

# Strongly Lensed Supernovae: New cosmological and astrophysical probes in the time-domain era

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with: LSST DESC + iPTF/ZTF Cosmo WG

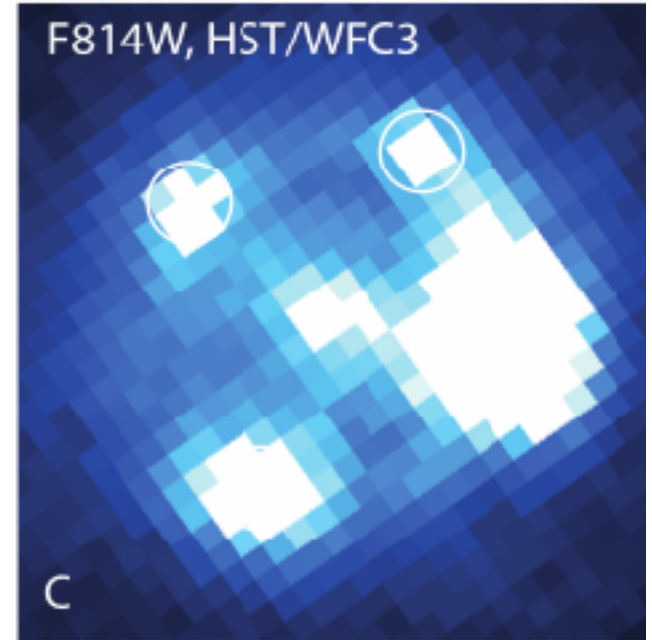


SOTU Seminar, TIFR, 19 December 2023





# Outline



## Motivation

### Recent discoveries

[SD+20b](#); [Mortsell, SD et al., 2021](#); [Johansson, .. SD, et al. 2021](#); [Goobar et al., Nat. As, 2023](#)

### Towards precision cosmology with LSST

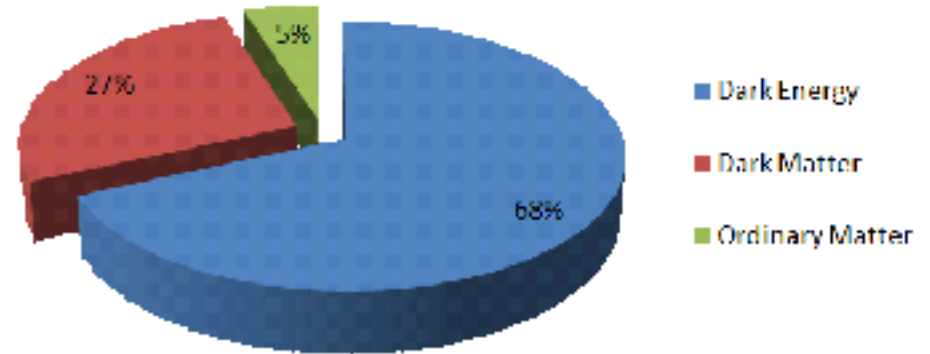
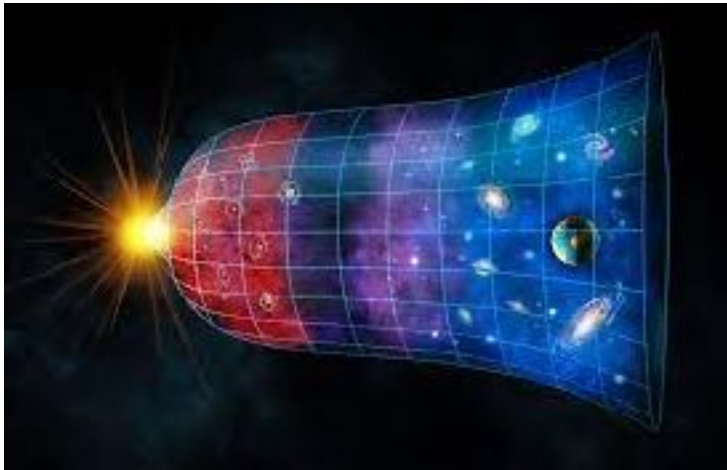
[Arendse, SD, et al.; SD+Pierel in prep., Birrer, SD, Shajib, 2022](#)

### Ongoing work: time-delay inference, JWST, curvature

[Crascall-Kennedy, SD, et al. in prep., Hu + SD to be submitted](#)



# Motivation



Well tested standard model:  $\Lambda$ CDM

Theoretical problems: what is dark energy + dark matter?

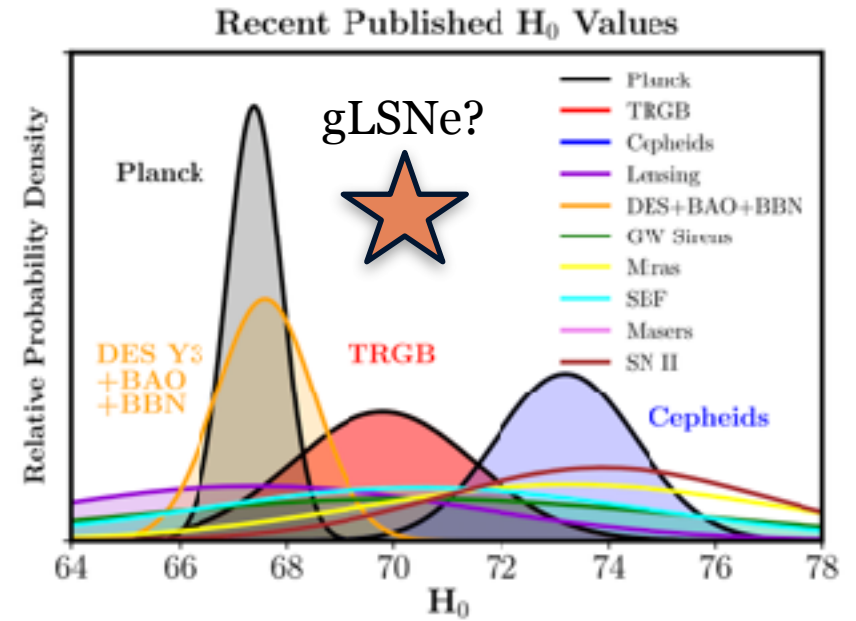
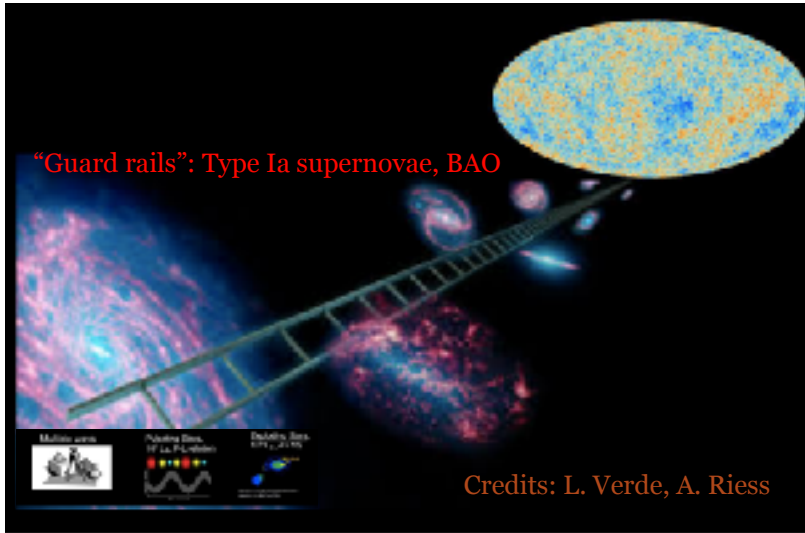
Observational problems appearing with increased precision



# Motivation

- $H_0$ : Absolute scale of the universe
- End-to-end test of background expansion

See also Scolnic+2023



- New physics?
- Unknown Systematics?



Need independent methods

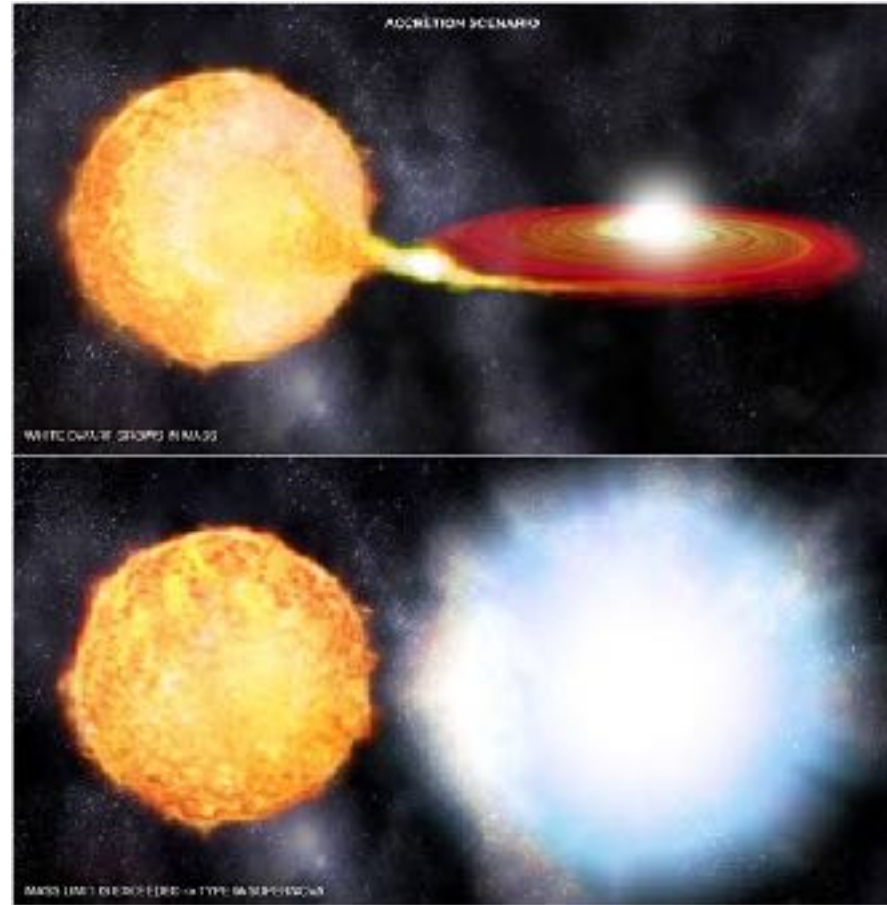
- Novel absolute distance measurement (e.g. lensed transients)





