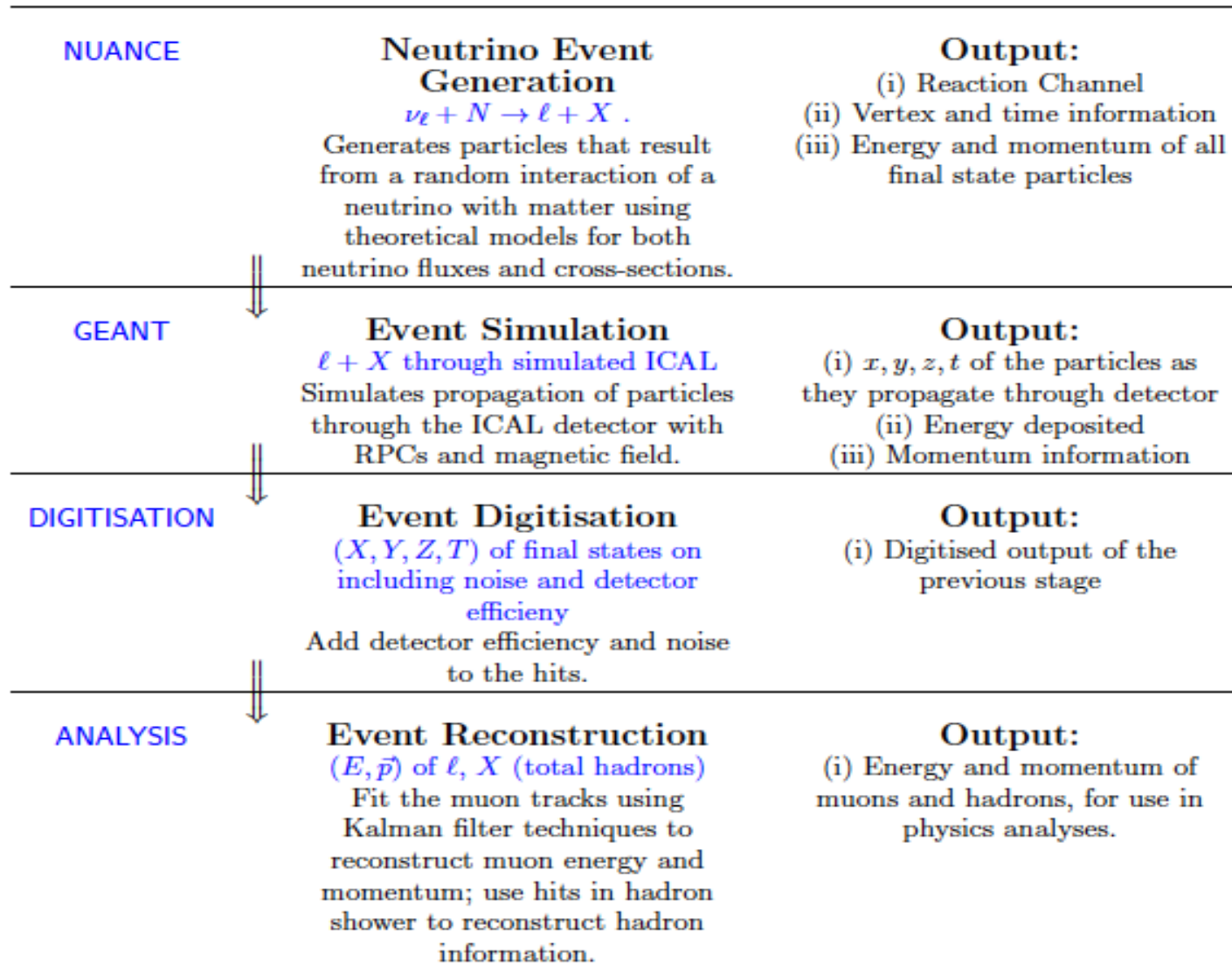
Angular distribution of neutrino flux



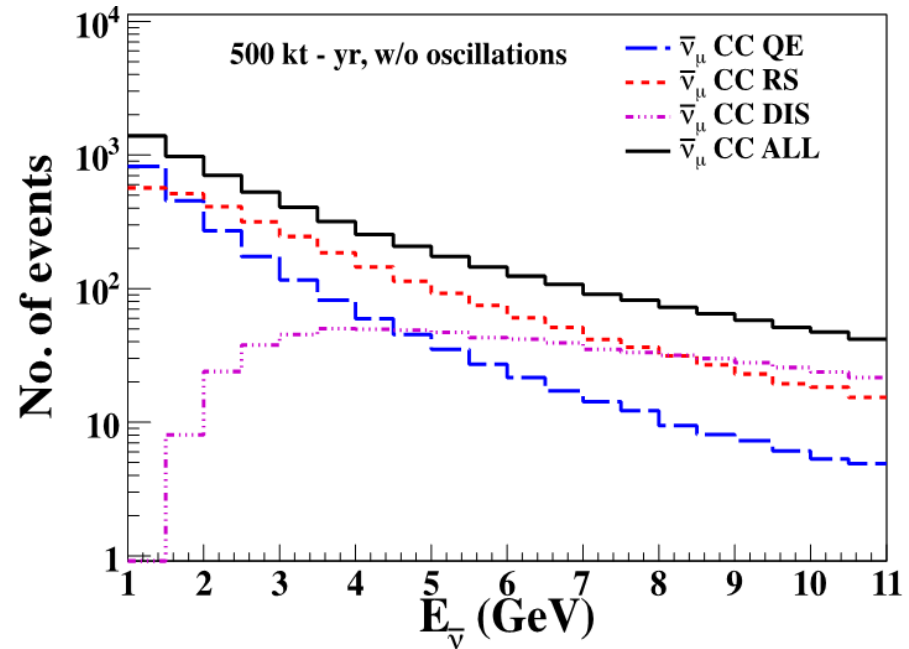
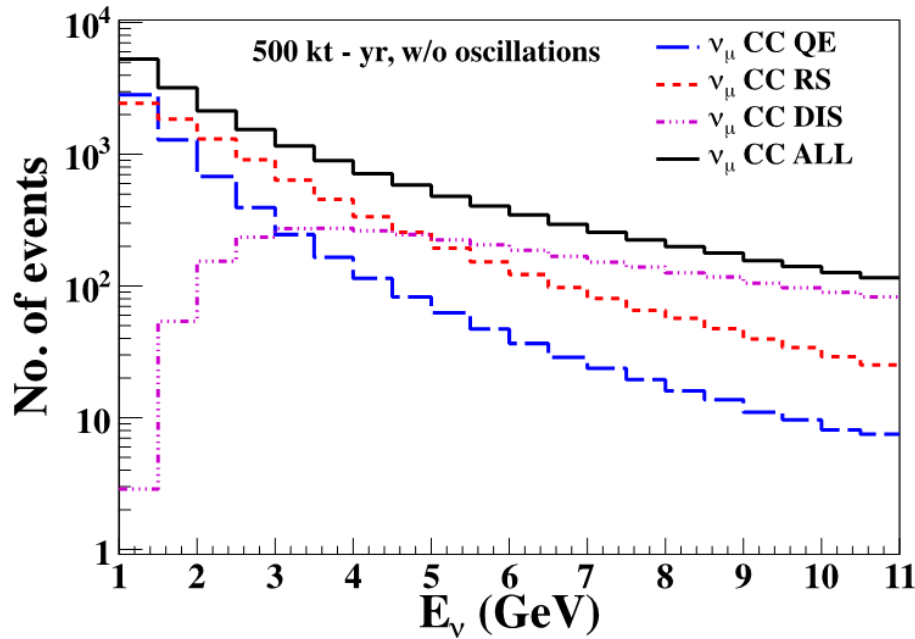
Athar, Honda, Kajita, Kasahara, Midorikawa, arXiv:1210.5154 [hep-ph]

