

Hot and Dense Matter in the RHIC-LHC Era TIFR, Mumbai January 12th 2008

## Radiative energy loss: problems and some new developments

Néstor Armesto Departamento de Física de Partículas and IGFAE Universidade de Santiago de Compostela Contents:



I. Introduction: radiative energy loss.

2. Successes and problems.

3. Recent attempts to go beyond (arXiv: 0710.3073 [hep-ph], JHEP to appear, with L. Cunqueiro, C.A. Salgado, *Santiago*, and W.-C. Xiang, *Wuhan and Bielefeld*).

4. Summary.

Radiative energy loss: problems and new ideas



# I. Introduction: radiative energy loss

### I.I.Theoretical setup.

I.2. Models.

Radiative energy loss: problems and new ideas

# I.I.Theoretical setup:





Two parameters define the medium: qhat or gluon density plus mean free path, and length (geometry, dynamical expansion).

Radiative energy loss: problems and new ideas: I. Introduction

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# I/2. BDMPS/GLV: static medium. (Majumder, nucl-th/0702066):



$$\omega \frac{dI}{d\omega d\mathbf{k}_{\perp}} = \frac{\alpha_s C_F}{(2\pi)^2 \omega^2} 2 \operatorname{Re} \int_0^\infty dy_I \int_{y_I}^\infty d\bar{y}_I e^{i\bar{q}(y_I - \bar{y}_I)} \\ \times \int d\mathbf{u} e^{-i\mathbf{k}_{\perp} \cdot \mathbf{u}} \exp\left(\frac{1}{2} \int_{\bar{y}_I}^\infty d\xi n(\xi) \sigma(\mathbf{u})\right) \\ \times \frac{\partial}{\partial \mathbf{y}} \cdot \frac{\partial}{\partial \mathbf{u}} \int_{\mathbf{y}=0=\mathbf{r}(y_I)}^{\mathbf{u}=\mathbf{r}(\bar{y}_I)} \mathcal{D}\mathbf{r} \exp\left[i \int_{y_I}^{\bar{y}_I} d\xi \frac{\omega}{2} \left(\dot{\mathbf{r}}^2 - \frac{n(\xi) \sigma(\mathbf{r})}{i\omega}\right)\right]$$

Exact solution unknown, two approximations: I. Harmonic oscillator (Brownian motion): multiple soft scatterings.  $\int d\xi n(\xi)\sigma(\mathbf{r}) \simeq \frac{1}{2}\hat{q}(\xi)\mathbf{r}^2$ 

2. Opacity expansion: N=I, single hard scattering, corrects Brownian motion. Comparison for massless and massive: SW '03,ASW '04. Radiative energy loss: problems and new ideas: I. Introduction

# I.2. Models (II):

**Physical Process** 

3. AMY: rates order  $\alpha_s$ , dynamical medium, no interference of emissions in/out medium, expansion.

4. GW(M): FF in DIS on nuclei, first corrections in , modification of DGL/AP splitting functions, virtuality (see also Majumder et al. '07).

 $\theta \sim k_{\perp}/k$   $p \gg T \xrightarrow{k_{\perp}/k} gT$   $t_{i}$   $t_{2}$   $t_{3}$   $t_{N}$  h = 0  $k_{\perp}/k$  h = 0  $h \gg T$  h = 0

$$\tilde{D}(z_{1,}\mu^{2}) = D(z_{1,}\mu^{2}) + \frac{\alpha_{s}}{2\pi} \int_{0}^{\mu^{2}} \frac{dl_{\perp}^{2}}{l_{\perp}^{2}} \int \frac{dy}{y} \left( \frac{1+y^{2}}{1-y} f(x,y,Q^{2},l_{\perp}) + V.C. \right) D(z_{1}/y,\mu^{2})$$

$$f = \frac{C_{A} 2\pi \alpha}{l_{T}^{2} + k_{T}^{2}} \frac{\int dy \, dy_{1} \, dy_{2} \left\langle A \left| \overline{\psi}(y) F(y_{1}) F(y_{2}) \psi(0) \right| A \right\rangle e^{i \, factors}}{N_{c} \, f^{A}(x)}$$



Radiative energy loss: problems and new ideas: 1. Introduction



Any number of gluon lines can attach like this.