# What are Type Ia supernovae?<sup>5</sup>

NOT standard but calibratable -> small scatter, reduce lensing uncertainties

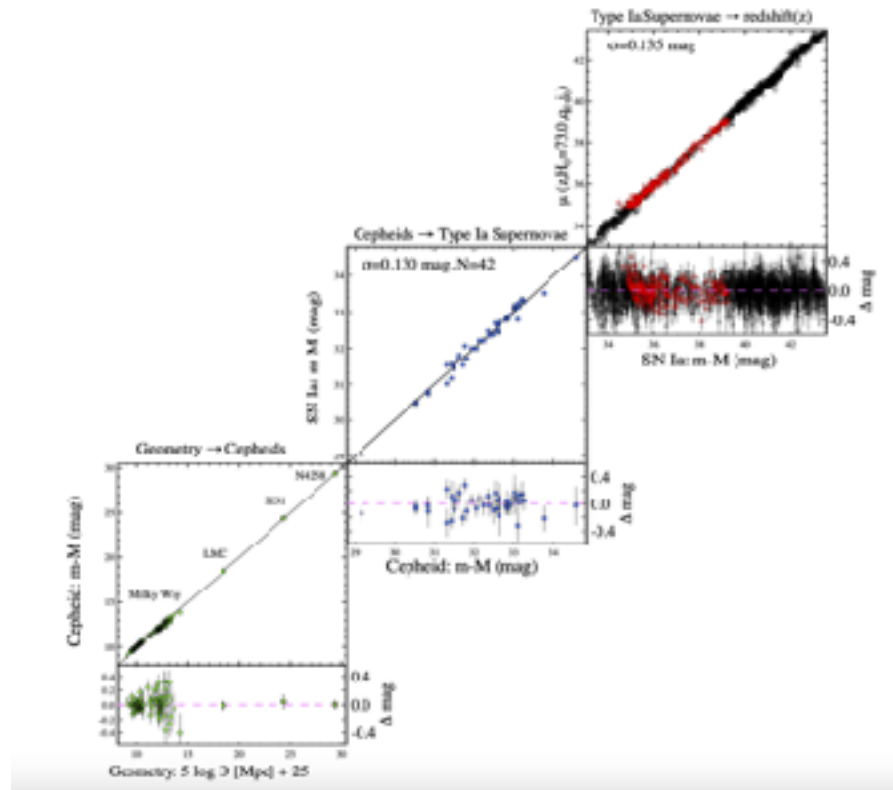


dark energy systematics -> lensing helps study high-z SN physics



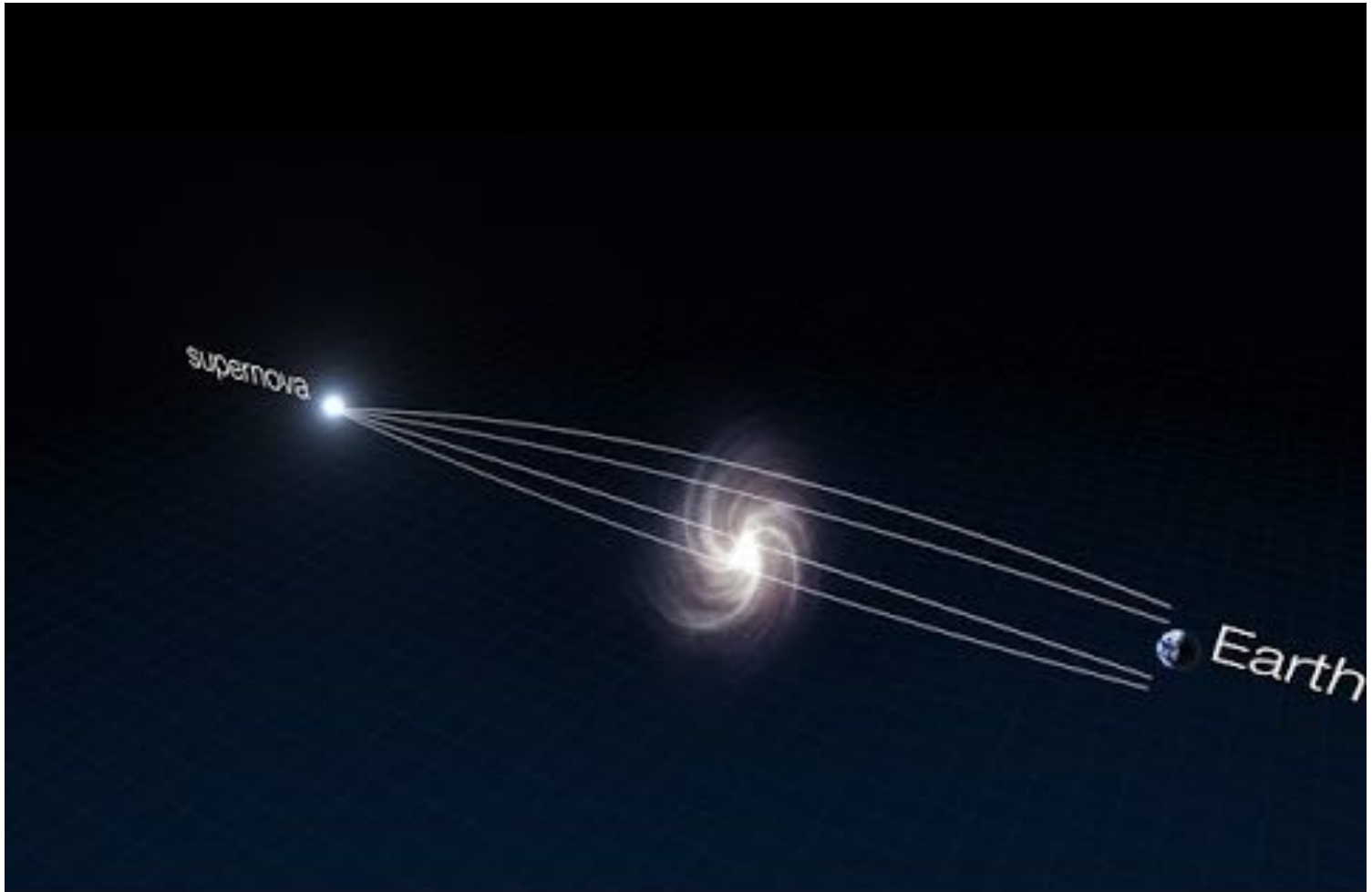
# Cosmic Distance Ladder

- Current most accurate method for direct  $H_0$  inference
- Type Ia supernova luminosity calibrated to Cepheids / TRGB
- Cepheids/TRGB calibrated to local distances, e.g. parallax, DEBs





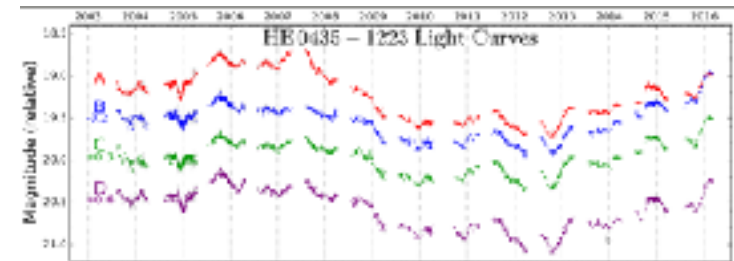
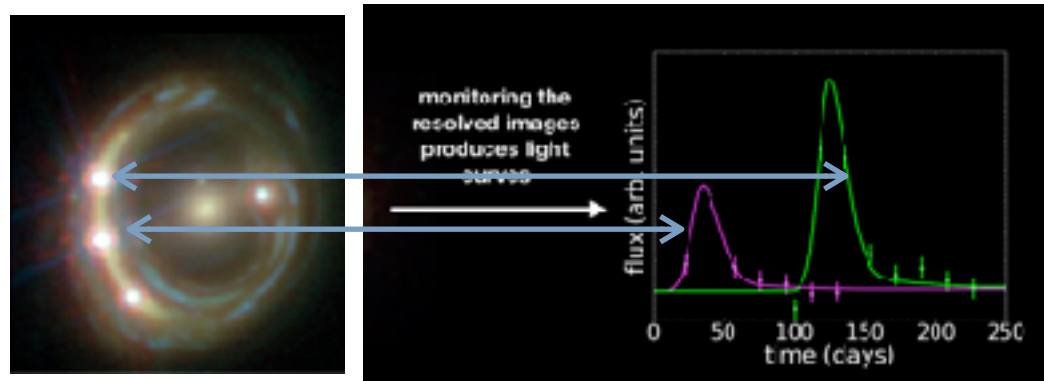
# Strongly lensed transients





# Time-delay cosmography

Typical lensed SN and QSO light curves



$$\Delta t \propto D_{\Delta t} \times \phi_{\text{lens}} \rightarrow D_{\Delta t} \propto \frac{1}{H_0}$$

Time delay    Time-delay distance    Lens potential (from mass model)