Physics potential of INO-ICAL

Simulation framework

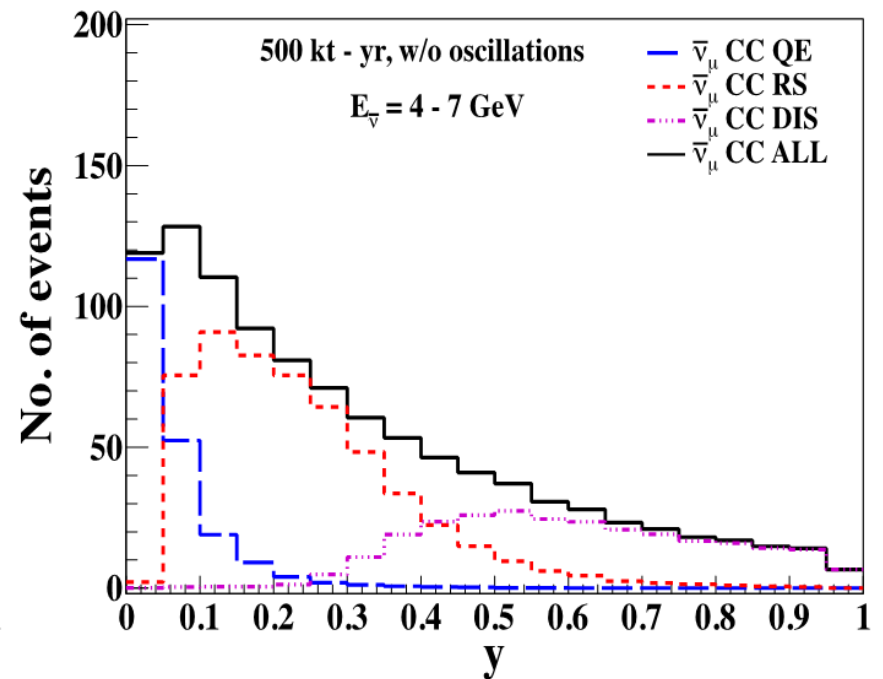
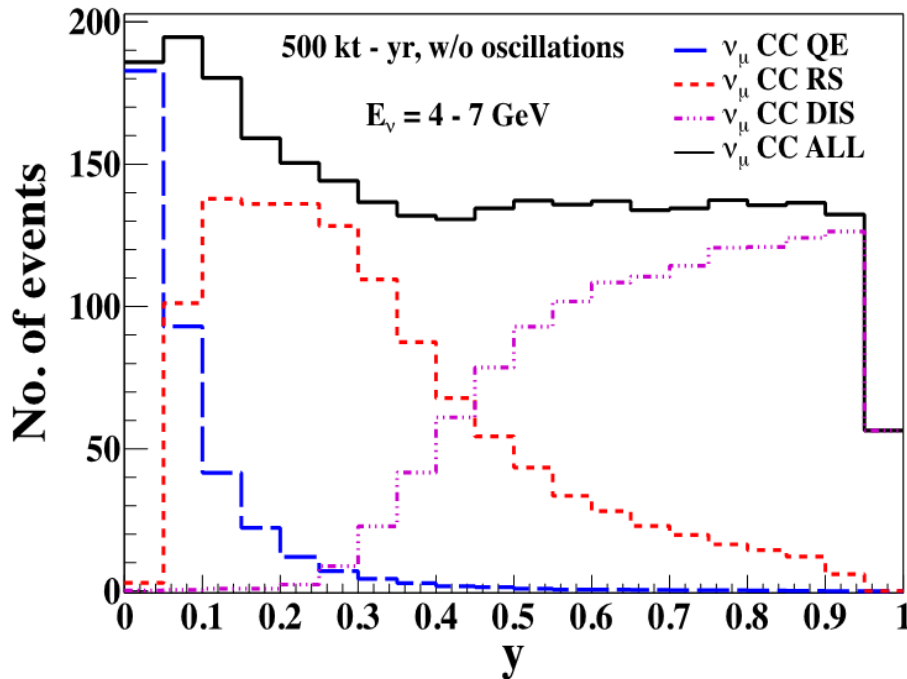


Event distribution (NUANCE)



Relative contributions of three cross-section processes to the total events in the absence of oscillation and without detector efficiency and resolutions