### 2. Successes and problems:

2.1. Light hadrons: R<sub>AA</sub> and back-toback suppression. :-)

2.2. Non-photonic electrons and more differential observables. :-(

2.3. qhat: dependence on medium modeling :-(

2.4. Limitations of the formalism. :-(

Radiative energy loss: problems and new ideas



2.1. RAA and btb for light:

$$Q(p_{\perp}) = \frac{d\sigma^{\text{med}}(p_{\perp})/dp_{\perp}^2}{d\sigma^{\text{vac}}(p_{\perp})/dp_{\perp}^2}$$

 $= \int d\Delta EP(\Delta E) \left( \frac{d\sigma^{\text{vac}}(p_{\perp} + \Delta E)/dp_{\perp}^{2}}{d\sigma^{\text{vac}}(p_{\perp})/dp_{\perp}^{2}} \right) \qquad D_{h/q}^{(\text{med})}(x,Q^{2}) = \int_{0}^{1} d\epsilon P(\epsilon) \frac{1}{1-\epsilon} D_{h/q} \left( \frac{x}{1-\epsilon},Q^{2} \right) \\ \text{BDMS 'OI; Wang et al '96}$ 

### Medium modeling $\rightarrow <\tau_0 qhat>=1-15GeV^2$



Zhang et al '07  $\langle \hat{q}_0 \tau_0 \rangle \approx 2 \div 3 \text{ GeV}^2$ 



Radiative energy loss: problems and new ideas: 2. Successes and problems



2.1. RAA and btb for light:

$$Q(p_{\perp}) = \frac{d\sigma^{\mathrm{med}}(p_{\perp})/dp_{\perp}^{2}}{d\sigma^{\mathrm{vac}}(p_{\perp})/dp_{\perp}^{2}}$$

$$D_{h/q}^{(\text{med})}(x,Q^2) = \int_0^1 d\epsilon P(\epsilon) \frac{1}{1-\epsilon} D_{h/q} \left(\frac{x}{1-\epsilon}\right) d\epsilon P(\epsilon) \frac{$$

### BDMS '01; Wang et al '96 Medium modeling $\rightarrow <_{T_0}qhat >= |-|5GeV^2|$







Radiative energy loss: problems and new ideas: 2. Successes and problems

 $= \int d\Delta E P(\Delta E) \left( \frac{d\sigma^{\rm vac}(p_{\perp} + \Delta E)/dp_{\perp}^2}{d\sigma^{\rm vac}(p_{\perp})/dp_{\perp}^2} \right)$ 

### 2.2. e's, differential observ.:

BAA

10<sup>-1</sup>

- Heavy quarks radiate less: non-photonic electrons not conclusive: benchmark (Armesto et al '05), hadronization (Adil et al '06), collisional (Djordjevic et al '06), resonances (van Hees et al '06), dynamical medium (Djordjevic et al. '08),...
- PseudoFF not well understood: no broadening at high pt in the near side, trigger bias?



Radiative energy loss: problems and new ideas: 2. Successes and problems

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#### Néstor Armesto 2.3. qhat: medium modeling USC DE SANTIAGO DE COMPOSTE

$$\hat{q}(\xi) = K \cdot 2 \cdot \epsilon^{3/4}(\xi)$$

$$\langle \hat{q} \rangle = \frac{2}{L^2 - \tau_0^2} \int_{\tau_0}^{L} d\tau \tau \hat{q}_0 \frac{\tau_0}{\tau} \simeq \frac{2\tau_0 \hat{q}_0}{L} \approx \frac{\hat{q}_0}{2 \div 5} \quad \text{Gyulassy et al. '01,}$$

| Phenomenological implementation         | qhat (GeV²/fm)                 |
|---|--------------------------------|
| fixed length                            | <~l (average)                  |
| Woods-Saxon (PQM)                       | 4-14 (average)                 |
| dynamical medium<br>(Djordjevic et al.) | decreases                      |
| flow (Armesto et al., Baier<br>et al.)  | no effect                      |
| dilution                                | increases, factor 2-5          |
| hydro (Eskola et al., Bass<br>et al.)   | K~3-4, late times<br>important |

