

# *Solar Model Parameters and Direct Measurement of Solar Neutrino Fluxes*

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# **OUTLINE**

- Standard Solar Model and its parameters
- Motivations for using neutrinos as probes to SSM parameters
- Neutrino flux – present knowledge and future measurements
- Techniques to determine SSM parameters using neutrino flux measurements
- Results
- Summary

# *Standard Solar Model & its parameters*

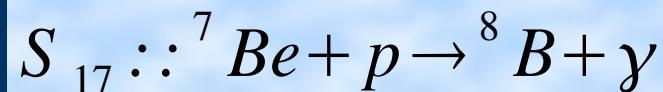
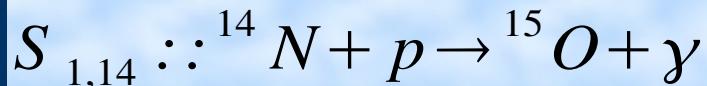
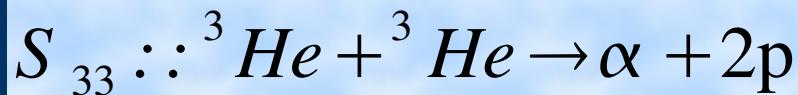
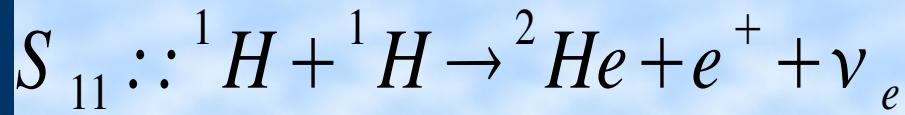
- Homogeneous cloud of **H**, **He**, small fraction of **heavier elements**
- Mass = **M sun**
- Evolved throughout a time **t = t sun**
- To get present day value of **L sun** and **R sun**

**Hydrostatic equilibrium**  
**Energy transport primarily by photon diffusion**

- Energy generations by nuclear reactions
- Abundance changes caused by nuclear reactions

# ***SSM & its parameters (contd..)***

## **Low E Cross Sections**



## **Astronomical parameters**

- **Solar Age**
- **Solar Luminosity**
- **Initial heavy element abundance: Z / X**
- **Diffusion**
- **Opacity**

# Solar Neutrino Flux – present knowledge and future measurements ( ${}^8 B$ )

•	<b>SSM Prediction:</b>	$\phi_B = 5.79(1 \pm 0.23) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$
•	<b>SNO NC:</b>	$\phi_B = 4.94(1 \pm 0.088) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$
•	<b>Global analysis of Solar+KamLAND data:</b>	$\phi_B = 4.88(1 \pm 0.036) \times 10^6 \text{ cm}^{-2} \text{ s}^{-1}$
		$\frac{\Delta \phi_B}{\phi_B}$
	<b>Data Set</b>	
	<b>Solar</b>	<b>4.4%</b>
	<b>+ KamLAND</b>	<b>3.6%</b>
	<b>+ SNO-III</b>	<b>3.2%</b>
	<b>+ pp</b>	<b>2.5%</b>
	<b>+ SPMIN</b>	<b>1.7%</b>

# *Solar Neutrino Flux – present knowledge and future measurements*

$(^7Be, pp)$

- $\phi_{^7Be}/(10^9 \text{ cm}^{-2} \text{ s}^{-1})$

**SSM : 4.16(1 +/- 0.12)**

$$\frac{\Delta \phi_{Be}}{\phi_{Be}}$$

**Solar+KL:**

worse

**+Borexino(10%): 10%**

**(+Borexino(5%)): 5.5%**

- $\phi_{pp}/(10^{10} \text{ cm}^{-2} \text{ s}^{-1})$

**SSM : 5.94(1 +/- 0.01)**

$$\frac{\Delta \phi_{pp}}{\phi_{pp}}$$

**Solar+KL:**

worse

**+ Lum Constraint: 2%**

**+Borexino(5%): 0.5%**

# *Dependence of Neutrino flux on SSM parameters [BP04]*

<http://www.sns.ias.edu/~bahcall>

$$\phi_B = Const \times S_{11}^{-2.59} S_{33}^{-0.40} S_{34}^{+0.81} S_{1,14}^{+0.01} S_{17}^{1.0} S_{e-7}^{-1.0} \\ L^{+6.76} \tau^{+1.28} Opac^{-2.93} D^{-2.2} (Z/X)^{+1.36}$$