Completely Independent of unlensed SNe

\* NOT \* luminosity distance

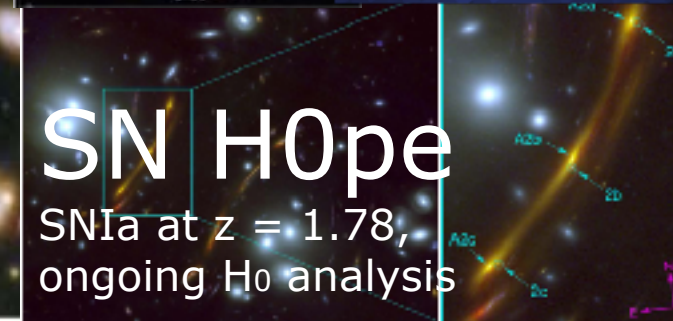
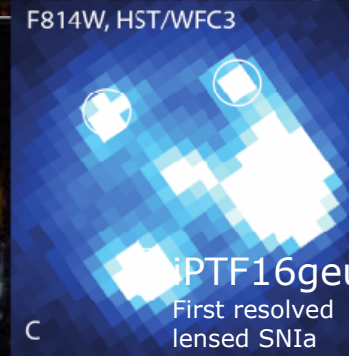
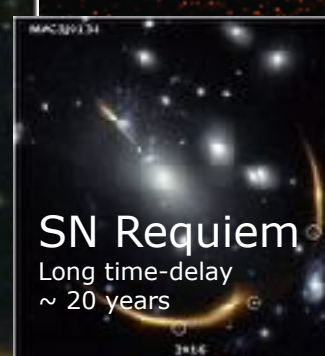
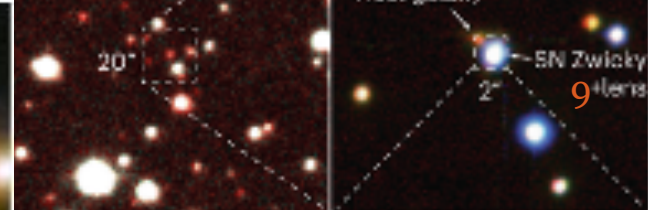
- Independent discovery method to lensed quasars
  - gISNe => “standardisable candle”
  - Lower impact of microlensing systematics
- First proposed in Refsdal 1964 (for SNe, used for QSOs)



# Recent Discoveries

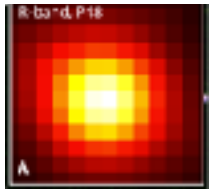


**SN Refsdal**  
IIP-87A-like;  $H_0$  to  $\sim 7\%$





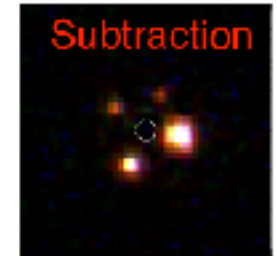
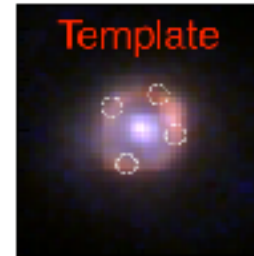
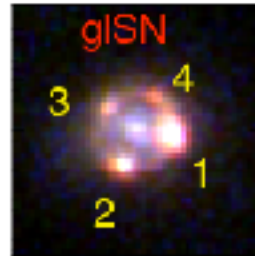
# First Resolved lensed SN Ia



Discovery in unresolved data

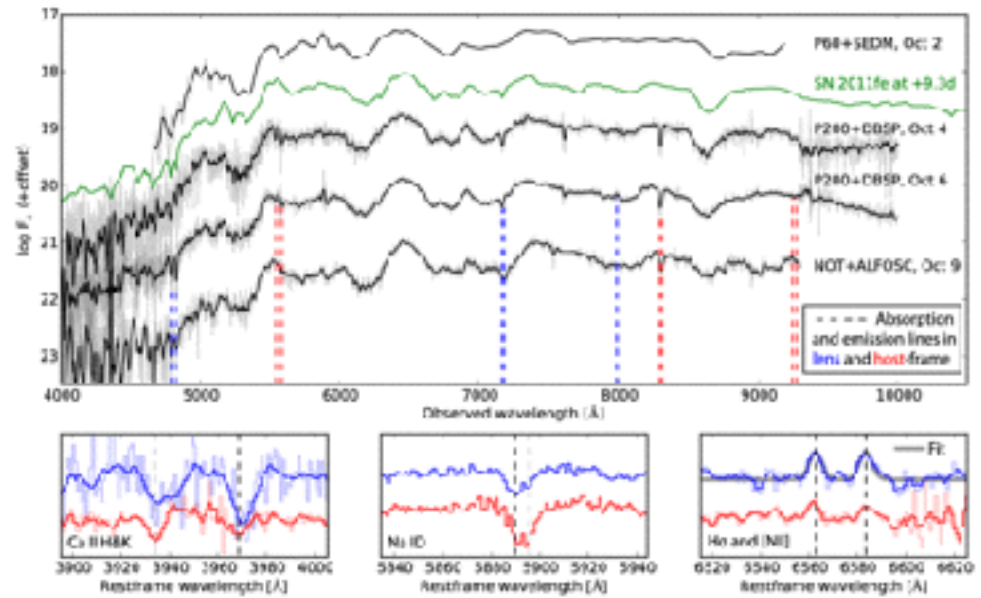
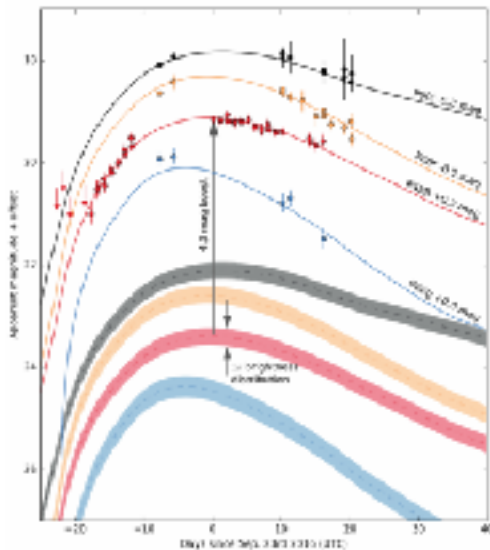


Follow-up: HST / AO



HST/WFC resolved image, template and subtraction => not possible for QSOs!

>50 times brighter than normal SNIa  
at  $z \sim 0.4$ : a  $30\sigma$  outlier!  
(Goobar+2017)



Higher resolution spectra -> second redshift from lens (Goobar+2017)



# Time-Delay Estimation

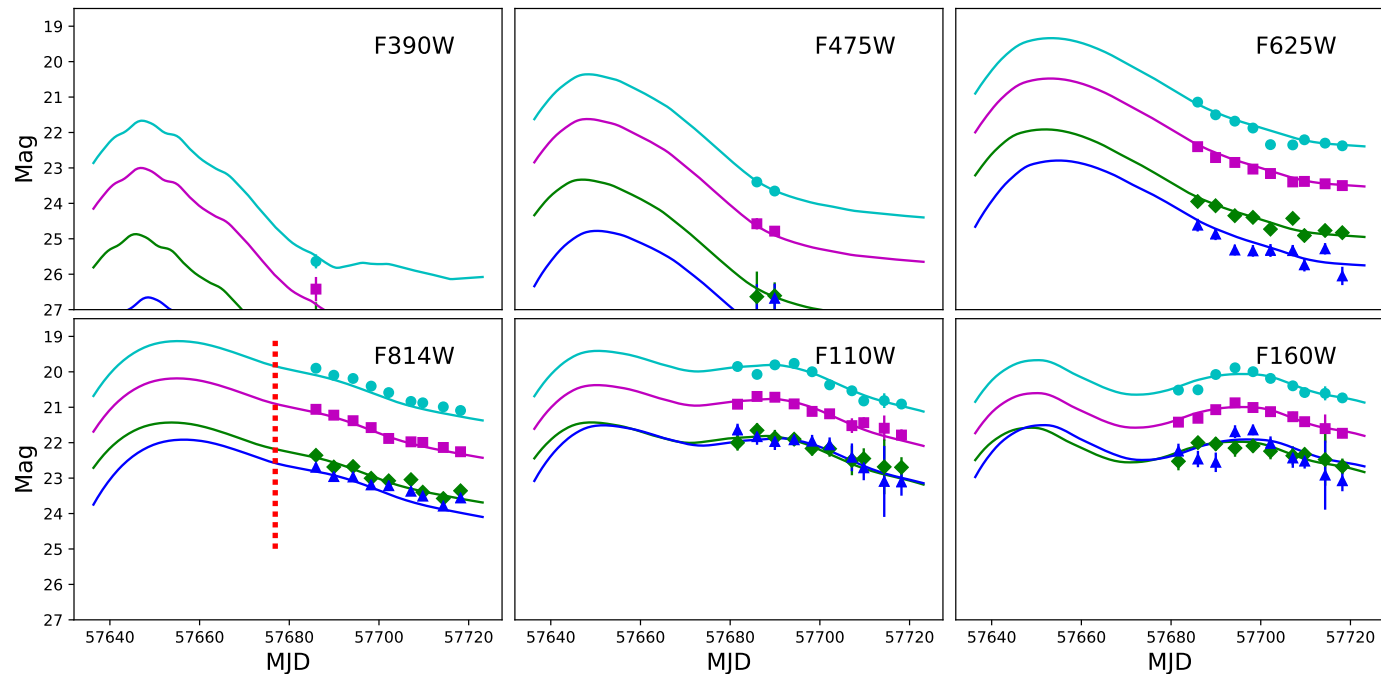
Very small time-delays (~ 1 day):  
Not ideal for measuring  $H_0$