Radiative energy loss: problems and new ideas: 2. Successes and problems



# 2.4. Limitations of the formalism:



- Calculations done in the high-energy approximation: only soft emissions.
- Energy-momentum conservation imposed a posteriori in the single inclusive spectrum (SGLV; Salgado et al. '03).
- Multiple gluon emission: Quenching Weights (BDMS '01), independent (Poissonian) gluon emission: assumption!

$$P(\Delta E) = \sum_{n=0}^{\infty} \frac{1}{n!} \left[ \prod_{i=1}^{n} \int d\omega_i \frac{dI(\omega_i)}{d\omega} \right] \delta \left( \Delta E - \sum_{i=1}^{n} \omega_i \right) \exp \left[ -\int_0^{\infty} d\omega \frac{dI}{d\omega} \right]$$

- No role of virtuality in medium emissions (but GWM!).
- Medium and vacuum treated differently.

Radiative energy loss: problems and new ideas: 2. Successes and problems



## 3. Beyond:

Recent attempts to go beyond (arXiv:0710.3073 [hep-ph], JHEP to appear, with L. Cunqueiro, C.A. Salgado, Santiago and W.-C. Xiang, Wuhan and Bielefeld); also with G. Corcella (Pisa).

Motivation: to check radiative eloss, more differential and unbiased observables (particle correlations and jets) have to be studied (others: Borghini et al. '05-..., Wang et al. '01-..., Vitev '05, Polosa et al. '06) → Monte Carlo for in-medium parton branching.

3.1. Medium-modified splitting functions (SF) and Sudakovs.

- 3.2. Medium-modified DGLAP evolution of frag. funct. (FF).
- 3.3. Preliminary: PYTHIA with in-medium branching.

Radiative energy loss: problems and new ideas

# 3.1. Medium-modified SF and Use Composed States Sudakovs:

In the vacuum, the formalism gives collinear  $(z \rightarrow I)$  SFs:

$$\frac{dI^{\text{vac}}}{dz \, d\mathbf{k}_{\perp}^{2}} = \frac{\alpha_{s}}{2\pi} \frac{1}{\mathbf{k}_{\perp}^{2}} P^{\text{vac}}(z), \quad P^{\text{vac}}(z) \simeq \frac{2C_{R}}{1-z} \qquad \omega = (1-z)E \text{ and } \mathbf{k}_{\perp}^{2} = z(1-z)t$$
In the medium, we make the analogy (ansatz!!!) (Polosa et al. '06):  

$$P^{\text{tot}}(z) = P^{\text{vac}}(z) + \Delta P(z,t)$$

$$\Delta P(z,t) \simeq \frac{2\pi t}{\alpha_{s}} \frac{dI^{\text{med}}}{dz dt}$$
Medium-modified Sudakovs:  

$$\Delta_{i}(t) = \exp\left[-\int_{t_{0}}^{t} \frac{dt'}{t'} \int_{z_{\min}(t')}^{1-z_{\min}(t')} dz \frac{\alpha_{s}(t',z)}{2\pi} \sum_{j} P_{i\rightarrow j}(z,t')\right] \prod_{\substack{v \in \mathbb{C} \\ v \neq v = 1 \\ v$$

0.1

10

3-flavor coupling with scale k<sub>T</sub><sup>2</sup>; different small-z extensions.

