

Neutrino physics: an introduction

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Outline

- 1 Motivation
 - What are neutrinos
 - Current excitement in neutrinos

- 2 Aspects of neutrino physics

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What are neutrinos

- **Particles that accompany radioactive β decay**
- Byproducts of nuclear reactions
- The most abundant particles
- The lightest massive particles
- The most weakly interacting particles
- Particles that break left-right (mirror) symmetry maximally
- Particles that may be their own antiparticles
- Particles that have always given unexpected new results

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Current excitement in neutrinos

- Solar neutrino problem
- Atmospheric neutrino problem
- Cosmology: structure formation, CMBR anisotropy
- Matter-antimatter asymmetry

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Aspects of neutrino physics

- Postulation of neutrino: β decay kinematics
- Discoveries of the three neutrinos ν_e, ν_μ, ν_τ
- Direct measurement of neutrino mass
- Violation of left-right symmetry: Wu's experiment
- Atmospheric neutrino problem:
quantum mechanics of neutrino oscillations
- Solar neutrino problem:
matter effects on neutrino oscillations
- Open questions in neutrino physics
- Current activities in neutrino physics

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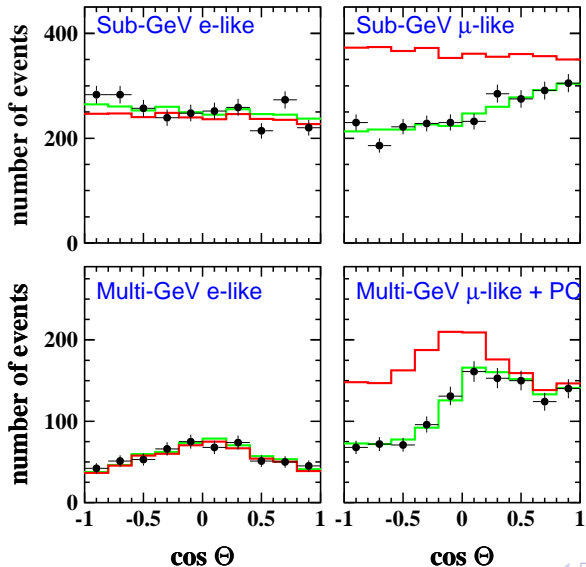
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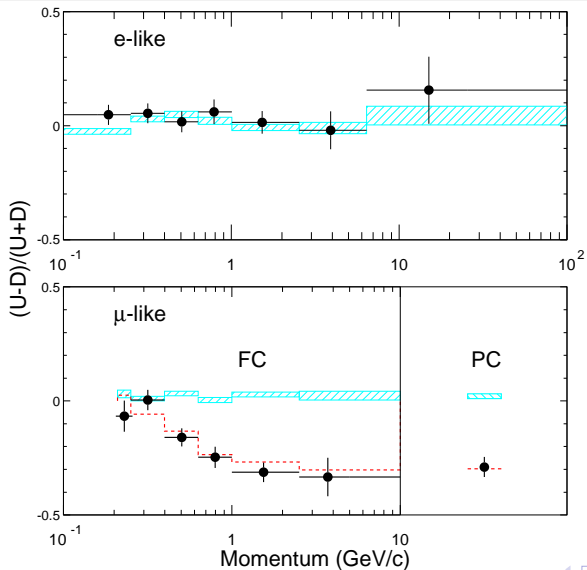
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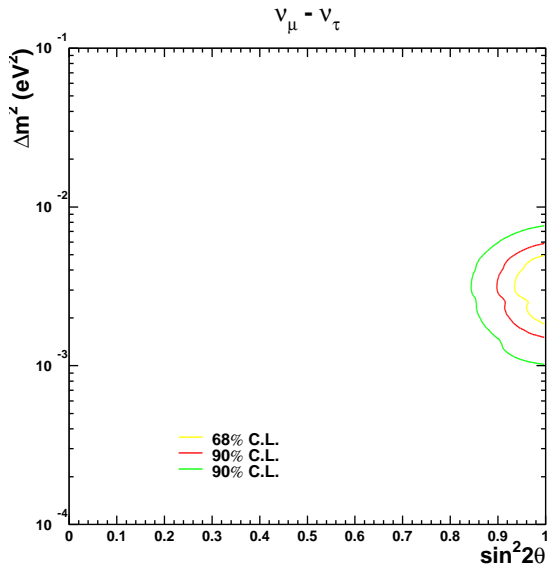
Zenith angle dependence