Max. light simulations  
=> five times smaller error

Just a few orbits: errors ~ 1d

Long wavelength lever arm for extinction constraints

**SD+20b**

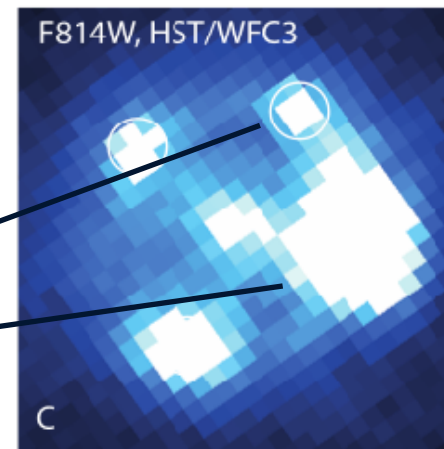




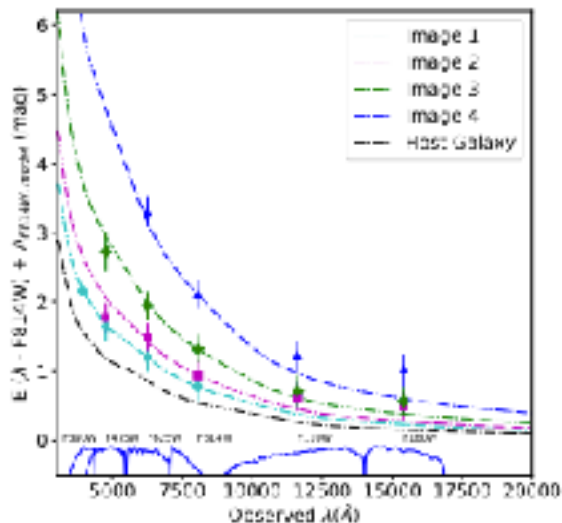
# iPTF16geu: Magnification + extinction

Important probe of dust in lens galaxy  
LoS  $R_v <$  Milky Way values

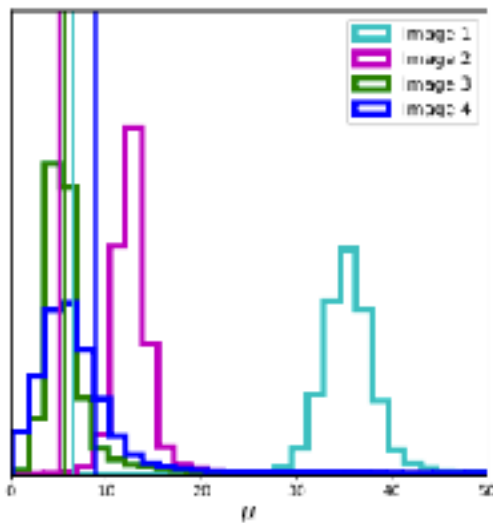
Model independent lensing magnification



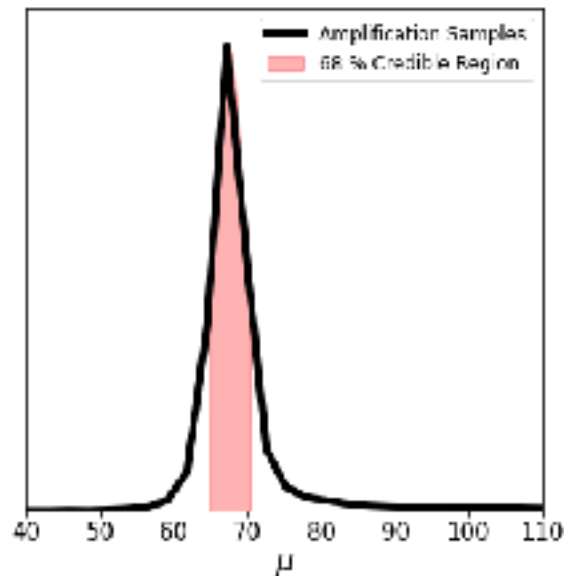
Preliminary magnification ( $\mu$ )  $\sim$  52  
With extinction correction 67 $\pm$ 3



Spectroscopy in Johansson, ..., SD, + '21

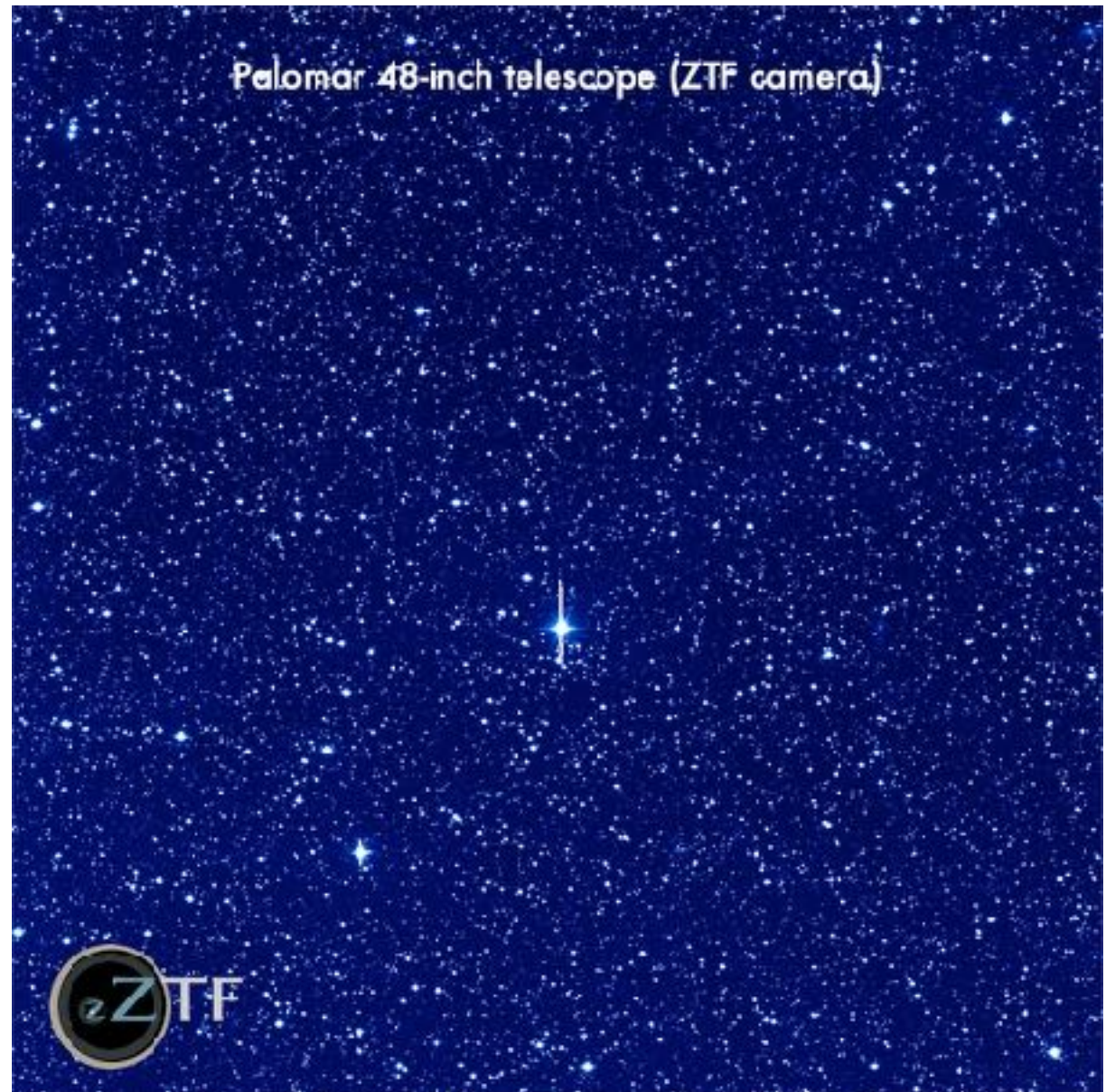


Modelling details in Mortzell, ..., SD, + '21





# First lensed SN with ZTF

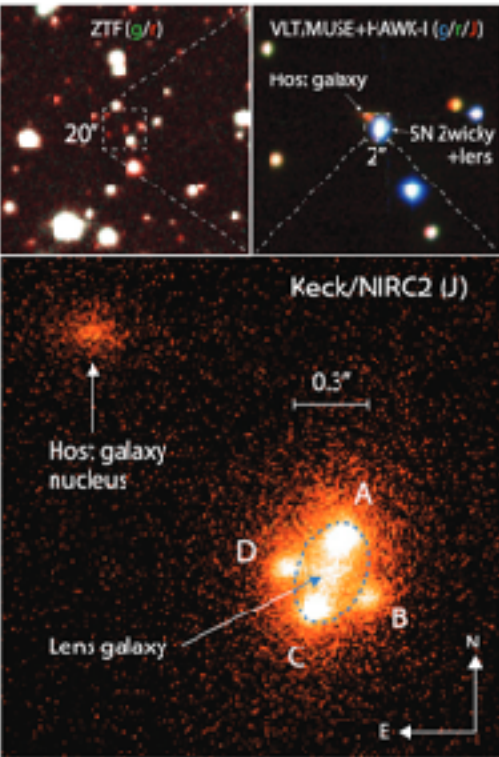


Credits: J. Johansson,  
N. Arendse



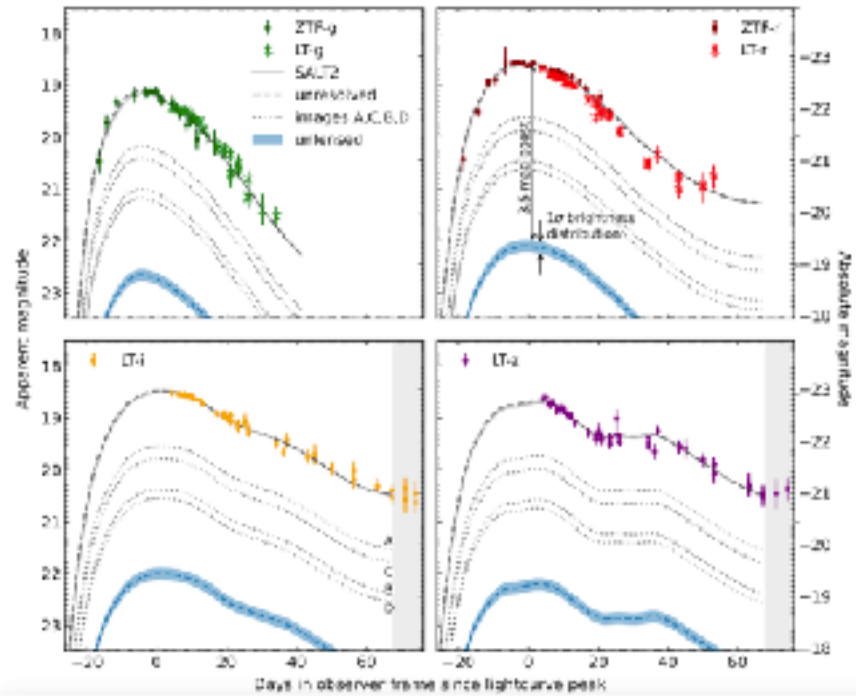
1"

Multiband P48+ LT data  
Accurate extinction constraints  
PI: Dhawan, Perley



~ 0.3"

0.01"