Radiative energy loss: problems and new ideas: 3. Beyond

 $10^{3}$  t (GeV<sup>2</sup>)

10<sup>2</sup>

# 3.2. Medium modified DGLAP evolution of FF (I):

Medium-modified DGLAP evolution of FF (from KKP IC):

$$D(x,t) = \Delta(t)D(x,t_0) + \Delta(t)\int_{t_0}^t \frac{dt_1}{t_1}\frac{1}{\Delta(t_1)}\int \frac{dz}{z}P(z)D\left(\frac{x}{z},t_1\right)$$



Radiative energy loss: problems and new ideas: 3. Beyond

 $D^{\mathrm{med}}(x,t_0) = D^{\mathrm{vac}}(x,t_0)$ 



Improvements: virtuality in medium emissions, medium and vacuum treated on the same footing, energy momentum conservation.

Drawbacks: formation time of the gluons does not affect the medium length seen by the radiating partons; no elastic scattering, no conversions included.





Radiative energy loss: problems and new ideas: 3. Beyond

## 3.2. Medium modified DGLAP evolution of FF (II):

Q<sup>2</sup>=1600 GeV<sup>2</sup>

10<sup>-1</sup>

Dashed - Quenching weights

10<sup>-4</sup>

10<sup>-5</sup>

10

Comparison with experimental data gives qhat~l GeV<sup>2</sup>/fm (as with QW for fixed L) or qhat~10 GeV<sup>2</sup>/ fm (for cylinder or sphere).



1 10<sup>1</sup>

10-2

10-3

10-4 Ē 10<sup>-5</sup>

10

10

10



0.4

0.2

0

1

z

10<sup>-1</sup>

1

z

### 3.3. Medium-mod. PYTHIA (I):

 We modify PYTHIA FSR (pyshow) routine introducing the medium-modified splittings and Sudakovs.

Good agreement between default and vacuum; medium enhancement at intermediate, decrease at large



Radiative energy loss: problems and new ideas: 3. Beyond

### 3.3. Medium-mod. PYTHIA (II):

For just one splitting, clear p<sub>T</sub> and angular broadening (Vitev '06, Salgado-Polosa '06): importance of multiple splitting.



Radiative energy loss: problems and new ideas: 3. Beyond

### 3.3. Medium-mod. PYTHIA (III):

At parton level and for high energy, high multiplicity enhancement and modest broadening.

![](_page_19_Figure_3.jpeg)

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### 3.3. Medium-mod. PYTHIA (IV)

A extreme example of hadronization kills most of the multiplicity enhancement, medium effects (soft stuff) less

![](_page_20_Figure_4.jpeg)

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# 4. Summary:

![](_page_21_Picture_2.jpeg)

- To check radiative energy loss as the explanation for jet quenching, differential probes needed: relation energy degradation / radiation enhancement / p<sub>T</sub> broadening.
- We have supplemented vacuum splitting functions with medium terms, based on an analogy with radiation spectra: virtuality, energy conservation, and vacuum and medium treated on the same footing.
- A modified DGLAP evolution for fragmentation functions has been performed (proof of principle), and its compatibility with QWs for high virtualities showed.
- A medium-modified parton shower is under development, required for correlations and jet shape studies: LHC.

Néstor Armesto 3. Determinations of  $\hat{q}(l)$ : USC UNIVERSIDADE SANTIAGO E COMPOSTE

- qhat is a natural parameter only in BDMPS.
- Extraction from a comparison with RAA.
- Phenomenological implementations are key: mean eloss rudimentary, distribution of energy losses better: quenching weights (BDMS, GLV '01).
- Fixed length (GLV; Arleo '02; SW'
   '03) gives ~< | GeV<sup>2</sup>/fm.