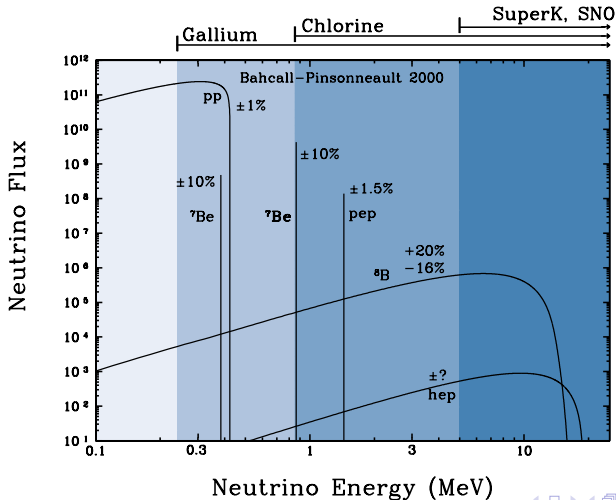
Up-down asymmetry



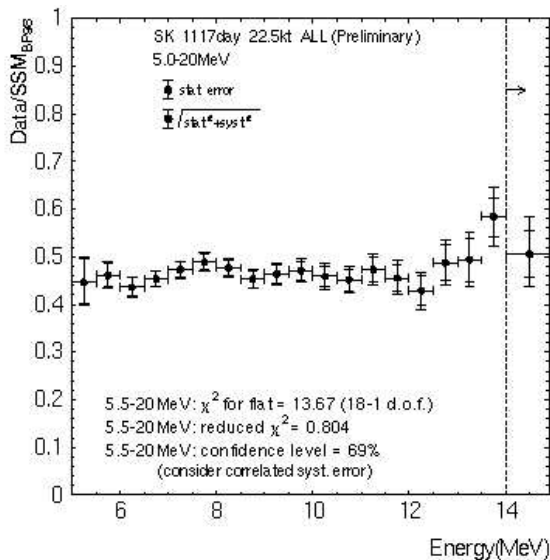
Solution to the atmospheric neutrino problem