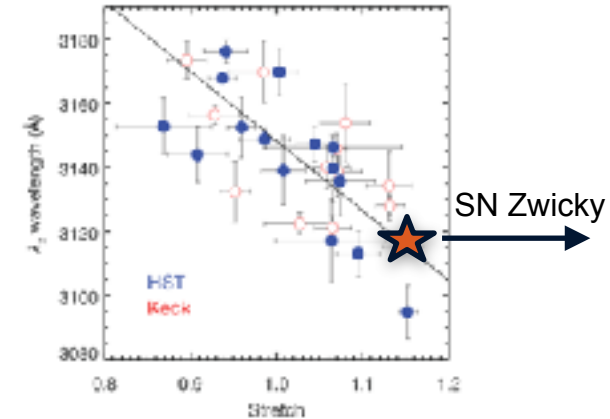
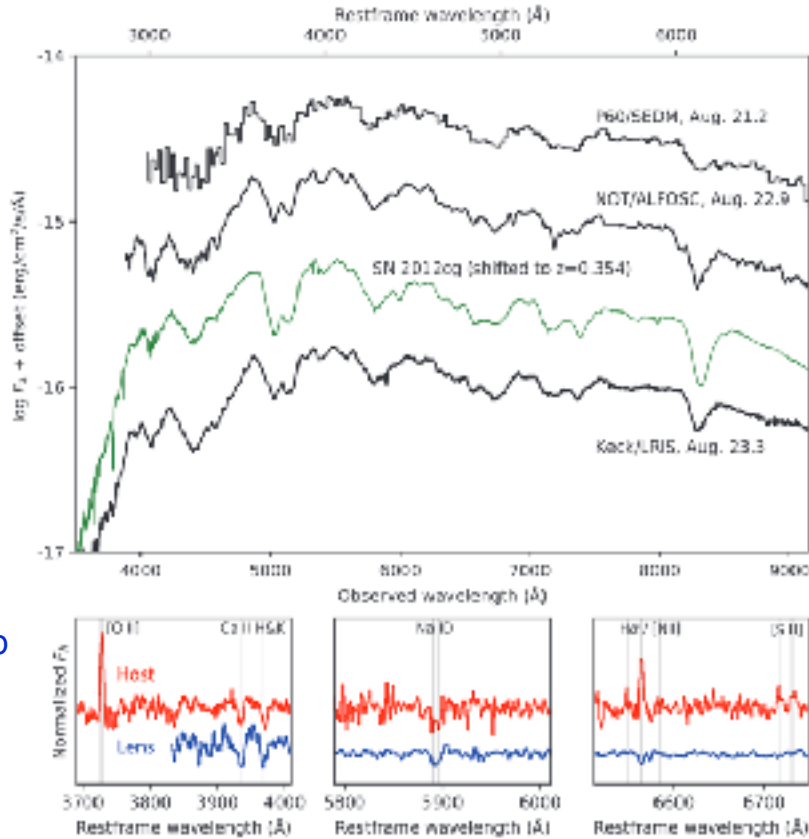
Small telescopes vital for bright lensed SNe

~ 3.5 mag > SN Ia at  $z_s = 0.354$

Low extinction in host + lens

Compact system  $\theta_E < 0.2''$ : study central stellar IMF

# Spectroscopy of lensed SNe



Comparison to low-z SNe sample from Maguire +2012

SN Zwicky spectra compared to local SN 2012cg

Johansson, SD+, in prep

No signs of cosmic spectroscopic evolution!

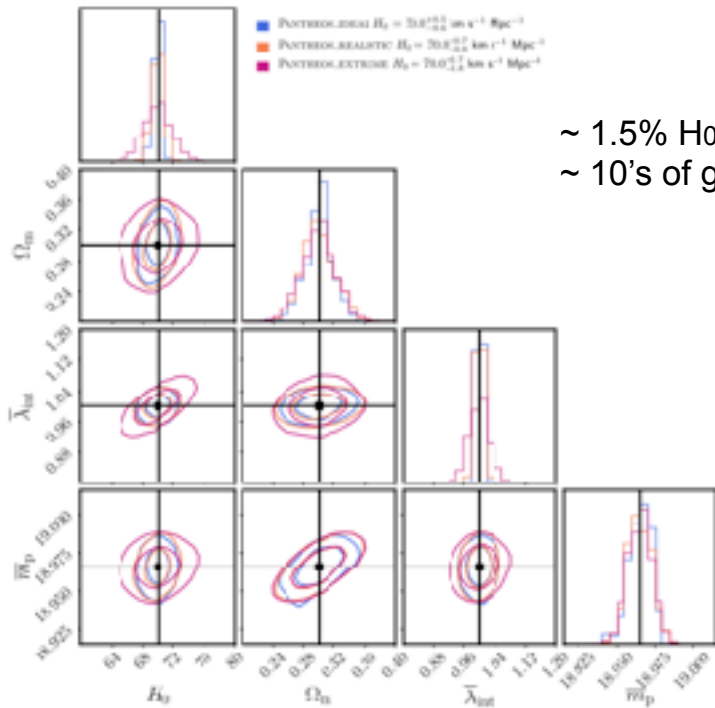
JWST NIRCcam + NIRSpec  
Nebular observations of SN Zwicky



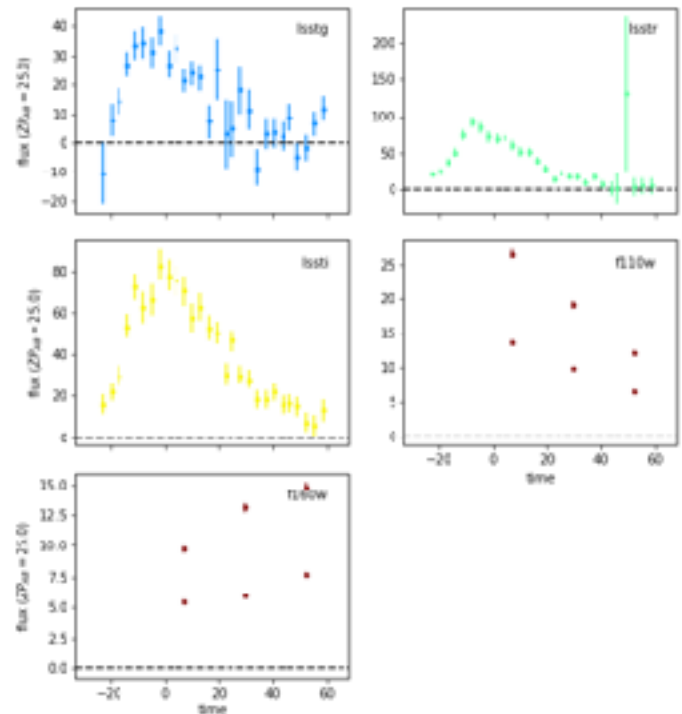
# What's next!

## Vera C. Rubin Observatory





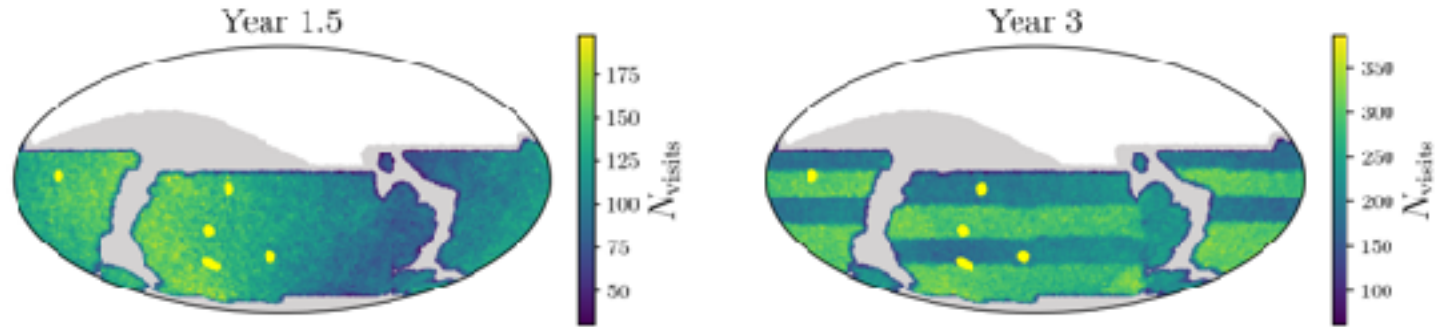
~ 1.5%  $H_0$  with LSST  
 ~ 10's of gLSNe



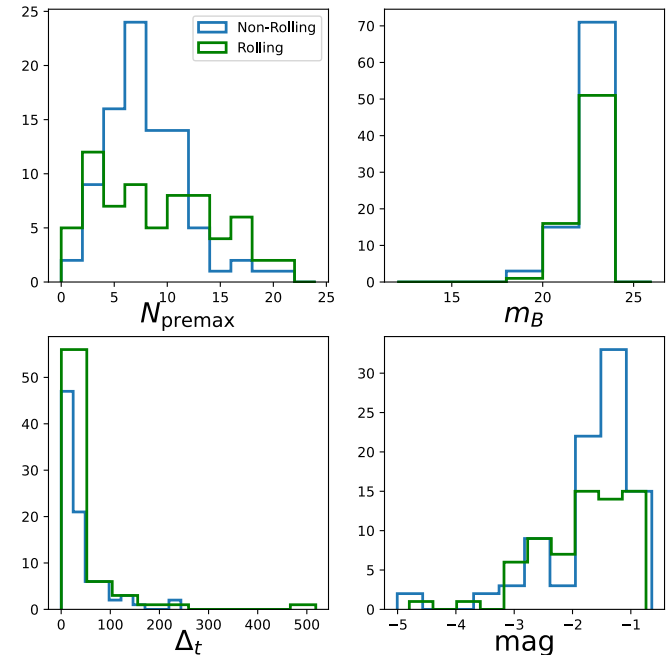
Birrer, SD, Shajib, 2022

Resolved photometry breaks degeneracies  
 Three epochs sufficient to 5%,  
 > 3 filters needed for < 3% time-delays

