Possible Evidence of Thermalization at RHIC

Outline

Why we need to address the issue of thermalization

What are the characteristic signatures of thermalization

What have we learned so far at RHIC

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Thermalization : Why Address This Topic ?

(A) To establish Quark-Gluon Plasma (QGP)

"For our purposes here, we take the QGP to be a (locally) thermally equilibrated state of matter in which quarks and gluons are deconfined from hadrons, so that color degrees of freedom become manifest over nuclear, rather than merely nucleonic, volumes."

STAR Collaboration : Nucl. Phys. A 757 (2005) 102

Thermalization : Why Address This Topic ?

(B) To explore the QCD Phase Diagram

Phase diagram is a type of graph used to show the equilibrium conditions between the thermodynamically distinct phases ³

Temperature in the phase diagram has a meaningful definition for a system in thermal equilibrium

One of the goals of HI program is to "Establish the QCD Phase Diagram"



Thermalization : Why Address This Topic ?

(C) To understand several physics conclusions at RHIC Most theories in our field are based on assumption of thermal equilibrium or close to thermal equilibrium

(3)

(1)

STAR : Phys. Rev. Lett. 99 (2007) 112301



Recombination/ coalescence

(2) Lattice QCD assumes thermalization and provides EOS for many phenomenology/theoretical work

Boltzman equation for homogenous system, no external force and with relaxation time approximation $-df/dt = f - f_0/\tau_{relax}$ $f = f_0 + (f_i - f_0) \exp(-\Delta t/\tau_{relax})$

f --> f_0 (equilibrium dist) if $\Delta t >> \tau_{relax}$ Most models now days work with conditions near equilibrium.

Thermalization

In physics, thermalization is the process of particles reaching thermal equilibrium through mutual interaction.

In general the natural tendency of a system is towards a state of equipartition of energy or uniform temperature. This raises the system's entropy.

Basic Features Of A Thermalized System

What happens when we have a thermalized system :

(A) Maximum Entropy - dS/dt = 0; S : Entropy, t : real time

(Very ideal case : Processes are reversible -Initial <===> Final)
-- To show experimentally is challenging (impossible?)

(B) Momentum-Space distributions reach equilibrium values

-- Can we access this experimentally ?

(C) Interactions among constituents are large or saturate.

-- Can we demonstrate this experimentally ?

Exponential distribution does not necessarily mean we have a thermal system.



 $\begin{array}{l} \mbox{Multi-particle production process} \\ (no assumption of thermalization) \\ dN/dyd^2p_T \sim Exp (- p_T/E/2N)^* \\ & \mbox{From multi-particle phase space factor} \end{array}$

Thermal system $dN/dyd^2p_T \sim Exp(-p_T/E/3N)$

Factor of 3/2: invariant momentum space (d³p/E) not equal to Thermodynamic Momentum space (d³p).

Experimentally difficult to distinguish the two.

*Assumption : average matrix element square is not strongly p_T dependent.

Slope of p_T distribution : Two contributions - Random part + Collective part



Random Part ~ E/N Intensive quantity independent of system size

Collective part can also occur for systems away from equilibrium

Indicates final state interactions, which will eventually drive system towards equilibrium.

 $\tau_{\text{scattering}} < \tau_{\text{expansion}}$

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Comparison with pp and dAu leads to interpretations being inconclusive ?



m_T-exponential Boltzman law. q characterizes degree of equilibrium -- Tallis statistics



Suggests no collectivity in pp and peripheral Au+Au collisions

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Z. Tang et al., Phys.Rev.C79:051901,2009.

Single Particle Momentum Distribution

T.T. Chou, C. N. Yang and E. Yen Phys. Rev. Lett. 54, 510–513 (1985)

> "A concept of partition temperature is introduced in high-energy collisions. It is a natural mathematical consequence of the Darwin-Fowler method, and neither requires nor implies thermal equilibrium. A collision at a given incoming energy is described as an incoherent superposition of collisions with different partition temperatures."

Probability = $(d^3p/E)g(p_{\perp}) \exp(-E/T_p)$ where T_p will be called the *partition temperature*.

cations: The concept of T_p originates in (i) the δ function in (1) representing energy conservation, and (ii) the Darwin-Fowler method of steepest descent. Both clearly will survive any modifications of (1), and $\exp(-E/T_p)$ will always be one of the factors of the single-particle distribution for the nonleading particles.