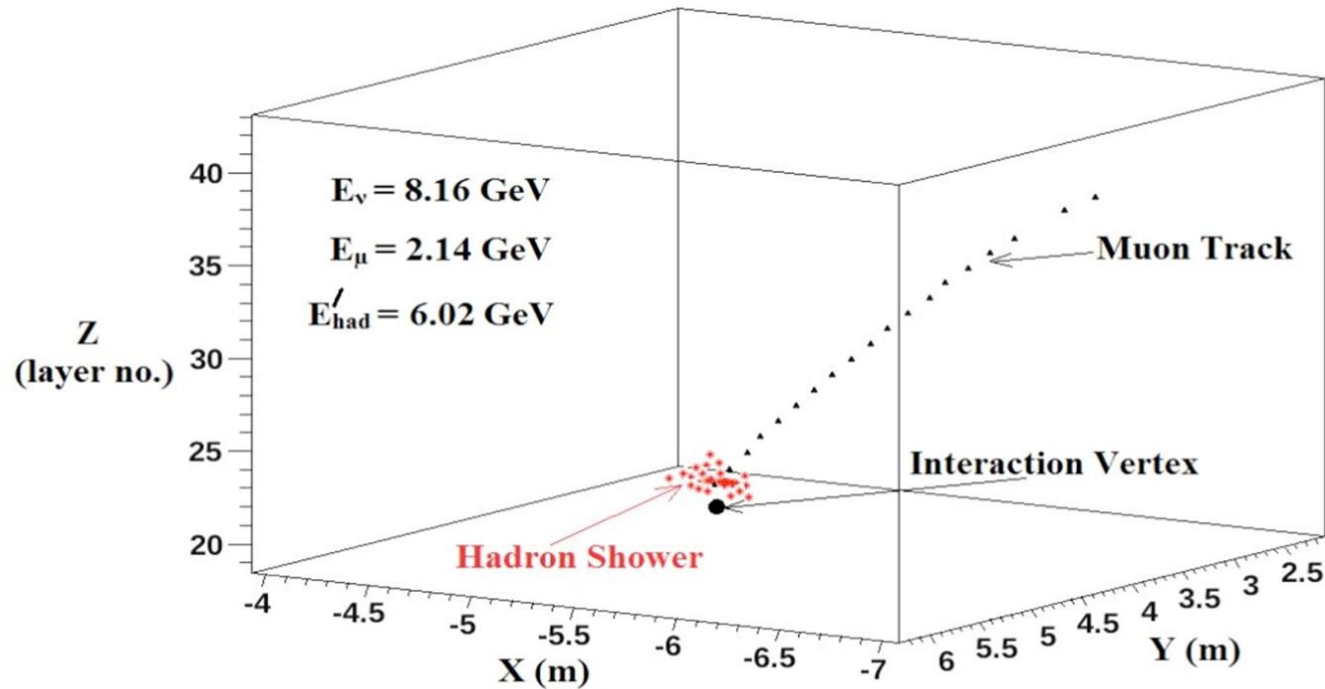
Inelasticity distribution



Inelasticities in individual events have a wide distribution

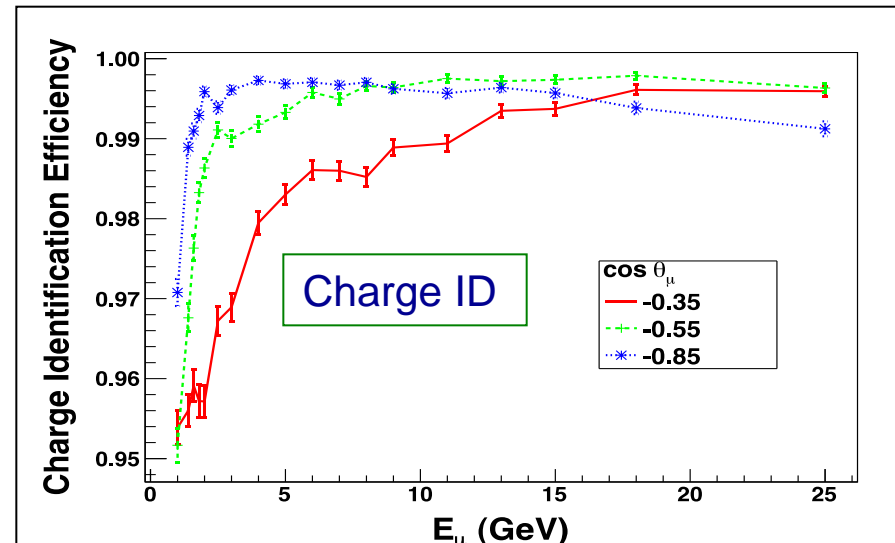
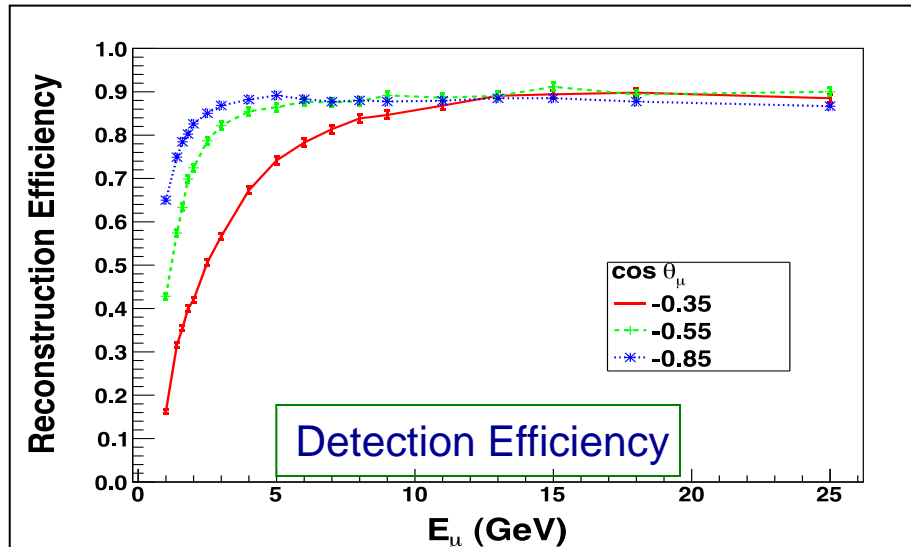
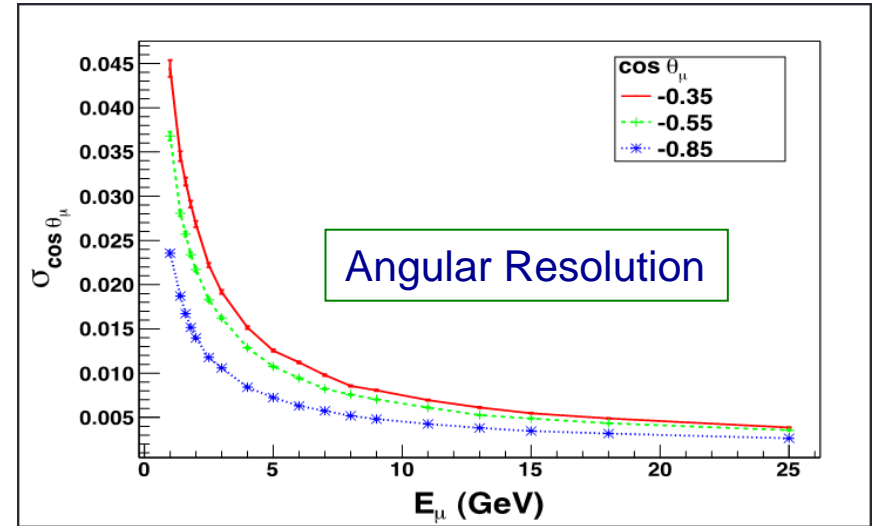
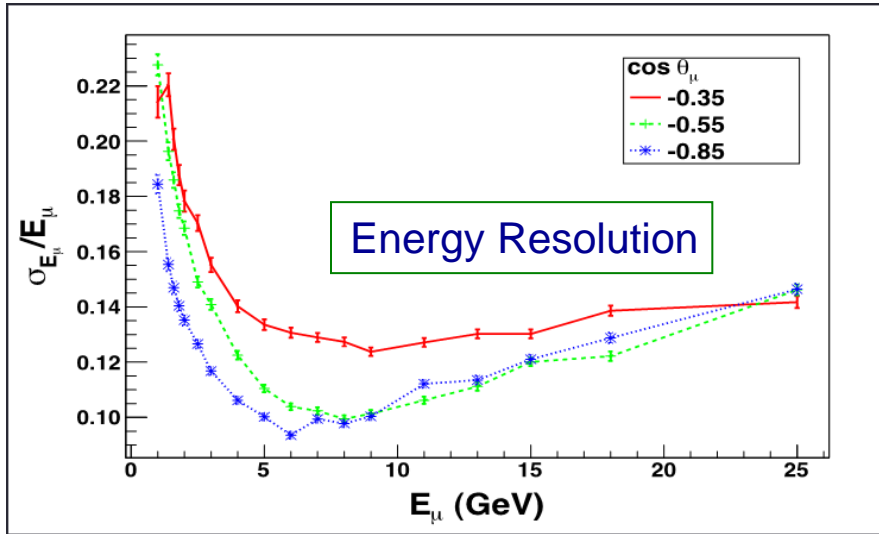
Important to measure inelasticity in individual events

