

Neutrino Physics: Lecture 5

Neutrino oscillations in constant matter density

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Outline

- 1 Effective Hamiltonian in $2 - \nu$ system
- 2 Matter potential for neutrinos
- 3 $\nu_e - \nu_\mu$ evolution in matter: MSW resonance
- 4 $\nu_\mu \leftrightarrow \nu_{\text{sterile}}$ in atmospheric neutrinos?

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Effective Hamiltonian in vacuum

$$\begin{aligned}\nu_1 \rightarrow \nu_1 : \quad H_{11} &= p + m_1^2/(2E) \\ \nu_2 \rightarrow \nu_2 : \quad H_{22} &= p + m_2^2/(2E)\end{aligned}$$

$\nu_1 - \nu_2$ basis (mass basis):

$$\begin{aligned}H &= \begin{pmatrix} p + m_1^2/(2E) & 0 \\ 0 & p + m_2^2/(2E) \end{pmatrix} \\ &= p + \frac{m_1^2 + m_2^2}{2E} + \begin{pmatrix} -\Delta & 0 \\ 0 & \Delta \end{pmatrix}\end{aligned}$$

$$\Delta \equiv \frac{m_2^2 - m_1^2}{4E}$$

$$H_{\text{eff}} = \begin{pmatrix} -\Delta & 0 \\ 0 & \Delta \end{pmatrix}$$

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Evolution in vacuum: two flavours

$$i \frac{d}{dt} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} -\Delta & 0 \\ 0 & \Delta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$\begin{pmatrix} \nu_1(t) \\ \nu_2(t) \end{pmatrix} = \begin{pmatrix} e^{i\Delta t} & 0 \\ 0 & e^{-i\Delta t} \end{pmatrix} \begin{pmatrix} \nu_1(0) \\ \nu_2(0) \end{pmatrix}$$

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$$\nu_1(t) = e^{i\Delta t} \nu_1(0)$$

$$\nu_2(t) = e^{-i\Delta t} \nu_2(0)$$

Evolution in flavour basis

Change of basis:

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

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Evolution:

$$\begin{pmatrix} \nu_\alpha(t) \\ \nu_\beta(t) \end{pmatrix} = U^\dagger \begin{pmatrix} e^{i\Delta t} & 0 \\ 0 & e^{-i\Delta t} \end{pmatrix} U \begin{pmatrix} \nu_\alpha(0) \\ \nu_\beta(0) \end{pmatrix}$$

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Forward scattering interactions with matter

$$\nu_e \rightarrow \nu_e : \quad H_{ee} = V_C + V_N$$

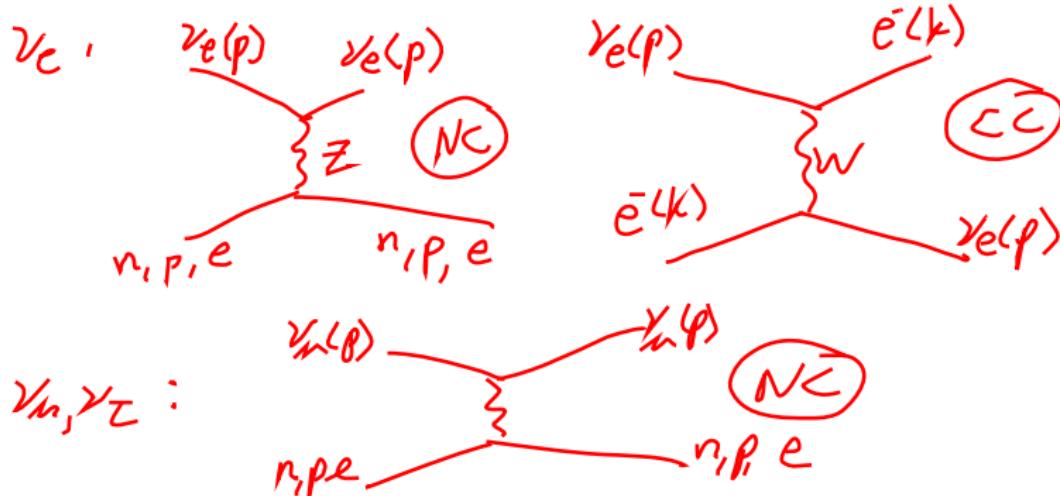
$$\nu_\mu \rightarrow \nu_\mu : \quad H_{\mu\mu} = V_N$$