STAR : Phys. Rev. C 74 (2006) 54902



Ideal hydrodynamics

 $R_x \sim 1/\sqrt{m_T}$

Viscosity breaks scaling

 $\begin{array}{l} R^2{}_L \thicksim \tau_0 \ T/m_T \ \text{---} \ 19/16 \ \Gamma_s/\tau_0 \\ \Gamma_s \ \text{: Sound attenuation length} \\ \tau_0 \ \text{--- Initial time} \end{array}$

Both Au+Au and p+p show scaling with $1/\sqrt{m_T}$

Conclusions inconclusive ?



Freeze-out Distributions in e⁺e⁻ and pp



Thermal description of hadron production in e+e- collisions. A. Andronic, et al., arXiv:0804.4132





Chemical Freeze-out Distributions

(A) Assumptions :

- 1. μ ,T are constant along freeze-out hyper surface
- Simultaneous freeze-out (but mean free path different for different particles)
- 3. Flow 4-velocity common to all particle species
- 4. Hadron masses and decay widths do not change with T



Loss of correlation due to interactions

For a thermal system :

Correlations (in momentum) reach equilibrium values Physical observables which are sensitive to interactions -Such as : Average elliptic flow saturate with density Momentum correlations saturate with density



Hints available : But we have we ruled out other physical processes (jets/minijets ?; PLB 567 (2003) 184)

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Freeze-Out Distributions

By definition a state in thermodynamic equilibrium has no knowledge of past

May be we should look at early time distributions to see if the partonic matter attended a thermalized state

Direct photon spectra



Early time Collectivity

Substantial portion of elliptic flow developed early



Transverse Momentum p₊ (GeV/c)

Low p_T : Heavier hadrons lower flow (~ hydrodynamic pattern)

High p_T : Flow grouped along baryon-meson lines (~ Hadronization by partonic recombination)
 All p_T : Flow similar for hadrons with strange and light quark (~ developed at partonic stage)
 Does this collectivity reflect enough partonic interactions to claim Thermalization ?

Saturation Of Interactions : v₂ Saturate

Part of it in STAR : Phys. Rev. C 66 (2002) 34904



Hints available, but U+U collisions in Run 11 Or LHC Heavy Ion program will hopefully settle the issue L. Kumar: ICPAQGP 2010



Interactions : Jet and Bulk



The away side jet traverses a large amount of matter. The interactions seem to drive particles from the two sources, jet fragmentation and the bulk medium, toward equilibration. This may in turn imply a high degree of thermalization within the medium itself.



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STAR: PRL 105 (2010) 022302

R. Gavai & S. Gupta arXiv:1001.3796

Event-by-event net-proton distribution

H. Ritter: ICPAQGP 2010

Assumptions: Net-proton ~ net-Baryon Thermalization Modelling: Chemical Freeze-out



Fluctuations

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Conclusions

Measurements @ top RHIC energies	Thermalization	Remark
PID spectra and mean p_T	Seems to	Not unique, Simple multi-particle production gives same feature. Tallis distribution ?
Particle Ratios	Seems to	Not unique & pp, e⁺e⁻, pA same
НВТ	Seems to	Not unique, pp same behaviour
PID v ₂ , away side <p<sub>T></p<sub>	Could be	Need models to confirm
V_2/ϵ vs. 1/S dN/dy V_2 vs. p_T	Hints	More measurements, V_2 vs. p_T saturates beyond 39 GeV
Fluctuations	Seems to	Agreement with Lattice QCD and HRG
Direct photon	Seems to	Hydro-based models explain data
p_T Correlations	Hints	Measurements in heavy quark sector

So far difficult to get a direct evidence from experiments. Cumulative evidences from joint theory-experiment comparisons is the way to establish thermalization at RHIC or need data from charm sector

Outlook: Possible direct evidence for light quark thermalization



Pressure, Flow, ...

Thermodynamic identity

 σ – entropy p – pressure U – energy V – volume $\tau = k_BT$, thermal energy per dof

 $\tau d\sigma = dU + pdV$

In A+A collisions, interactions among constituents *and* density distribution lead to: pressure gradient ⇒ collective flow

- ⇔ number of degrees of freedom (dof)
- ⇔ Equation of State (EOS)
- ⇔ cumulative *partonic* + *hadronic*