# Detectability in LSST

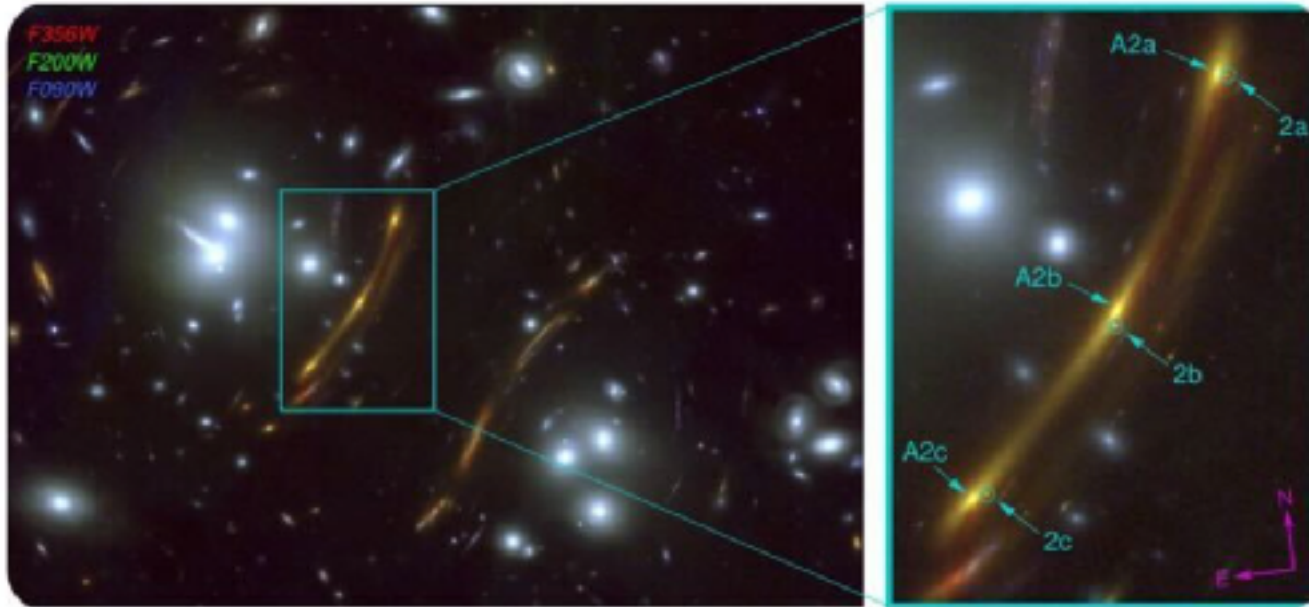


Baseline v3.0 cadence  
 Many 10's expected per year:  $\sim 20$  with large  $\Delta t$   
 Rolling has fewer  $\rightarrow$  denser sampling



Several with early sampling for discovery  
 Bright for 4m spec classification  
 Long time delays for cosmology





Triply imaged SN Ia at  $z = 1.78$

Cluster lens, long expected time-delay

Multiband follow-up with NIRCcam (F090W  $\rightarrow$  F444W)

Expected  $H_0$  at 10%  $\rightarrow$  7% with reference epochs

Important to test spectroscopic time-delays



# Spectroscopic Time-Delay <sup>20</sup>

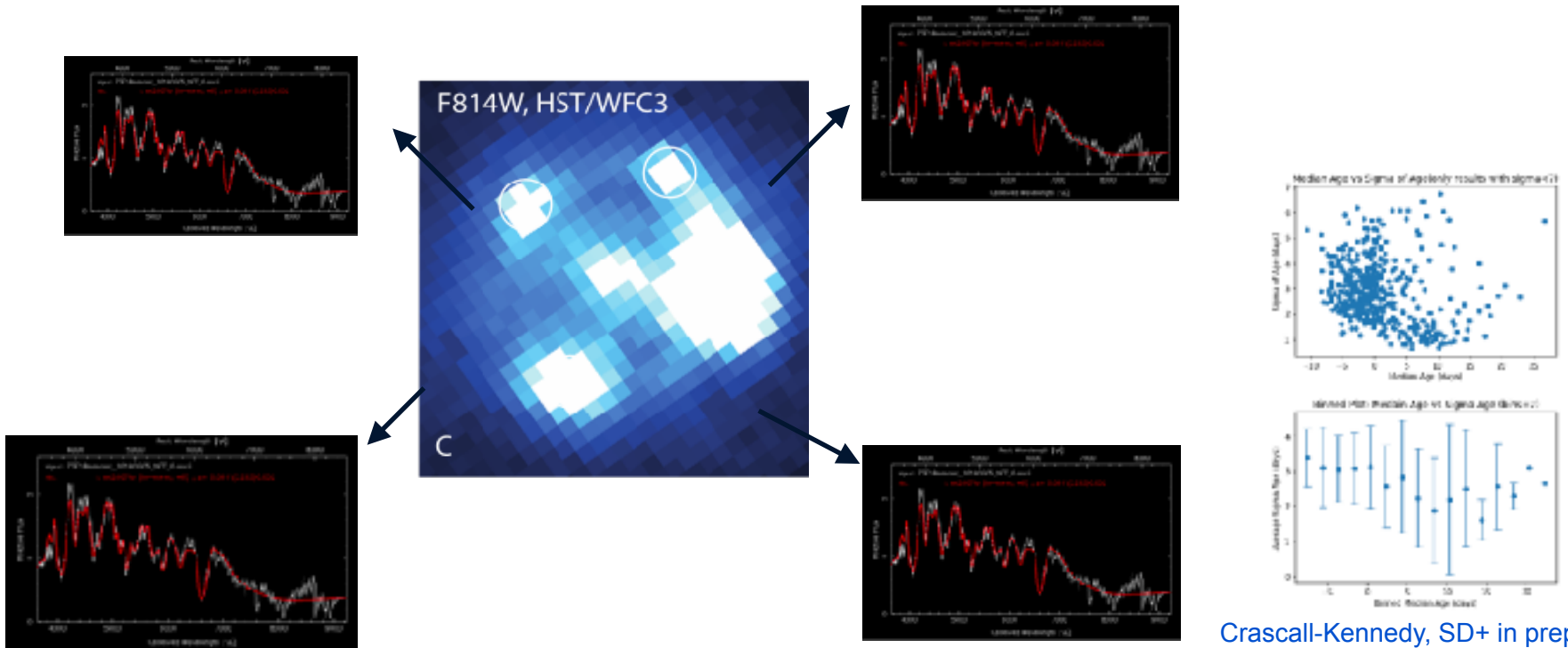
Spectroscopic time-delay: “one-shot” method (Johansson+2021 for 16geu)

Cross-correlate against template spectra

Small errors (< 2d) near maximum light

Works similarly well for core-collapse SNe

Example SNID fit for a typical low-z SNIa spectrum





# What's ongoing: JWST!

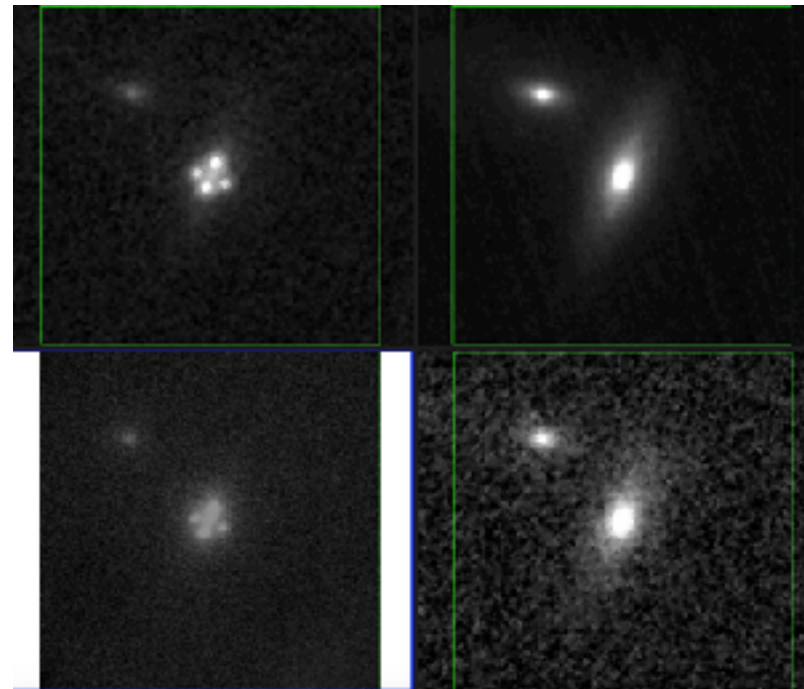
Nebular phase (~ 1 year post explosion) observations

Four filter photometry + spectra (NIRCam + NIRSpec IFU)

Important for post-explosion host + lens modelling

Nebular spectra -> shed light on progenitors

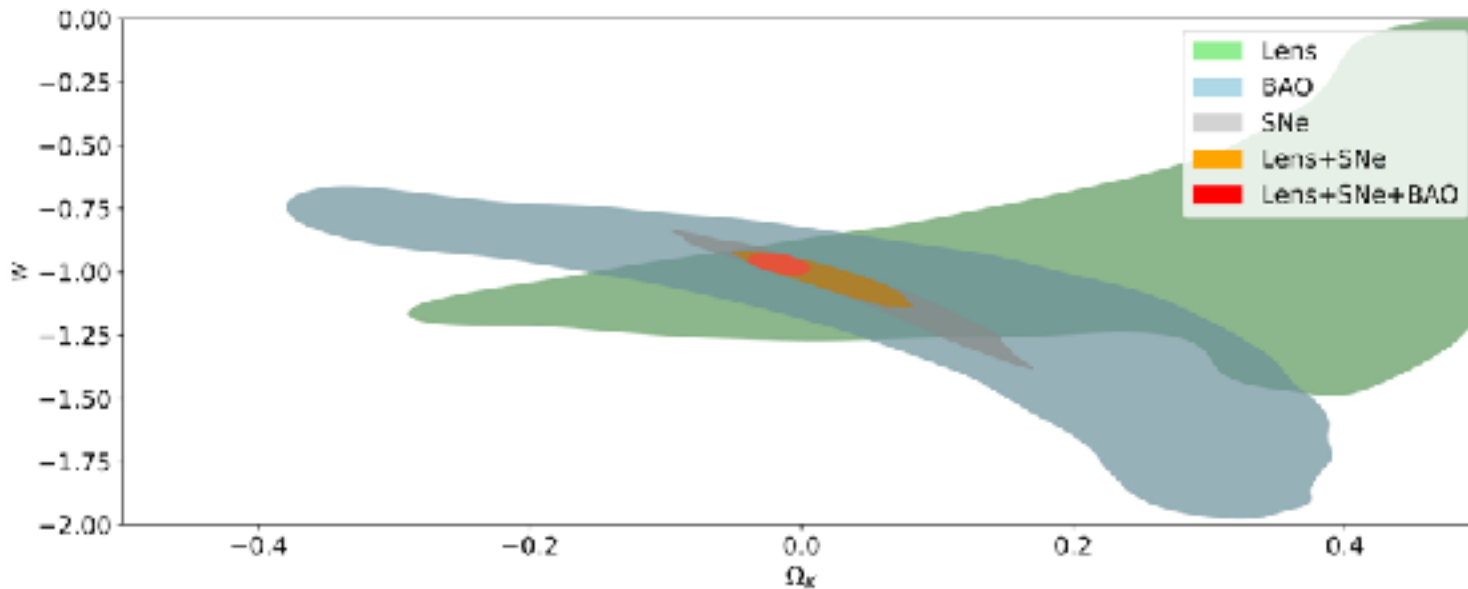
NIRCam image  
of host + lens





# What's ongoing: Spatial Curvature Constraints

- Simultaneous dark energy and curvature constraints
- Weak constraints from lensing  $\rightarrow$  orthogonal to SNe+BAO
  - Breaks degeneracy : Improves  $\Omega_k$  by factor 2