$$\nu_\tau \rightarrow \nu_\tau : \quad H_{\tau\tau} = V_N$$

$$V_C = \sqrt{2} G_F n_e \quad , \quad V_N = -G_F n_n / \sqrt{2}$$

Detailed explanation: J. Linder, hep-ph/0504264

Forward scattering interactions with matter



$\nu_\mu, \nu_\tau :$

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Effective potential in matter

Three flavours:

$$V = \begin{pmatrix} V_C + V_N & 0 & 0 \\ 0 & V_N & 0 \\ 0 & 0 & V_N \end{pmatrix}$$

Two flavours (ν_e - $\nu_{\mu/\tau}$):

$$V_f = \begin{pmatrix} V_C & 0 \\ 0 & 0 \end{pmatrix}$$

Two flavours (ν_μ - ν_τ):

$$V_f = \begin{pmatrix} 0 & 0 \\ 0 & 0 \end{pmatrix}$$

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Evolution in flavour basis: constant density

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$$H_f = \begin{pmatrix} -\Delta \cos 2\theta + V_C & \Delta \sin 2\theta \\ \Delta \sin 2\theta & \Delta \cos 2\theta \end{pmatrix}$$

Evolution:

$$\begin{pmatrix} \nu_e(t) \\ \nu_\mu(t) \end{pmatrix} = e^{-i H_f t} \begin{pmatrix} \nu_e(0) \\ \nu_\mu(0) \end{pmatrix}$$

Not simple....

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Matter basis: where H is diagonal

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_m & \sin \theta_m \\ -\sin \theta_m & \cos \theta_m \end{pmatrix} \begin{pmatrix} \nu_{1m} \\ \nu_{2m} \end{pmatrix} = U_m \begin{pmatrix} \nu_{1m} \\ \nu_{2m} \end{pmatrix}$$

In matter basis:

$$H_m = U_m^\dagger H_f U_m = \text{Constant} \times I + \begin{pmatrix} -\Delta_m & 0 \\ 0 & \Delta_m \end{pmatrix}$$

Flavour evolution:

$$\begin{pmatrix} \nu_{1m}(t) \\ \nu_{2m}(t) \end{pmatrix} = \begin{pmatrix} e^{i\Delta_m t} & 0 \\ 0 & e^{-i\Delta_m t} \end{pmatrix} \begin{pmatrix} \nu_{1m}(0) \\ \nu_{2m}(0) \end{pmatrix}$$

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Evolution in flavour basis again

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Exactly like vacuum oscillations, with “m” suffix added !

$$P_{ee} = 1 - \sin^2(2\theta_m) \sin^2(\Delta_m L)$$

Need to determine Δ_m, θ_m

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Diagonalising H_f

$$H_f = \begin{pmatrix} -\Delta \cos 2\theta + V_C & \Delta \sin 2\theta \\ \Delta \sin 2\theta & \Delta \cos 2\theta \end{pmatrix}$$

Solve

$$H_m = U_m^\dagger H_f U_m = \text{Constant} + \begin{pmatrix} -\Delta_m & 0 \\ 0 & \Delta_m \end{pmatrix}$$

Mixing angle:

$$\tan 2\theta_m = \frac{\Delta \sin 2\theta}{2\Delta \cos 2\theta - V_C}$$

Mass squared difference in matter:

$$\Delta_m = \sqrt{(\Delta \cos 2\theta - V_C)^2 + (\Delta \sin 2\theta)^2}$$

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Mass squared difference in matter:

$$\Delta_m = \sqrt{(\Delta \cos 2\theta - \frac{V_C}{2})^2 + (\Delta \sin 2\theta)^2}$$