A typical CC event in ICAL

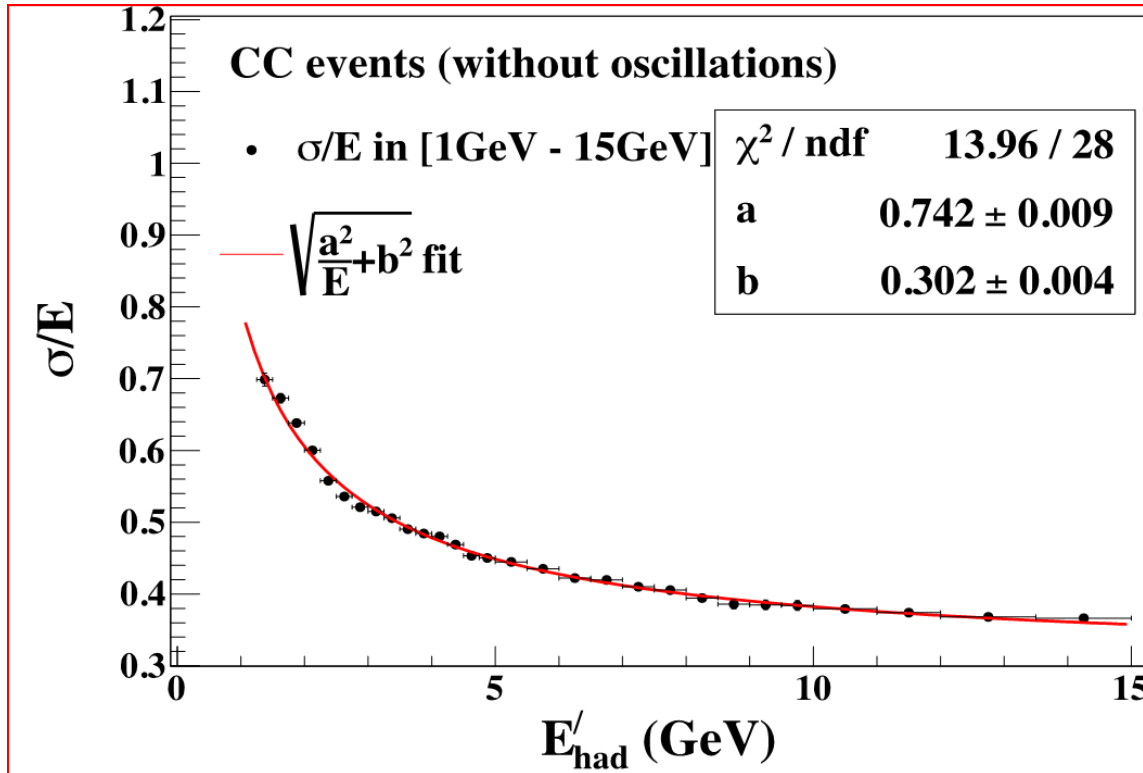


Using GEANT4 simulation

Detector response to muons



Detector response to hadrons



$$E_h = E_\nu - E_\mu \text{ (from hadron hit calibration)}$$

Hadron energy resolution: 85% at 1 GeV and 36% at 15 GeV

Typical analysis technique

We define the Poissonian χ^2 for μ^- events as :

$$\chi^2 = \min_{\xi_l} \sum_{i=1}^{N_{E'_{\text{had}}}} \sum_{j=1}^{N_{E_\mu}} \sum_{k=1}^{N_{\cos \theta_\mu}} \left[2(N_{ijk}^{\text{theory}} - N_{ijk}^{\text{data}}) - 2N_{ijk}^{\text{data}} \ln \left(\frac{N_{ijk}^{\text{theory}}}{N_{ijk}^{\text{data}}} \right) \right] + \sum_{l=1}^5 \xi_l^2,$$

where

$$N_{ijk}^{\text{theory}} = N_{ijk}^0 \left(1 + \sum_{l=1}^5 \pi_{ijk}^l \xi_l \right).$$

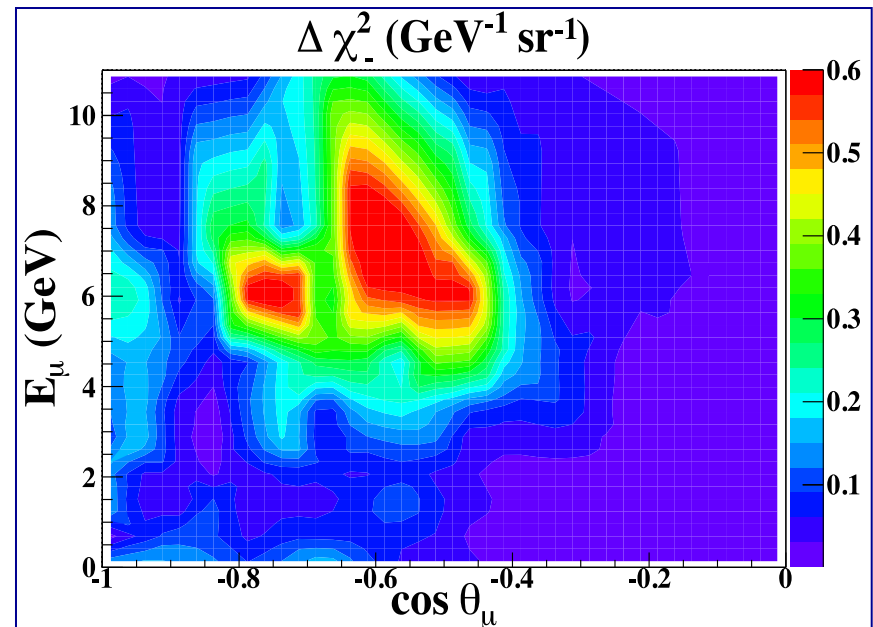
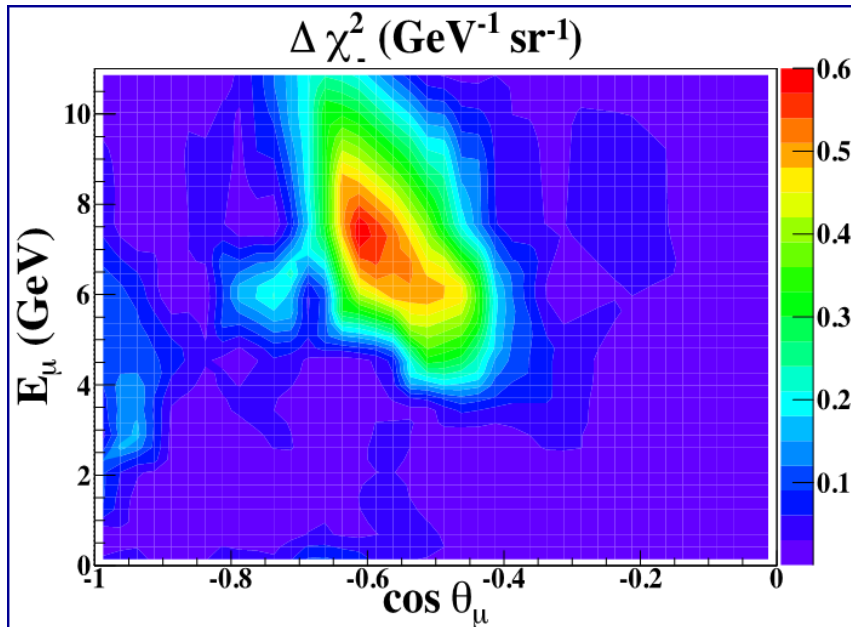
Observable	Range	Bin width	Total bins
E_μ (GeV)	[1, 4)	0.5	6
	[4, 7)	1	3
	[7, 11)	4	1
$\cos \theta_\mu$	[-1.0, -0.4)	0.05	12
	[-0.4, 0.0)	0.1	4
	[0.0, 1.0]	0.2	5
E'_{had} (GeV)	[0, 2)	1	2
	[2, 4)	2	1
	[4, 15)	11	1

- 1) Overall 5% systematic uncertainty
- 2) Overall flux normalization: 20%
- 3) Overall cross-section normalization: 10%

- 4) 5% uncertainty on the zenith angle dependence of the fluxes
- 5) Energy dependent tilt factor:
 $\Phi_\delta(E) = \Phi_0(E) [E/E_0]^\delta \approx \Phi_0(E) [1 + \delta \ln E/E_0]$
 where $E_0 = 2$ GeV and
 δ is the 1σ systematic error of 5%