On the determination of the transport coefficient.

$$Q(p_{\perp}) = \frac{d\sigma^{\mathrm{med}}(p_{\perp})/dp_{\perp}^2}{d\sigma^{\mathrm{vac}}(p_{\perp})/dp_{\perp}^2} = \int d\Delta E P(\Delta E) \left(\frac{d\sigma^{\mathrm{vac}}(p_{\perp} + \Delta E)/dp_{\perp}^2}{d\sigma^{\mathrm{vac}}(p_{\perp})/dp_{\perp}^2}\right)$$

### BDMS '01; Wang et al '96

$$D_{h/q}^{(\text{med})}(x,Q^2) = \int_0^1 d\epsilon P(\epsilon) \frac{1}{1-\epsilon} D_{h/q}\left(\frac{x}{1-\epsilon},Q^2\right)$$

![](_page_22_Figure_9.jpeg)

#### Néstor Armesto 3. Determinations of $\hat{q}(I)$ : USC DETERMINATION OF $\hat{Q}(I)$ : USC DETERMINATI

### • A Woods-Saxon

geometry (production plus 'medium') gives larger values and leads to saturation: fragility (Dainese et al, Eskola et al '04).

- Surface bias (Muller '03).
- <qhat>=4-14 GeV<sup>2</sup>/fm.
- Energy constraints
   (Baier et al '06); energy
   dependence (Casalderrey)

et al '07

![](_page_23_Figure_7.jpeg)

#### Néstor Armesto 3. Determinations of $\hat{g}(\Pi)$ : $USC_{UNVERSIDADE}$ Determinations of $\hat{g}(\Pi)$ : $USC_{UNVERSIDADE}$ DE SANTIAGO DE COMPOSTELA

- Hard probes '06: AMY gives 2, GLV gives <1, MW give 3-4 GeV<sup>2</sup>/fm: all at initial time.
- Dilution: introduced effectively (GLVW '01, SW '02)  $\langle \hat{q} \rangle = \frac{2}{L^2 - \tau_0^2} \int_{\tau_0}^{L} d\tau \tau \hat{q}_0 \frac{\tau_0}{\tau} \simeq \frac{2\tau_0 \hat{q}_0}{L} \approx \frac{\hat{q}_0}{2 \div 5}$

• Flow (Armesto et al '04) doesn't lower qhat (Baier et al '06).

• A dynamical medium decreases qhat (AMY?, Djordjevic et al '07).  $\hat{q}(\xi) = K \cdot 2 \cdot \epsilon^{3/4}(\xi)$ 

 A dynamical expansion (Hirano-Nara '03; Ruppert-Renk '05, '06; Majumder et al '07; Qin et al '07) lowers qhat with respect to a static medium; still K>I; late time effect?

#### Néstor Armesto 3. Determinations of $\hat{g}(V)$ : USCUNVERSIDADE DE SANTIAGO DE COMPOSTEL

 Non-photonic electrons not conclusive: benchmark (Armesto et al '05), hadronization inside (Adil et al '06), collisional (Djordjevic et al '06)...

 IAA or away side pseudoframentation function (Wang '03) tend to favor low values of qhat (Renk '06; Loizides '06; Zhang et al '07): punch-through. (ĝo To)

![](_page_25_Figure_4.jpeg)

![](_page_25_Figure_5.jpeg)

![](_page_25_Figure_6.jpeg)

### 4. An exercise (I):

![](_page_26_Picture_2.jpeg)

(with Carlos A. Salgado, Rome La Sapienza)

Quantification of the effect on qhat of some of the phenomenological ingredients, based on R<sub>AA</sub> for central, using a pQCD spectrum and QW.

$$\mathcal{Q}(p_{\perp}) = \frac{d\sigma^{\mathrm{med}}(p_{\perp})/dp_{\perp}^{2}}{d\sigma^{\mathrm{vac}\,(p_{\perp})/dp_{\perp}^{2}}}$$

$$= \int d\Delta E P(\Delta E) \left( \frac{d\sigma^{\rm vac}(p_{\perp} + \Delta E)/dp_{\perp}^2}{d\sigma^{\rm vac}(p_{\perp})/dp_{\perp}^2} \right)$$

$$P(\Delta E) = \sum_{n=0}^{\infty} \frac{1}{n!} \left[ \prod_{i=1}^{n} \int d\omega_i \frac{dI(\omega_i)}{d\omega} \right] \delta \left( \Delta E - \sum_{i=1}^{n} \omega_i \right) \exp \left[ -\int_0^{\infty} d\omega \frac{dI}{d\omega} \right] \omega_c = \frac{1}{2} \hat{q} L^2, \quad R = \omega_c L, \quad L/\lambda = 1$$

single hard

#### multiple soft

![](_page_26_Figure_9.jpeg)