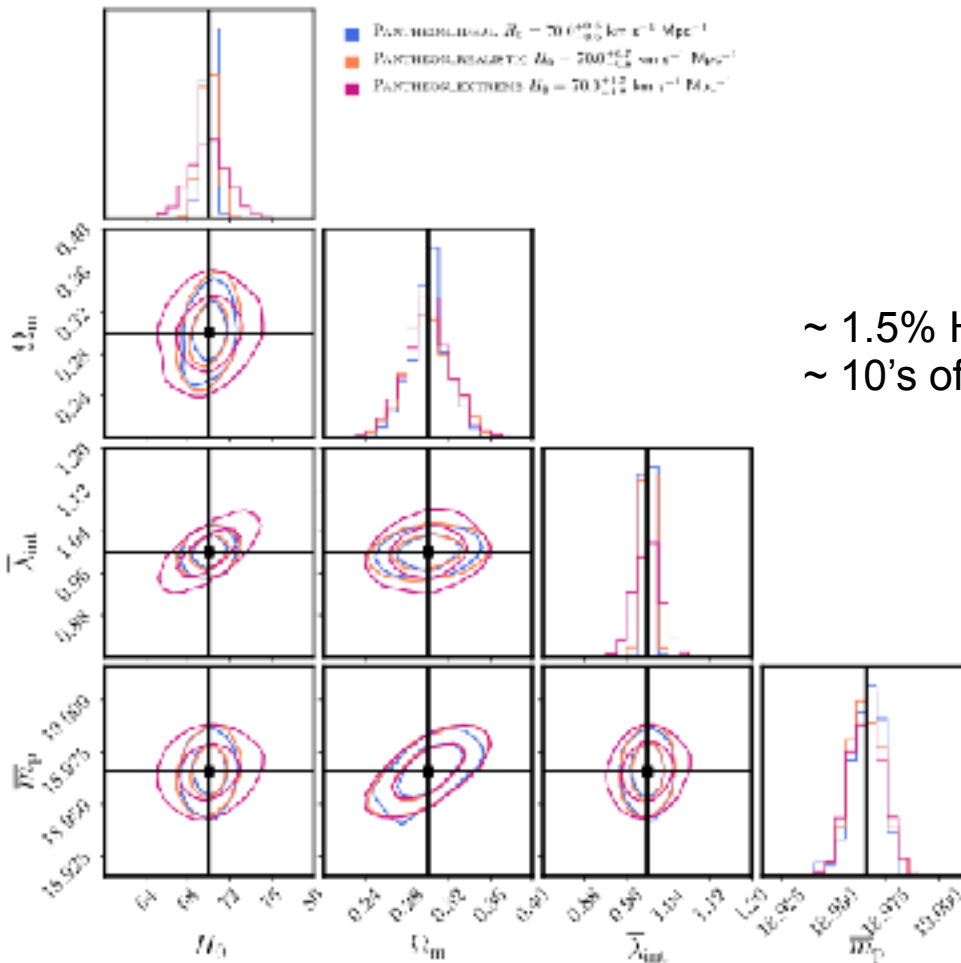
# Conclusions

- Current survey discoveries
  - 16geu: Deviation of extinction from MW
  - No sign of cosmic spectroscopic evolution
  - Compact lens of SN Zwicky -> exotic DM?
- Forecasts for cosmology
  - Independent  $H_0$  at 1.5% with LSST
  - Detect a large sample with feasible spectroscopy
- New inference methods
  - Spectroscopic time-delay at  $< 2\sigma$  error
- Excellent complement to SNeIa for curvature

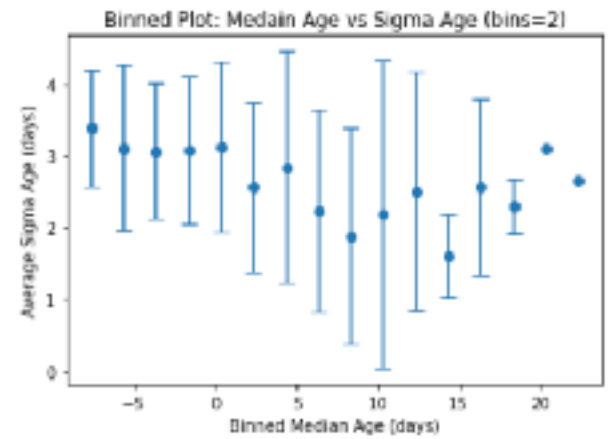
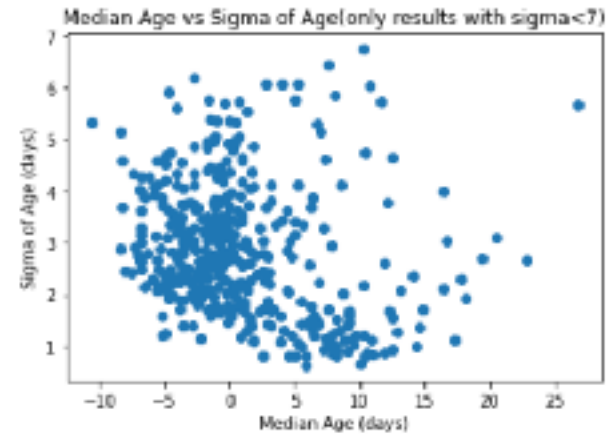
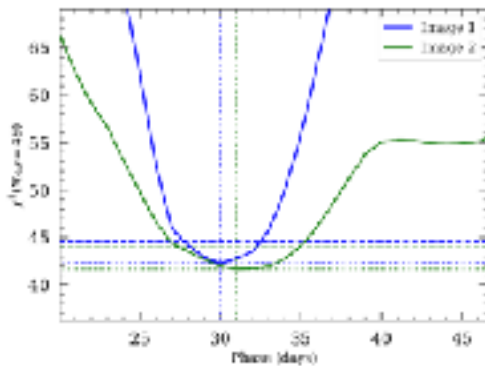
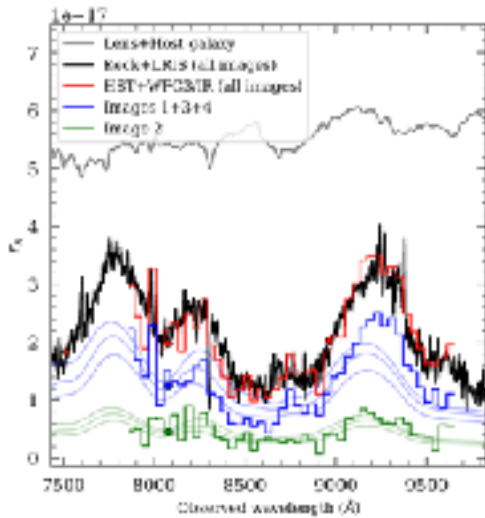




# Gearing up for LSST



~ 1.5%  $H_0$  with LSST  
 ~ 10's of gLSNe



Spectroscopic time-delay: “one-shot” method (Johansson+2021, JWST NIRSpec)

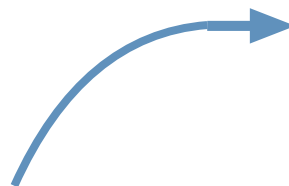
SNID: low errors (< 2d) with maximum light spectra  
Similarly low for CCSNe

(L. Crascall-Kennedy, SD in prep.)

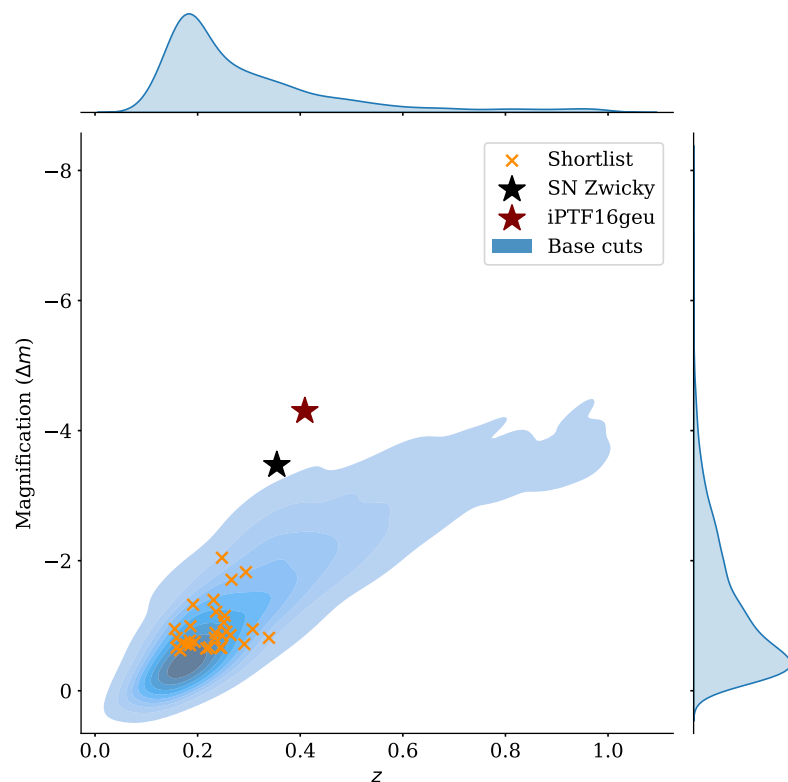


# ZTF archive search for lensed SNe

- Systematic search for strongly lensed SNe Ia in the ZTF archive.
- Applying cuts based on simulations.
  - Cuts on photo-z, distance to host, peak absolute magnitude, SN and host colours.
- 31 930 alerts  $\rightarrow$  30 candidates.
- Paper out this year!