$$\phi_{Be} = Const \times S_{11}^{-0.97} S_{33}^{-0.43} S_{34}^{+0.86} L^{+3.40} \tau^{+0.69} Opac^{-1.49} \\ D^{-0.96} (Z/X)^{0.62}$$

$$\phi_{pp} = Const \times S_{11}^{+0.14} S_{33}^{+0.03} S_{34}^{-0.06} S_{1,14}^{-0.02} L^{+0.73} \tau^{-0.07} \\ Opac^{+0.14} D^{+0.13} (Z/X)^{-0.08}$$

$$\phi_i = C_i \prod_{all\ j} x_j^{\alpha_{ij}}$$

$$\alpha_{ij} = \frac{\partial \ln \phi_i}{\partial \ln x_j}$$

# *Determination of SSM parameters using measured neutrino flux*

$$\phi_i = C_i \prod_{all\ j} x_j^{\alpha_{ij}} \quad i=pp, \text{ pep, hep, B, Be, N, O, F}$$

K = # of known fluxes      N = total # of input parameters

Use the set of K eqns. to solve for any K input parameters (x), taking rest (N-K) parameters from their laboratory measurements.

$$(\Delta \ln x_j)^2 = \sum_{i=1,k} P_{ir}^2 (\Delta \ln \phi_i)^2 + \sum_{j \neq 1,K} Q_{jr}^2 (\Delta \ln x_j)^2$$
$$\Delta \ln x_r = \ln \left( 1 + \frac{\Delta x_r}{x_r} \right)$$

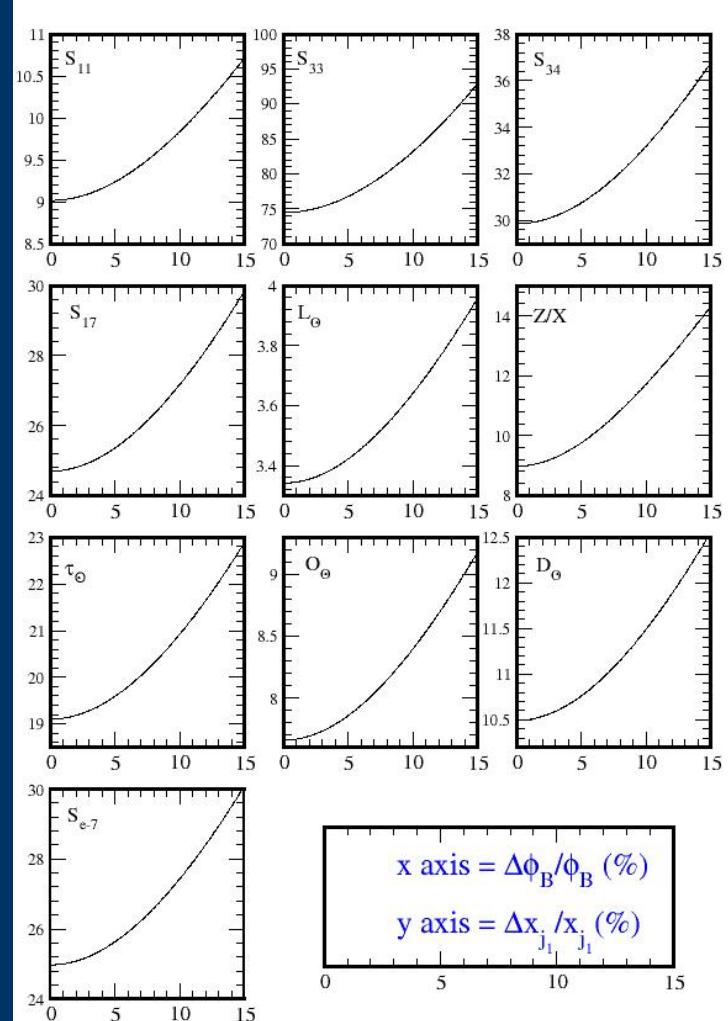
# *Results (Measured flux = B)*

- Good precision of  $S_{11}$ ,  $Z/X$ ,  $L$ , Opacity
- Uncertainty in  $Z/X$  is smaller than its currently estimated uncertainty in SSM
- Poor precision of  $S_{33}$  and  $S_{1,14}$