![](_page_26_Figure_10.jpeg)

DD@200, PHENIX DIO

### 4. An exercise (II): fixed length

![](_page_27_Picture_2.jpeg)

 $R_{AA}(p_T)$  for  $\pi^0$  at  $\eta = 0$ PHENIX: nucl-ex/0510023 0.9 AuAu at 200 GeV∕A, 0−10% 0.8 solid: multiple soft dashed: single hard 0.7 qhat=0.5, 1, 2 GeV<sup>2</sup>/fm 0.6 fixed L=4.3, 5.6 fm 0.5 0.4 0.3 0.2 0.1 0 18 8 10 12 14 16 20 p<sub>T</sub> (GeV)

![](_page_27_Figure_4.jpeg)

### 4. An exercise (III): Woods-Saxon

![](_page_28_Picture_2.jpeg)

$$\omega_c(\mathbf{r_0}, \phi) = \int_0^\infty d\xi \xi \hat{q}(\xi)$$
$$\langle \hat{q} L \rangle(\mathbf{r_0}, \phi) = \int_0^\infty d\xi \hat{q}(\xi)$$

 $\hat{q} \propto T_A T_B(x_0 + \xi \cos \phi, y_0 + \xi \sin \phi)$ 

### $R(\mathbf{r}_0,\phi) = 2\omega_c^2(\mathbf{r}_0,\phi)/\langle \hat{q}L \rangle(\mathbf{r}_0,\phi), \quad L = R/\omega_c, \quad \langle \hat{q} \rangle = 2\omega_c^2/(LR)$

![](_page_28_Figure_6.jpeg)

![](_page_28_Figure_7.jpeg)

#### Néstor Armesto **4.** An exercise (IV): hydro Hirano-Nara: 3+1 ideal hydro, for AuAu@200, b=3.1 fm, ideal EOS with N<sub>f</sub>=3, B<sup>1/4</sup>=247 MeV. **4.** An exercise (IV): hydro $\langle \epsilon \rangle(\tau_0) \simeq 27(36) \text{ GeV/fm}^3$ $\langle \epsilon^{3/4} \rangle(\tau_0) \simeq 1.5(2.2) \text{ GeV}^2/\text{fm}$ $\tau_0 = 0.6 \text{ fm}, \tau_{max} = 10.2 \text{ fm}$

![](_page_29_Figure_1.jpeg)

| Néstor Armesto<br>2.3. qhat: medium modeling USC<br>E COMPOSTELA |                        |                       |
|--|------------------------|-----------------------|
| $\hat{q}(\xi) = K \cdot 2 \cdot \epsilon^{3/4}(\xi)$             |                        |                       |
|  |                        | qhat (GeV²/fm)        |
|  | fixed length           | <=1 (average)         |
|  | Woods-Saxon            | 4-14 (average)        |
| Phenomenological   | dynamical medium       | decreases             |
| implementation   | flow                   | no effect             |
|  | dilution               | increases, factor 2-5 |
|  | hydro                  | K~3-4, late times     |
|  | I <sub>AA</sub> /pff   | favors low values     |
|  | non-photonic electrons | unconclusive          |
| Ť  | AMY                    | 2 (initial)           |
|  | M\A/                   | 2 2 (initial)         |

 Models
 MW
 2-3 (initial)

 multiple soft/single hard
 small decrease

 GLV
 <1 (initial)</td>

Radiative energy loss: problems and new ideas: 2. Successes and problems