Importance of hadron information

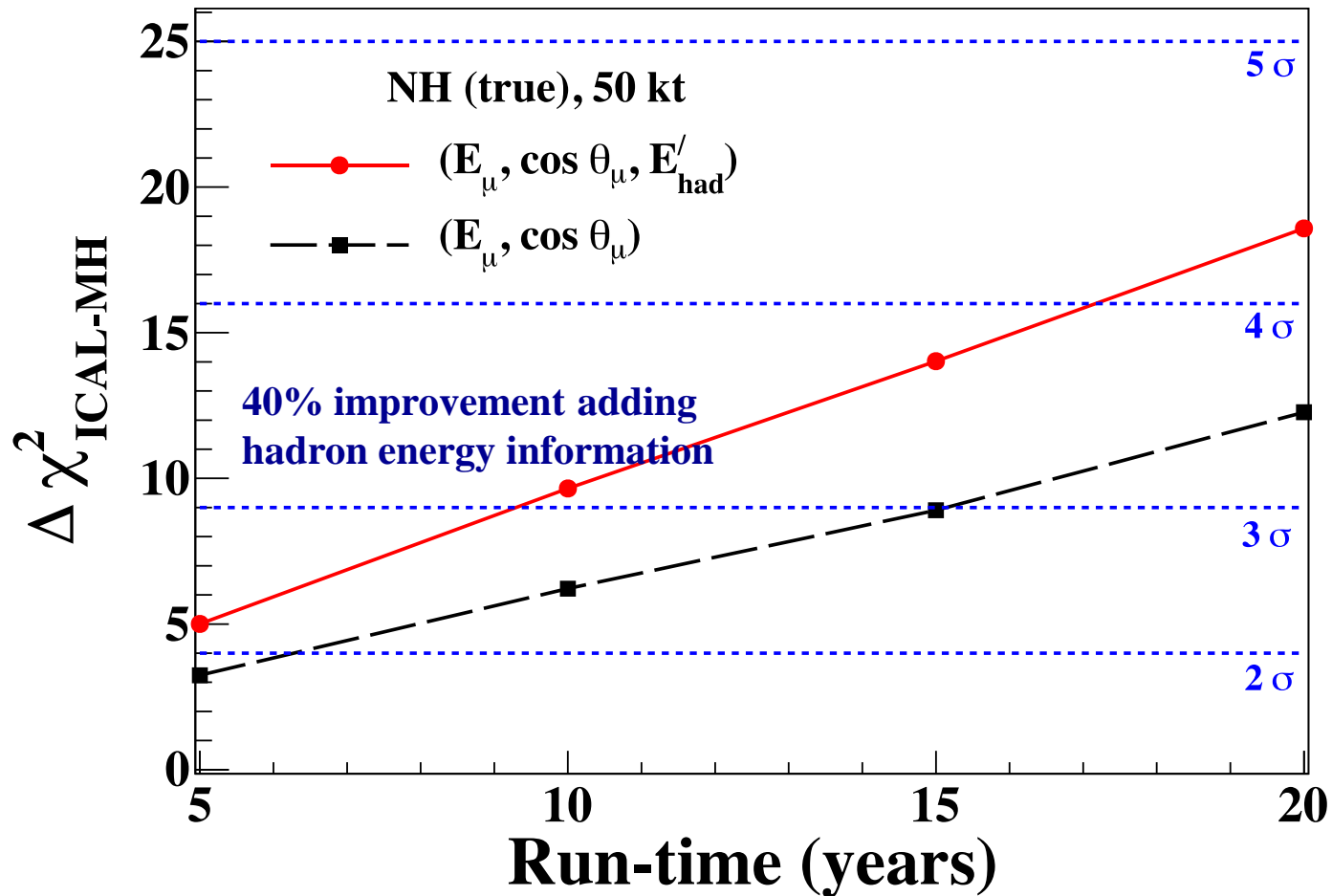
Distribution of $(\Delta\chi^2/\text{area}) [\chi^2 (\text{IH}) - \chi^2 (\text{NH})]$ for mass hierarchy discrimination considering μ^- events



Hadron energy information not used

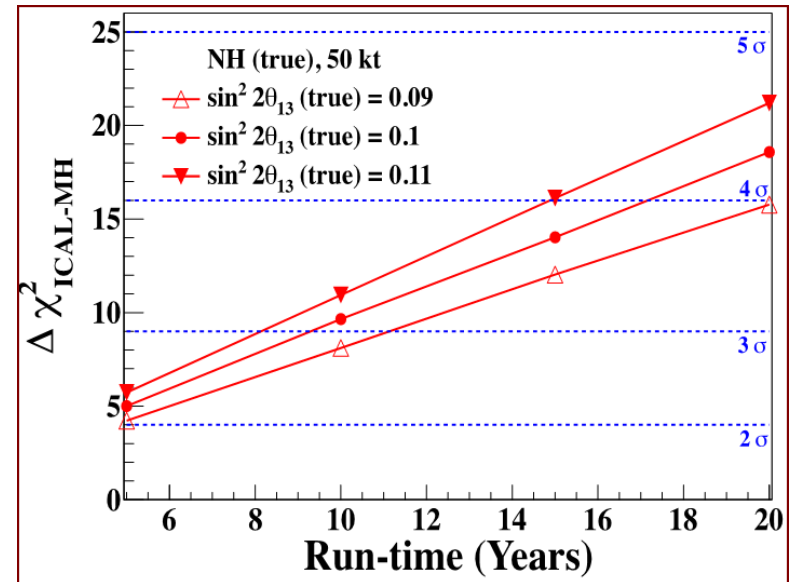
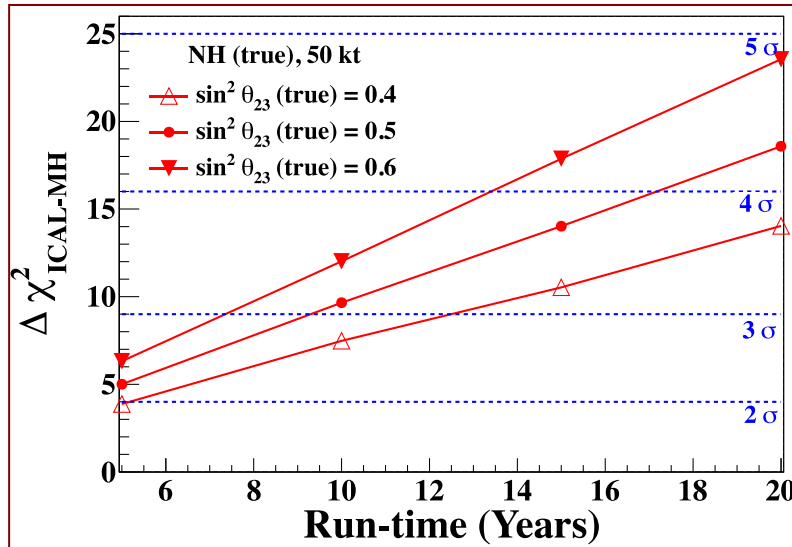
Hadron energy information used

Mass hierarchy sensitivity: ICAL alone



Devi, Thakore, Agarwalla, Dighe, arXiv:1406.3689 [hep-ph]