$x_j$	$\frac{\Delta x_j}{x_j} (\%)$			$\frac{\Delta x_j^0}{x_j^0} (%)$	$\frac{x_j}{x_j^0}$
	$\frac{\Delta \phi_B}{\phi_B} = 2\%$	$\frac{\Delta \phi_B}{\phi_B} = 3\%$	$\frac{\Delta \phi_B}{\phi_B} = 4\%$		
$S_{11}$	9.05	9.10	9.16	0.4	1.07
$S_{34}$	30.01	30.19	30.44	9.4	0.81
$S_{17}$	24.80	24.94	25.12	3.8	0.84
$L_\odot$	3.35	3.37	3.39	0.4	0.97
$Z/X$	9.11	9.27	9.48	15.0	0.88
Age	19.18	19.28	19.42	0.4	0.87
$O_\odot$	7.69	7.73	7.79	2.0	1.06
$D_\odot$	10.54	10.59	10.66	2.0	1.08
$S_{e-7}$	25.08	25.22	25.40	2.0	1.19

# *Results (Measured flux = B)*

- Model parameter uncerts depends strongly on 8B flux uncert if 8B flux uncert > 5%
- Model parameter uncert are stable below 4% of 8B uncert.



# *Results (Measured flux = B, Be)*

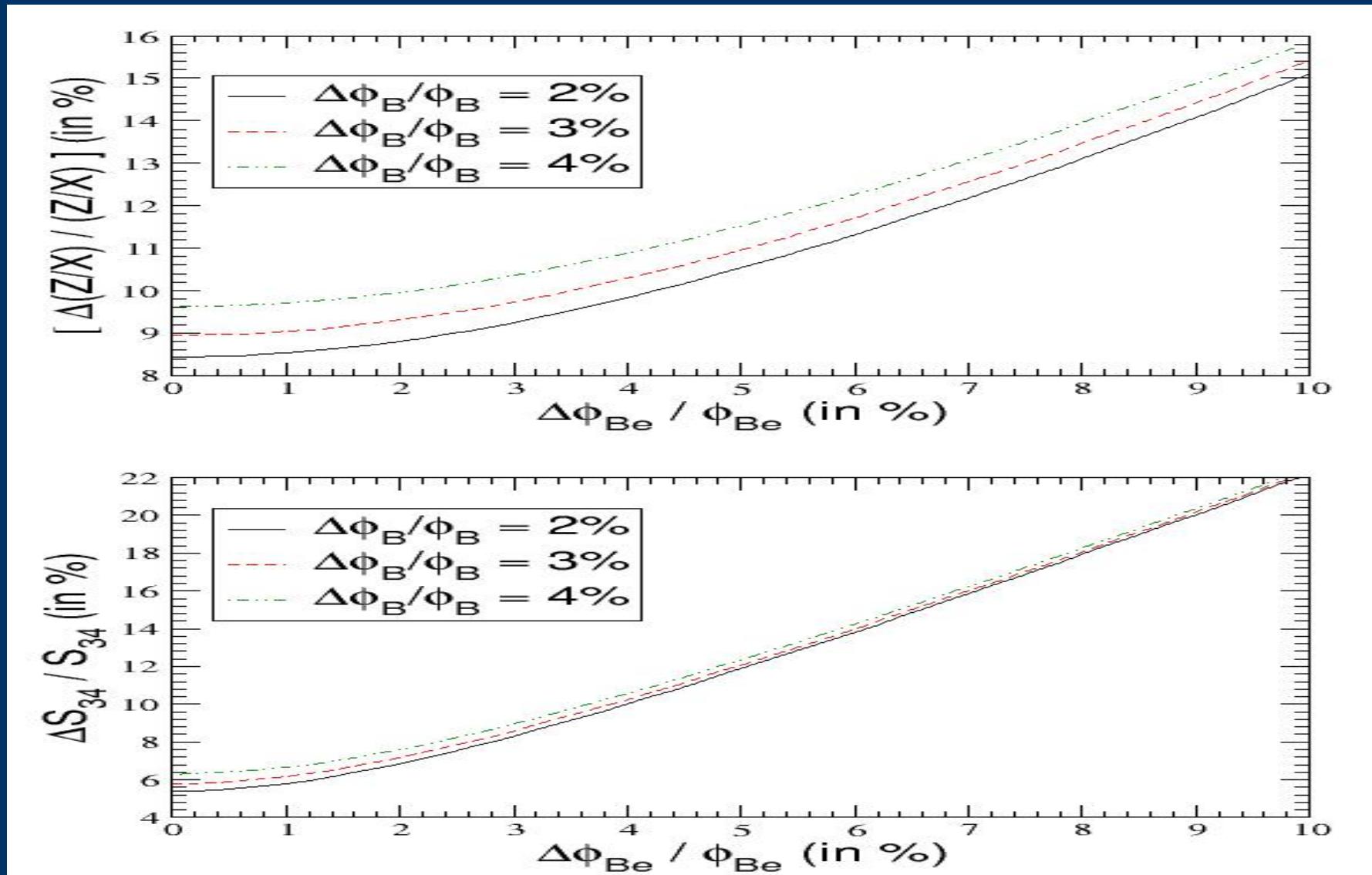
4% Uncert of B flux

6% Uncert of Be flux

- Good precision of S34 when attempted to measure with S11 or Z/X or D
- L is measured with good accuracy when attempted to measure with S17, S34, Se-7

combinations		$\frac{\Delta x_{j_1}}{x_{j_2}}$ (%)	$\frac{\Delta x_{j_2}}{x_{j_2}}$ (%)
$j_1$	$j_2$		
$S_{34}$	$S_{11}$	12.71	8.81
$S_{34}$	Z/X	14.25	12.26
$S_{34}$	$D_\odot$	13.79	11.35

# Results (Measured flux = $B$ , $Be$ )



# Results (Measured flux =B, Be, pp)

Uncerts in fluxes:

B -> 3%

Be -> 4%

pp -> 1%

Some Most effective combinations:

$\{S_{34}, L, S_{e-7}\} \rightarrow S_{34}$  (6.3%)  
(present~9.4%)

$\{S_{17}, L, Z/X\} \rightarrow Z/X$  (6.65%)  
(present~15%)

combinations			$\frac{\Delta x_{j_1}}{x_{j_2}}$ (%)	$\frac{\Delta x_{j_2}}{x_{j_2}}$ (%)	$\frac{\Delta x_{j_3}}{x_{j_3}}$ (%)
$x_{j_1}$	$x_{j_2}$	$x_{j_3}$			
$S_{11}$	$S_{34}$	$L_\odot$	6.16	7.98	1.62
$S_{11}$	$L_\odot$	$O_\odot$	13.67	1.72	12.25
$S_{34}$	$L_\odot$	$Z/X$	8.87	1.00	8.63
$S_{34}$	$L_\odot$	$O_\odot$	10.45	1.69	5.31
$S_{34}$	$L_\odot$	$D_\odot$	9.03	1.57	5.77
$S_{34}$	$L_\odot$	$\tau_\odot$	11.40	1.66	12.86
$S_{11}$	$S_{17}$	$L_\odot$	4.73	11.05	1.73
$S_{11}$	$L_\odot$	$S_{e-7}$	4.53	1.73	11.05
$S_{33}$	$S_{17}$	$L_\odot$	12.40	8.47	1.95
$S_{33}$	$L_\odot$	$S_{e-7}$	13.43	1.95	8.47
$S_{34}$	$S_{17}$	$L_\odot$	8.61	8.98	1.86
$S_{34}$	$L_\odot$	$S_{e-7}$	6.33	1.85	8.98
$S_{17}$	$L_\odot$	$Z/X$	9.26	1.21	6.65
$S_{17}$	$L_\odot$	$\tau_\odot$	9.05	1.85	7.34
$S_{17}$	$L_\odot$	$O_\odot$	9.21	1.84	3.40
$S_{17}$	$L_\odot$	$D_\odot$	9.74	1.74	4.68
$L_\odot$	$Z/X$	$S_{e-7}$	1.21	6.65	9.88
$L_\odot$	$\tau_\odot$	$S_{e-7}$	1.85	7.42	9.29
$L_\odot$	$O_\odot$	$S_{e-7}$	1.84	3.40	9.48
$L_\odot$	$D_\odot$	$S_{e-7}$	1.73	4.68	10.12