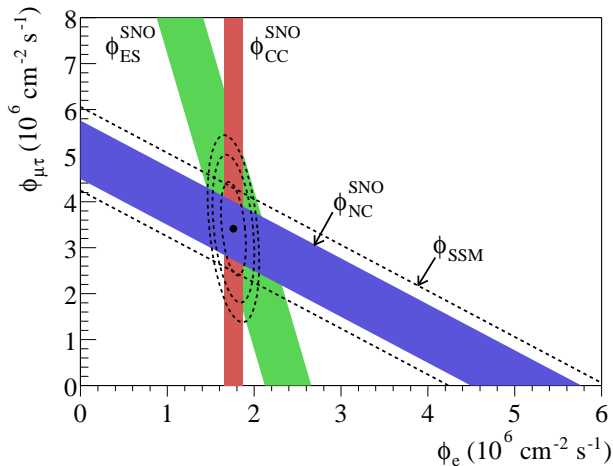
Spectra of solar neutrinos



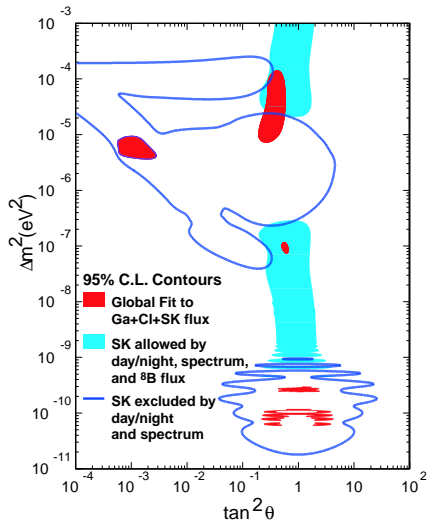
Survival probability of solar neutrinos



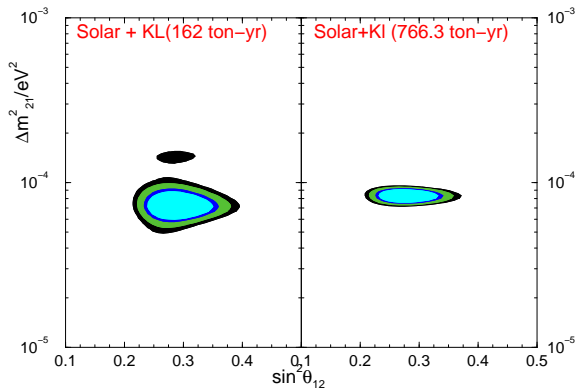
Solar ν_e convert to ν_μ and ν_τ



Solar neutrino parameters



Solar neutrino parameters



What we know about neutrinos

$$(\nu_e, \nu_\mu, \nu_\tau) \leftrightarrow (\nu_1, \nu_2, \nu_3)$$

Solar, Atmospheric and Reactor neutrino experiments \Rightarrow

- Mass squared differences: $\Delta m_{21}^2 \ll \Delta m_{31}^2 \approx \Delta m_{32}^2$

$$\Delta m_{\odot}^2 \approx (7.2 - 9.5) \times 10^{-5} \text{eV}^2, \quad \Delta m_{\text{atm}}^2 \approx (1.3 - 3.4) \times 10^{-3} \text{eV}^2$$

- Mixing matrix: $U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}$

$$|U_{e1}|^2 \approx 0.7, \quad |U_{e2}|^2 \approx 0.3, \quad |U_{\mu 3}|^2 \approx 0.5, \quad |U_{e3}|^2 < 0.05$$

(Two large angles and one small, possibly vanishing – and hence interesting – angle)

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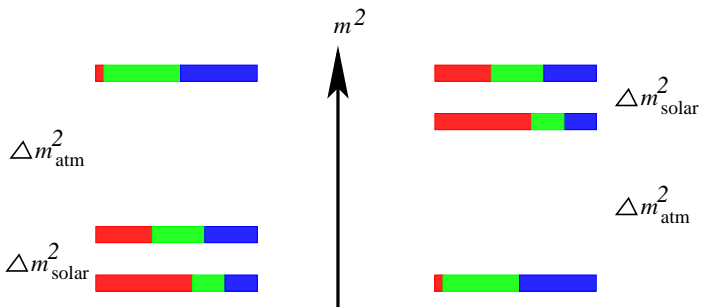
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What we do not know

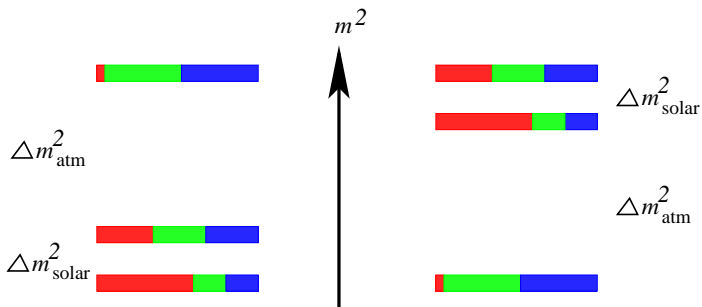
- Mass hierarchy: Normal or Inverted ?
(red ν_e , green ν_μ , blue ν_τ)



- CP violation
- Absolute neutrino masses

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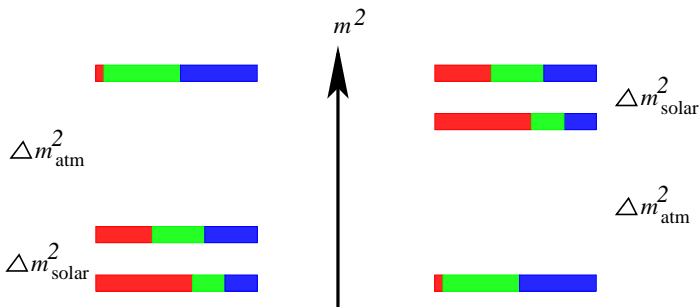
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Current activities in neutrino physics

- **Solar experiments:**
measuring the energy of the sun in neutrinos
- **Reactor experiments:**
Measuring the small mixing angle
- **Short baseline experiments:**
Confirming atmospheric oscillations with known ν fluxes
- **“Neutrino factories”:**
Long baseline experiments that span the Earth
- **Neutrino telescopes:** Looking for extremely energetic neutrinos from the cosmic rays
- **India-based Neutrino Observatory (INO):**
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