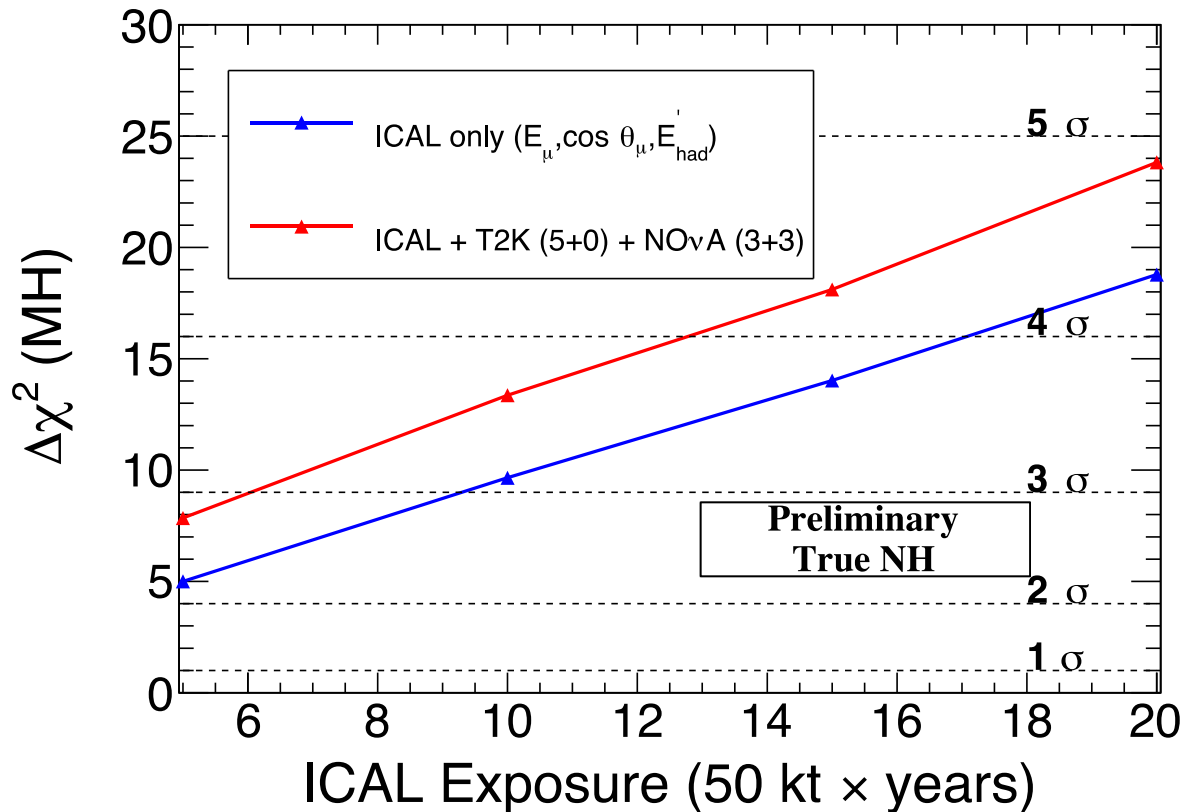
MH sensitivity dependence on mixing angles



Devi, Thakore, Agarwalla, Dighe, arXiv:1406.3689 [hep-ph]

50 kt ICAL can rule out the wrong hierarchy with median $\Delta\chi^2 \approx 7$ to 12 depending on the true values of θ_{23} and θ_{13} in 10 years

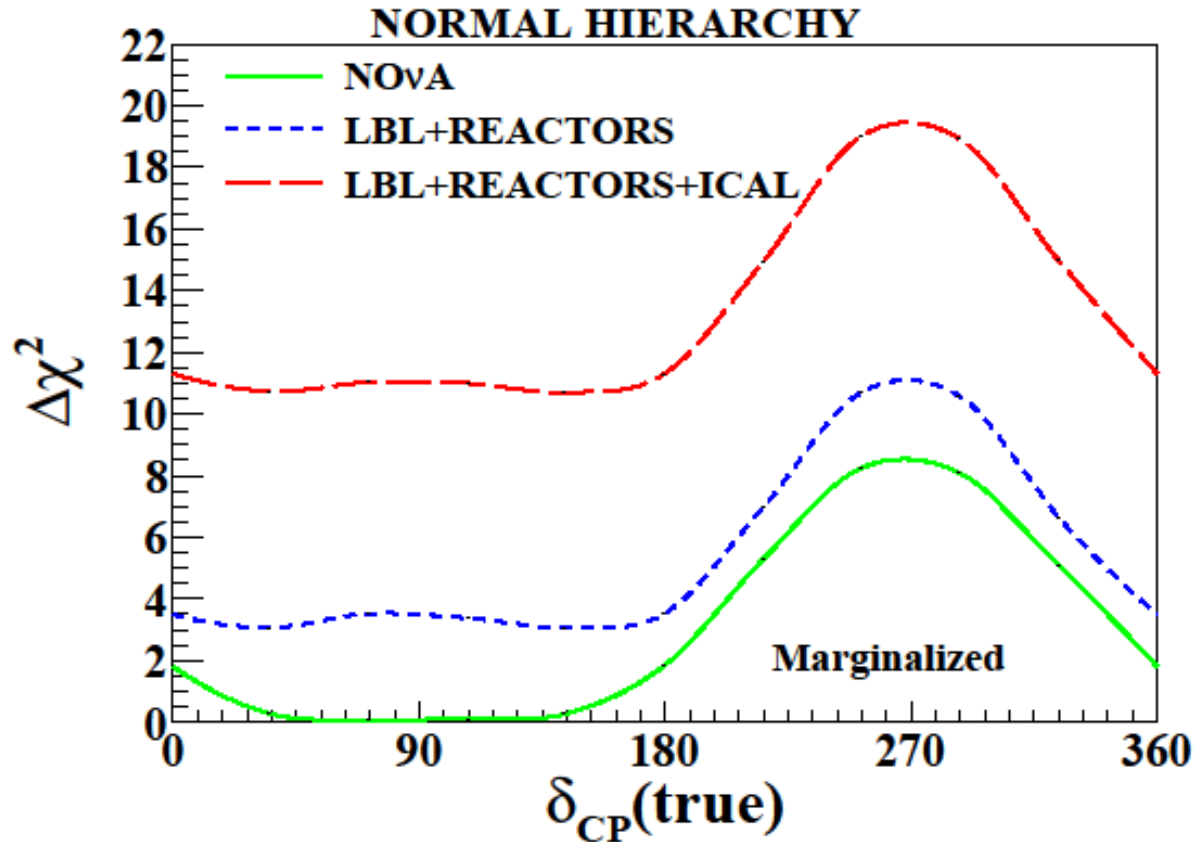
Synergy with T2K and NOvA (at $\Delta\text{CP}=0$)



Thakore, Agarwalla, work in progress

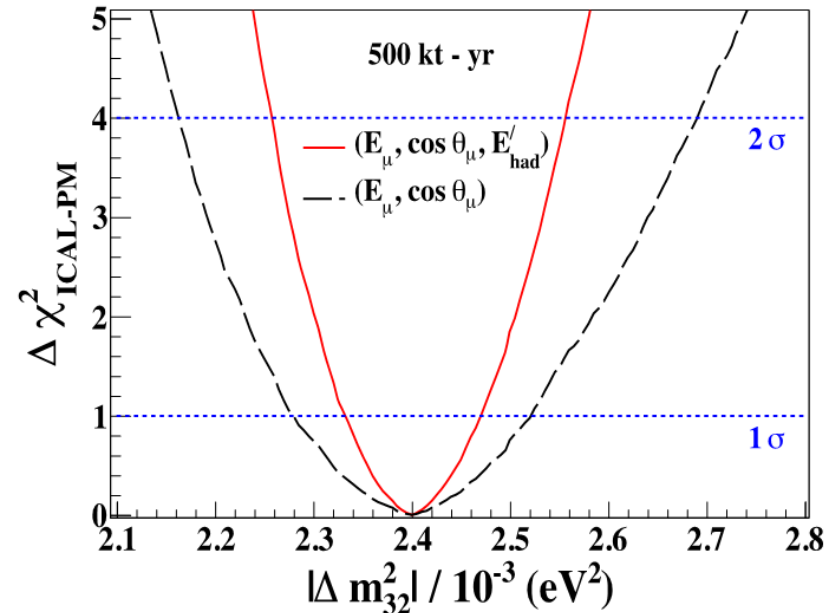
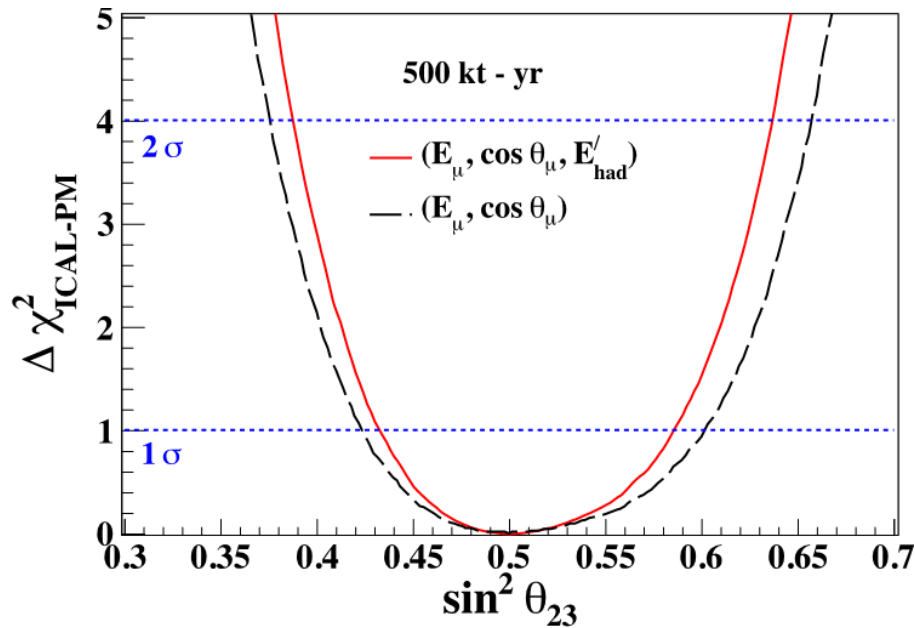
3σ median sensitivity can be achieved in 6 years ($\Delta\text{CP}=0$)