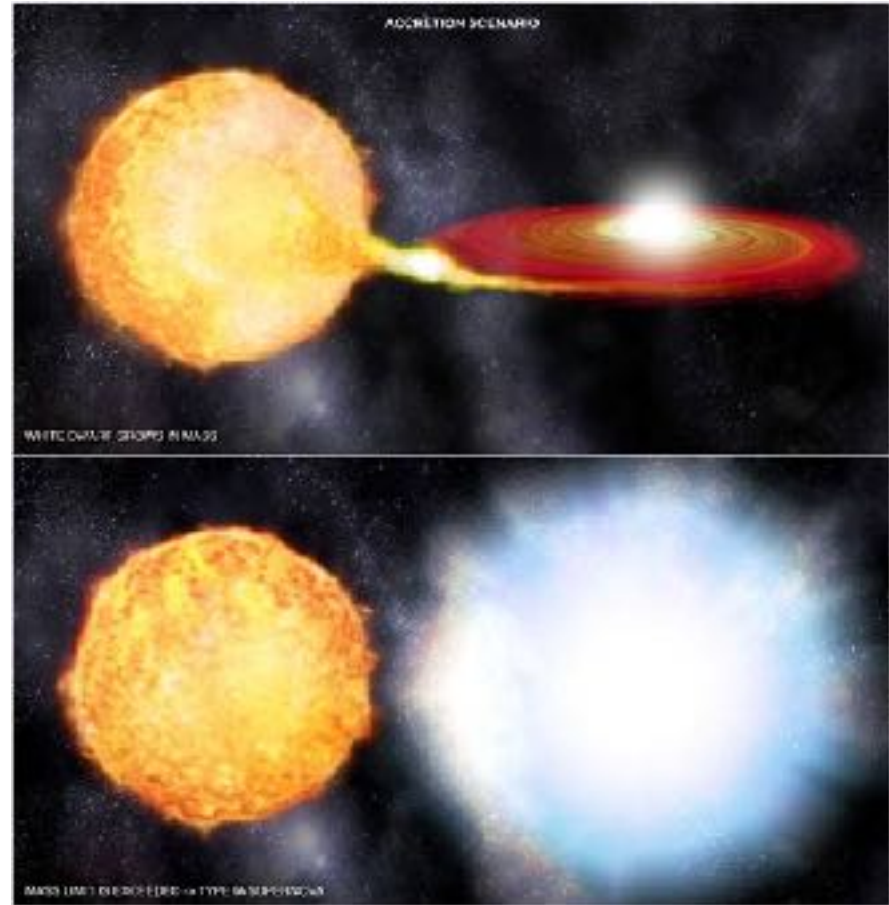
*Magnification vs.  $z$  for the best 7000 candidates (base cuts), and the shortlist candidates.*





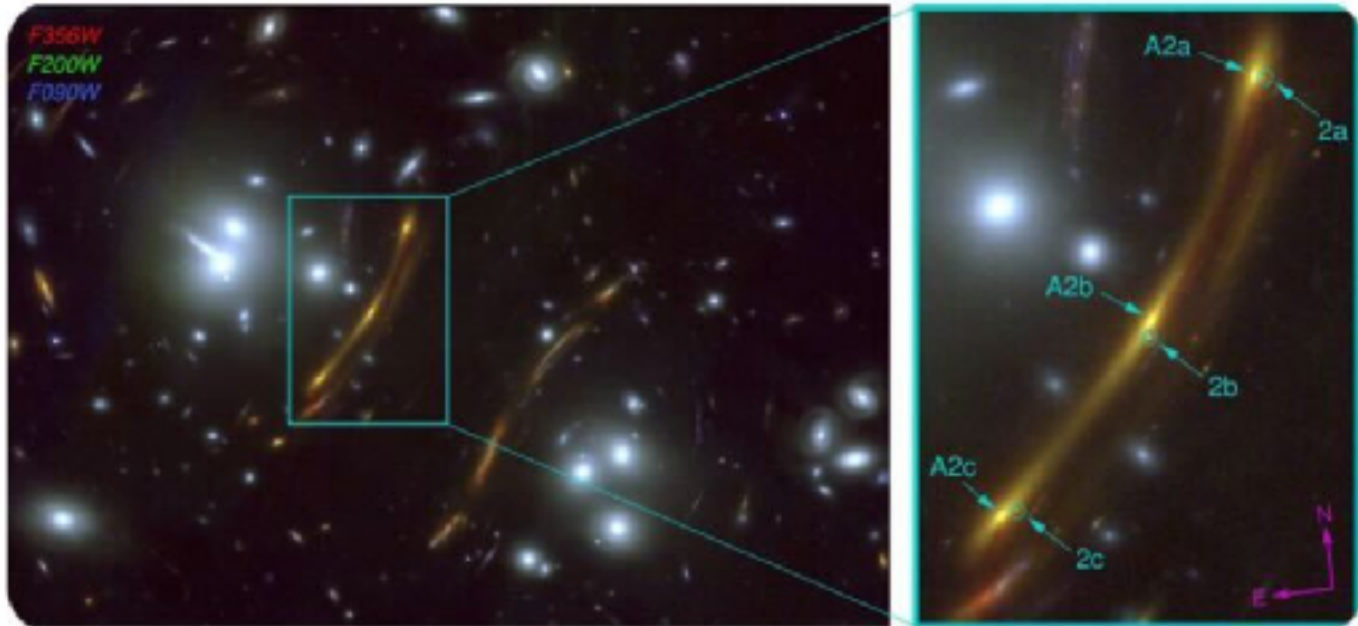
# What are Type Ia supernovae?<sup>27</sup>

NOT standard but calibratable -> small scatter, reduce lensing uncertainties



dark energy systematics -> lensing helps study high-z SN physics

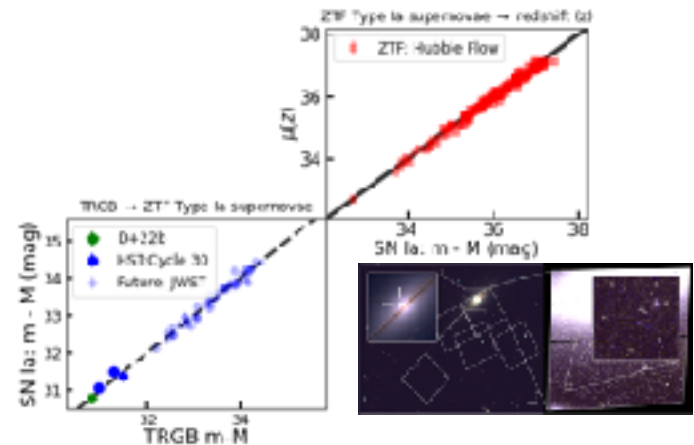
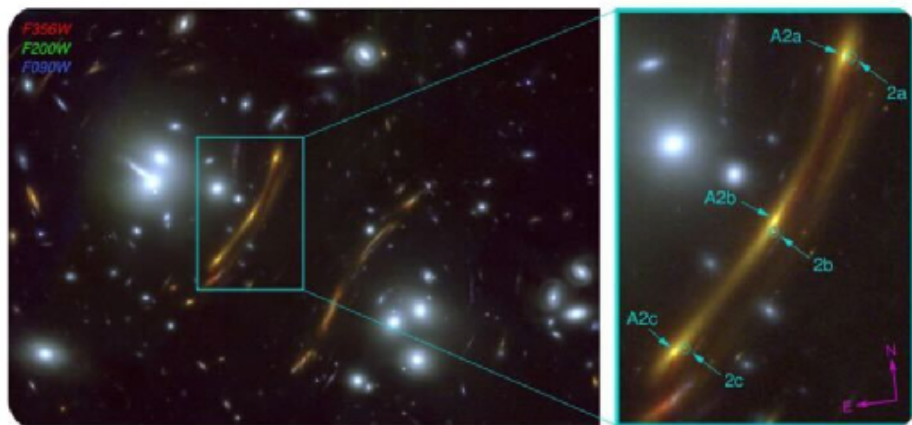
# SN Hope



Discovery in NIRCAM obs of PLCK G165.7+67 ; PEARLS program

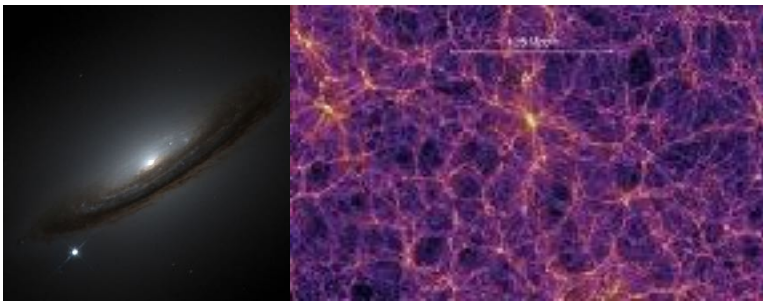
Cluster lens, long expected time-delay

Triply imaged SN Ia at  $z \sim 2$ ; Follow-up with DDT ongoing



## Strongly Lensed Supernovae

- Discovery with Vera C. Rubin Observatory
- Analyses with JWST IFU data
- Novel methods with current surveys



## Local Hubble Constant

- Uniform, large Type Ia supernova datasets
- New distance indicators

## Structure growth with SNe

- Bulk flows from local superstructures
- $\sigma_8$  with Type Ia supernovae