New amplitude, new wavelength

Vacuum → Resonance → High densities

$$\theta_m = 45^\circ$$

Δ_m minimum

$$\hookrightarrow \Delta \sin 2\theta$$

$$\sin^2 2\theta_m \approx 1$$

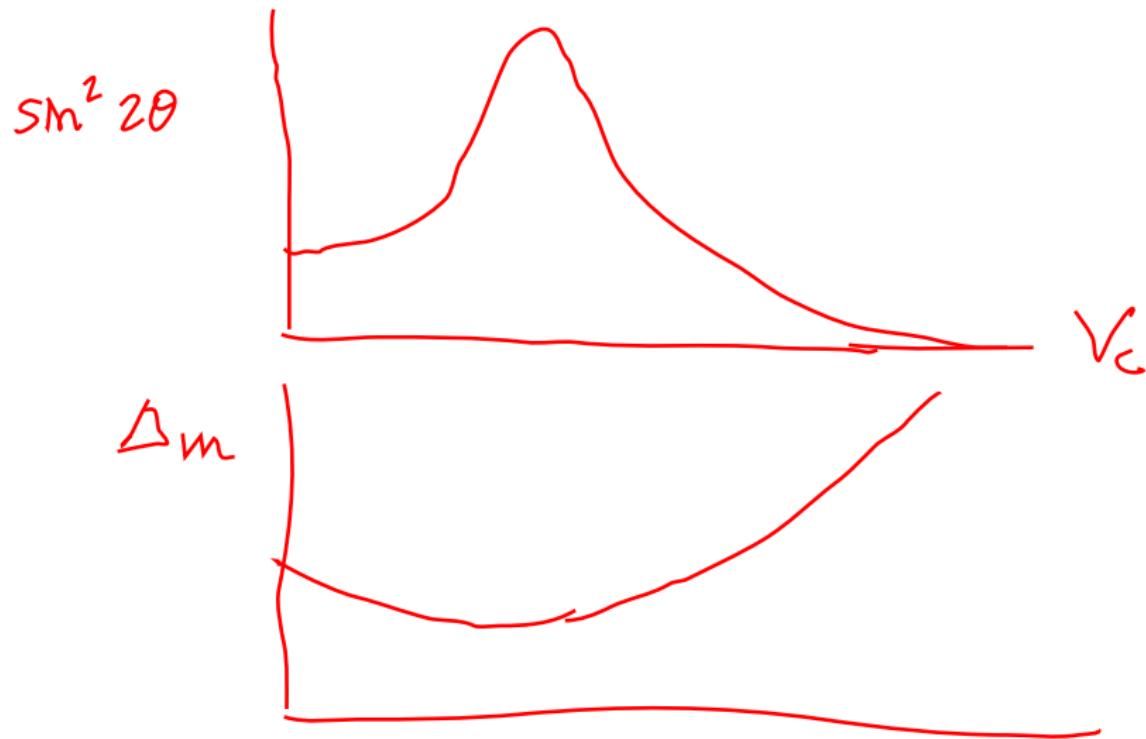
maximum

$$\theta_m = 90^\circ$$

$$\sin^2 2\theta_m \approx 0$$

NO oscillations

MSW resonance: minimum Δ_m , maximum $\sin^2 2\theta$



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Atmospheric oscillations: $\nu_\mu \rightarrow ?$

- A candidate: $\nu_\mu \leftrightarrow \nu_{\text{sterile}}$
- Matter potential ($\nu_\mu - \nu_s$ basis):

$$V_f = \begin{pmatrix} V_N & 0 \\ 0 & 0 \end{pmatrix}$$

- Oscillation probability different from vacuum oscillations

$$P_{\mu\mu} = 1 - \sin^2(2\theta_m) \sin^2(\Delta_m L)$$

- Oscillation amplitude depends on energy !

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Ruling out $\nu_\mu \leftrightarrow \nu_s$ in atmospheric neutrinos