Synergy at other DeltaCP values



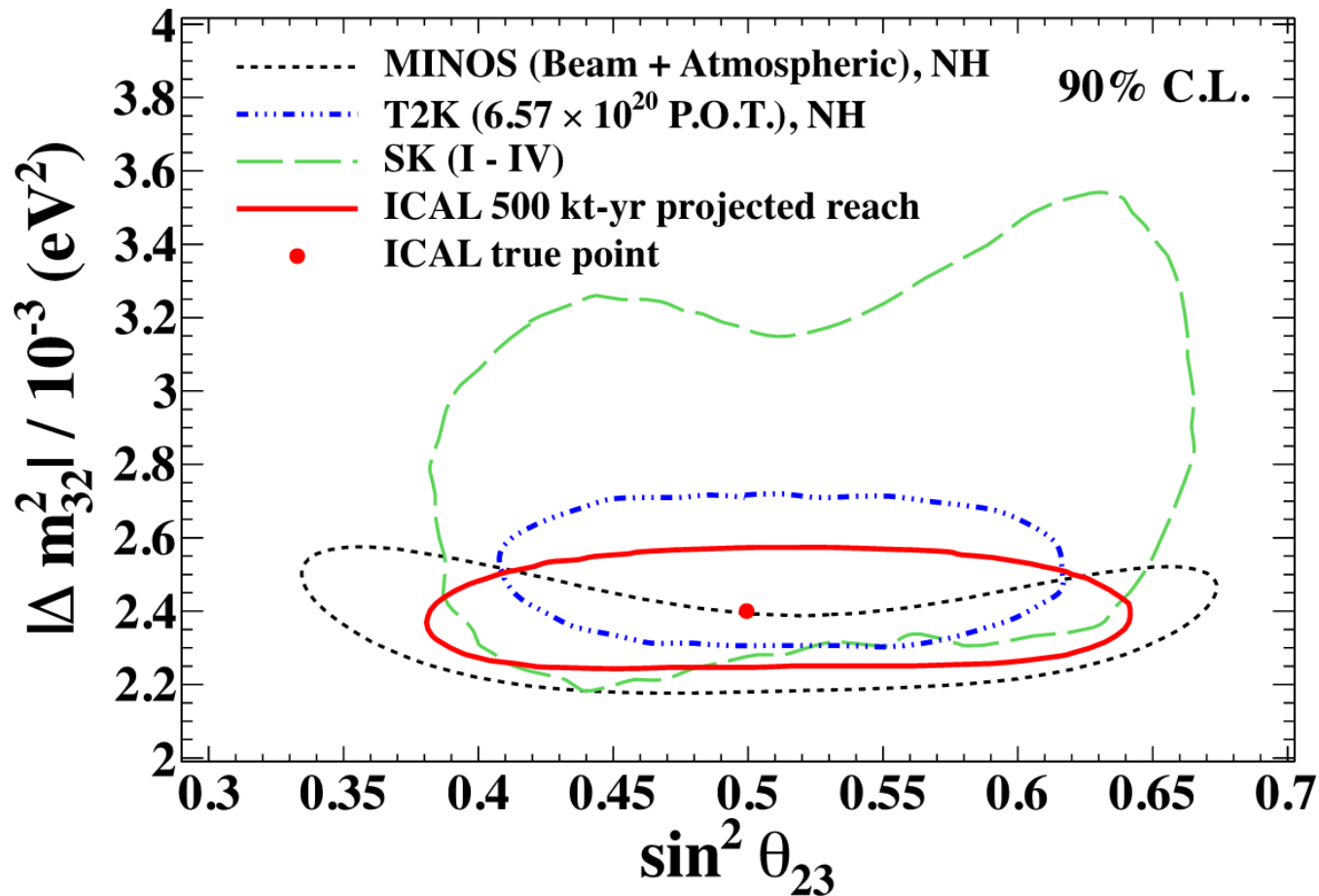
ICAL works even in the “unlucky” δ_{CP} range....

Sensitivity to atmospheric mixing parameters

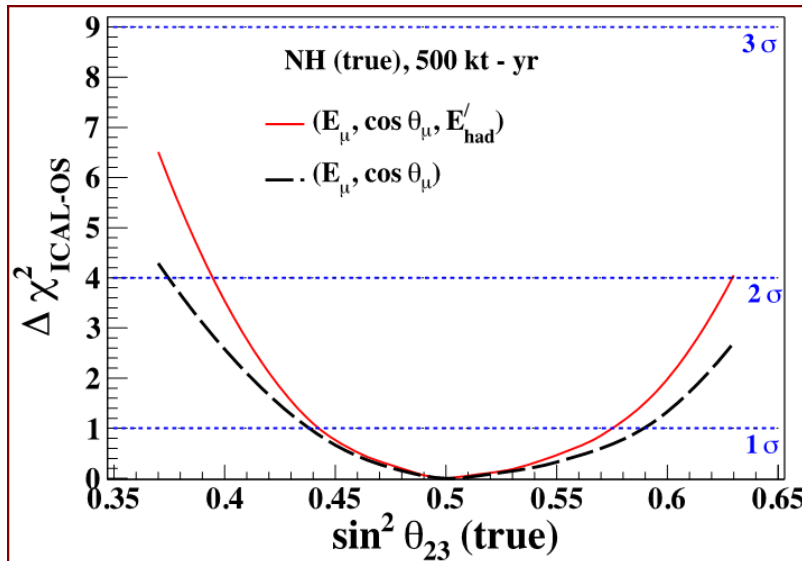
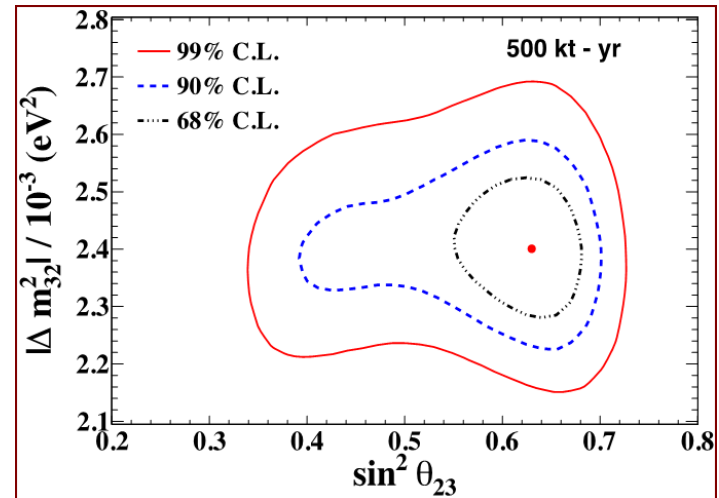
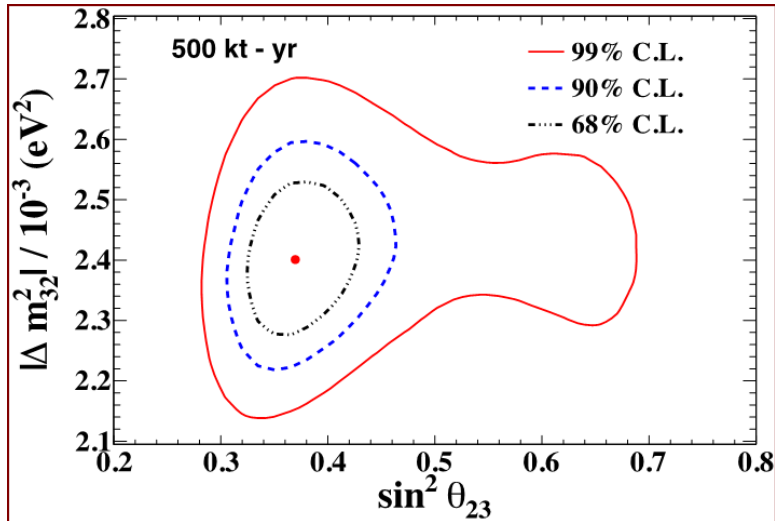


Significant improvement in the precision measurement of atmospheric mass splitting by adding hadron energy information with muon momentum

Reach for atmospheric mixing parameters



Octant sensitivity



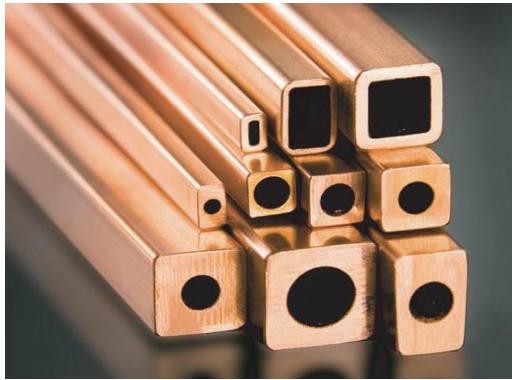
Median 2σ discovery of θ_{23} octant is possible if θ_{23} is sufficiently away from maximal value

Other analyses in progress

- *Search for sterile neutrinos*
- *CPT violation and Non-Standard Interactions*
- *Search for magnetic monopoles*
- *Search for dark matter from the Sun*
- *Long range forces*
- *Exploiting NC events*
- *Possibilities of electron detection*
-

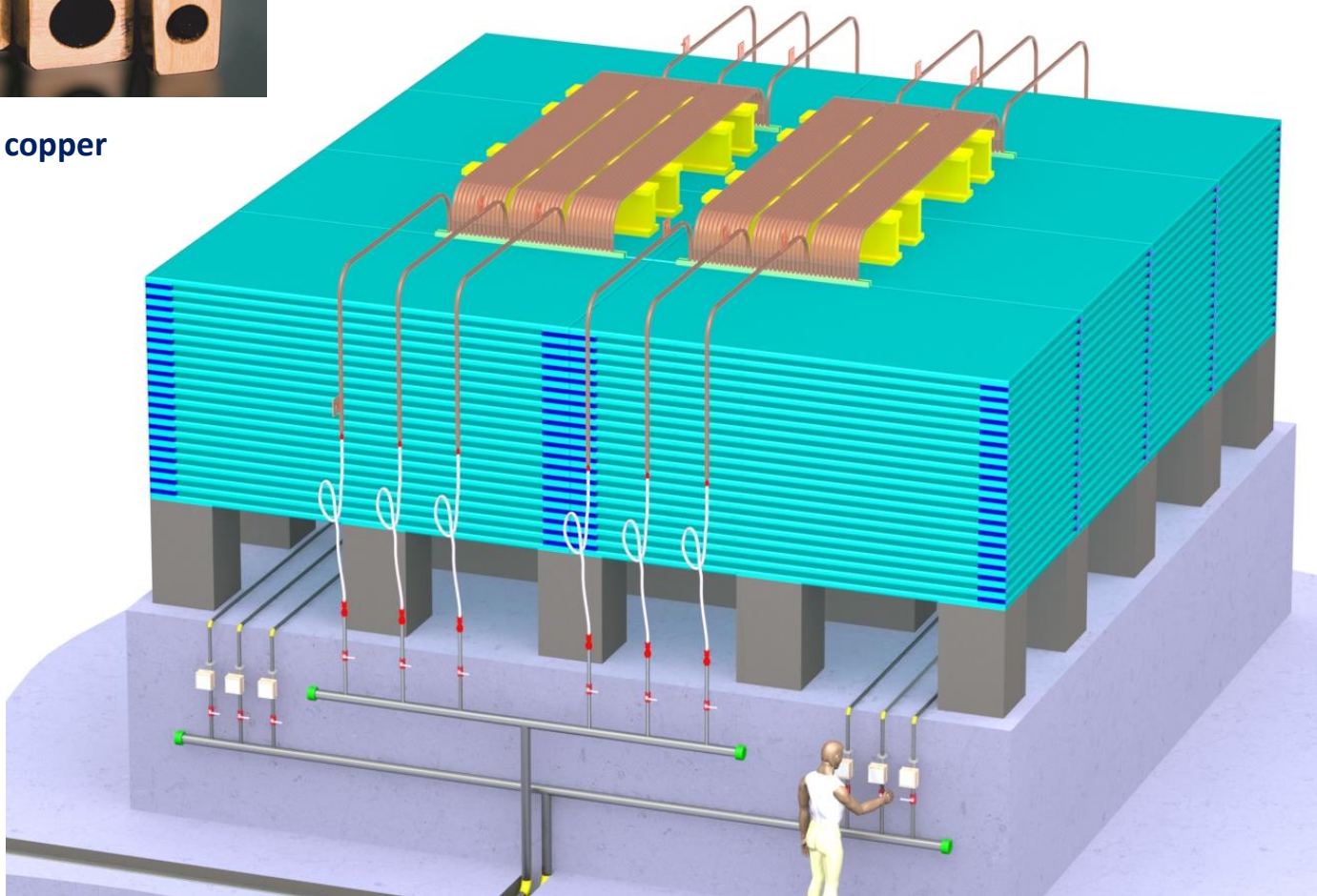
Current status of INO project

- **Site infrastructure development**
 - **Development of INO centre at Madurai city (110 km from underground lab)**
 - **Inter-Institutional Centre for High Energy Physics (IICHEP)**
 - **Construction of an 8m x 8m x 2.1 m engineering prototype module**
-
- **Detector R&D is now over**
 - **Detailed Project Report for Detector and DAQ system is ready**
 - **Soon go for industrial production of RPCs & associated front-end electronics**



Water cooled copper conductor

Prototype: the next step



The prospects

- Good muon tracking, charge ID and sensitivity to multi-GeV hadrons makes ICAL a unique multi-purpose detector
- Sensitive primarily to mass hierarchy, but also to many other new physics possibilities
- Synergy with ongoing and upcoming long baseline experiments
- Waiting for final go-ahead to start making the tunnel (t=0)
- The first module expected to start taking data at t=5 years

Thank You

- Collaborators are welcome...

<http://www.ino.tifr.res.in/ino>

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